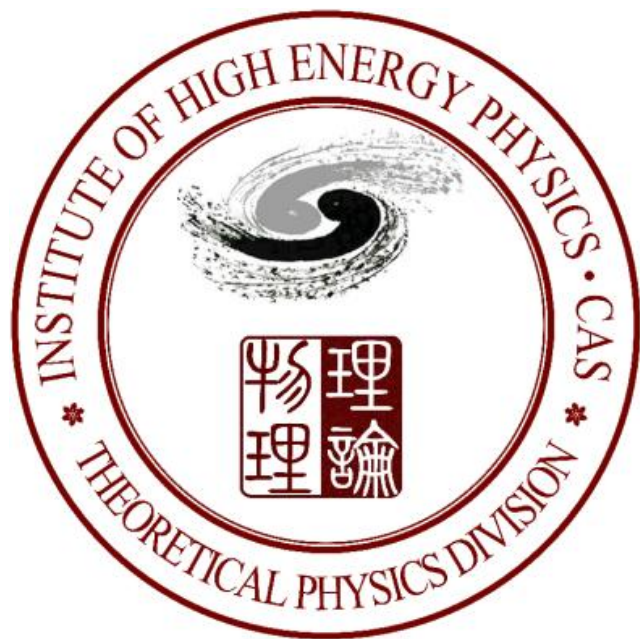


# Neutrinos——The Basics & Hot Topics

邢志忠

中科院高能所/国科大近物系

- ★ A brief history of neutrinos
- ★ Basic neutrino interactions
- ★ Dirac and Majorana masses
- ★ Flavor mixing & CP violation
- ★ Oscillation phenomenology
- ★ Neutrinoless double- $\beta$  decay
- ★ Typical seesaw mechanisms
- ★ Two types of cosmic neutrinos
- ★ Matter-antimatter asymmetry



# Leptons: a partial list

1

**1897:** Discovery of **electron** (J.J. Thomson)



**1928:** Prediction of positron (P.A.M. Dirac)



**1930:** Postulation of neutrino (W. Pauli)

**1932:** Discovery of positron (C.D. Anderson)



**1933:** Effective theory of beta decay (E. Fermi)

**1936:** Discovery of **muon** (J.C. Street et al; C.D. Anderson et al)

**1956:** Discovery of **electron antineutrino** (C.L. Cowan et al)



**1956:** Postulation of parity violation (T.D. Lee, C.N. Yang)



**1957:** Discovery of parity violation (C.S. Wu et al)

**1962:** Discovery of **muon neutrino** (G. Danby et al)



**1962:** Postulation of neutrino flavor conversion (Z. Maki et al)

**1967:** Standard model of leptons (S. Weinberg)



**1975:** Discovery of **tau** (M. Perl et al)



**2000:** Discovery of **tau neutrino** (K. Kodama et al)

# Quarks: a partial list

2

**1917:** Discovery of proton (E. Rutherford) **up and down**

**1932:** Discovery of neutron (J. Chadwick) **up and down**

**1947:** Discovery of Kaon (G. Rochester, C. Butler) **strange**

**1960:** The quark model (M. Gell-Mann; G. Zweig)

**1963:** The Cabibbo angle of quark mixing (N. Cabibbo)

**1964:** Discovery of CP violation (J.W. Cronin, V.L. Fitch)

**1964:** The Higgs mechanism (F. Englert, R. Brout; P. Higgs)

**1967:** The standard model (S. Weinberg)

**1970:** The GIM mechanism (S. Glashow et al)

**1973:** Asymptotic freedom (F. Wilczek, D. Gross; H. Politzer)

**1973:** The origin of CP violation (M. Kobayashi, T. Maskawa)

**1974:** Discovery of **charm** (C.C. Ting; B. Richter)

**1977:** Discovery of **bottom** (L. Lederman et al)

**1995:** Discovery of **top** (F. Abe et al)

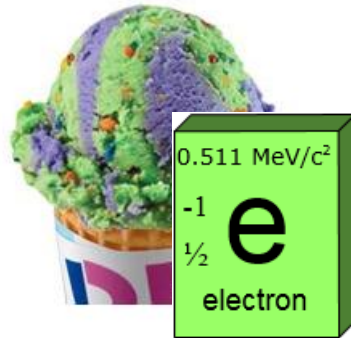


# Origin of "flavor"

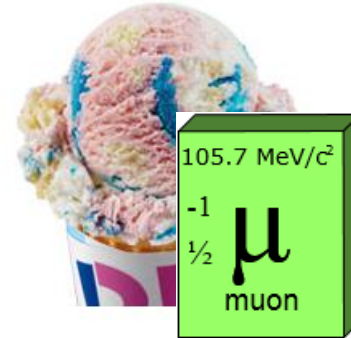
The term **Flavor** was coined by **Harald Fritzsch** and **Murray Gell-Mann** at a Baskin-Robbins ice-cream store in Pasadena in **1971**.



One of the most puzzling things in particle physics is **flavor mixing** !  
But this is normal for ice creams!



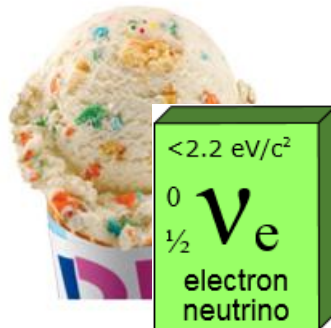
Rock 'n Pop Swirl Sherbet >



America's Birthday Cake® Ice Cream >



Cotton Candy Ice Cream >



Icing on the Cake® Ice Cream >



Wild 'n Reckless Sherbet >



Splish Splash® Sherbet >

# Lecture A4

- ★ Flavor mixing and CP violation
- ★ The PMNS flavor mixing matrix
- ★ What is behind:  $\mu$ - $\tau$  symmetry

# Flavor mixing

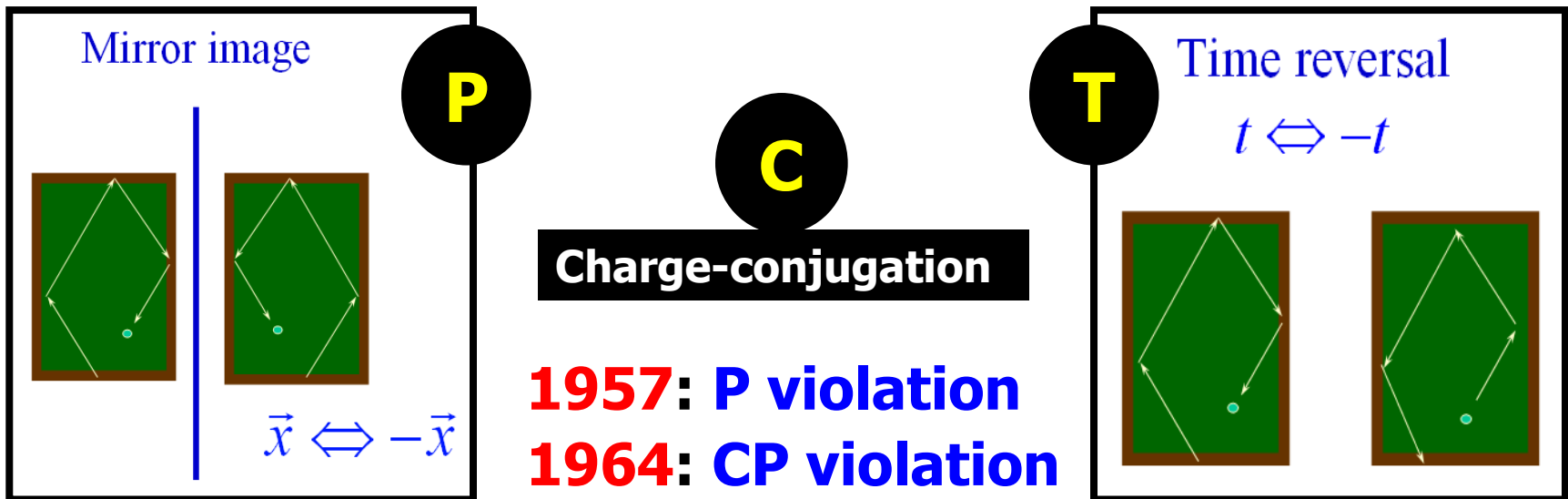
5

**Flavor mixing:** mismatch between **weak/flavor** eigenstates and **mass** eigenstates of fermions due to coexistence of **2** types of interactions.

**Weak eigenstates:** members of weak isospin doublets transforming into each other through the interaction with the **W** boson;

**Mass eigenstates:** states of definite masses that are created by the interaction with the Higgs boson (**Yukawa** interactions).

**CP violation:** **matter** and **antimatter**, or a reaction & its CP-conjugate process, are distinguishable --- coexistence of **2** types of interactions.



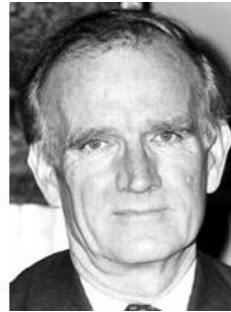


# Towards the KM paper

6

**1964:** Discovery of CP violation in K decays  
(J.W. Cronin, Val L. Fitch)

NP 1980



**1967:** Sakharov conditions for cosmological  
matter-antimatter asymmetry (A. Sakharov)

NP 1975



**1967:** The standard model of electromagnetic and  
weak interactions without quarks (S. Weinberg)

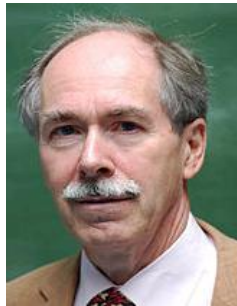
NP 1979



0 citation for the first 4 yrs

**1971:** The first proof of the renormalizability of the  
standard model (G. 't Hooft)

NP 1999



Progress of Theoretical Physics, Vol. 49, No. 2, February 1973

## *CP*-Violation in the Renormalizable Theory of Weak Interaction

Makoto KOBAYASHI and Toshihide MASKAWA

*Department of Physics, Kyoto University, Kyoto*

(Received September 1, 1972)



In a framework of the renormalizable theory of weak interaction, problems of *CP*-violation are studied. It is concluded that no realistic models of *CP*-violation exist in the quartet scheme without introducing any other new fields. Some possible models of *CP*-violation are also discussed.

**3 families allow for CP violation: Maskawa's bathtub idea!**

**"as I was getting out of the bathtub, an idea came to me"**



# Where or why

In the standard model, plus **3** right-handed  $\nu$ 's, where/why can flavor mixing and CP violation arise?

$$\mathcal{L}_{\nu\text{SM}} = \mathcal{L}_G + \mathcal{L}_H + \mathcal{L}_F + \mathcal{L}_Y$$

$$\mathcal{L}_G = -\frac{1}{4} (W^{i\mu\nu} W_{\mu\nu}^i + B^{\mu\nu} B_{\mu\nu})$$

$$\mathcal{L}_H = (D^\mu H)^\dagger (D_\mu H) - \mu^2 H^\dagger H - \lambda (H^\dagger H)^2$$

$$\mathcal{L}_F = \overline{Q}_L i \not{D} Q_L + \overline{\ell}_L i \not{D} \ell_L + \overline{U}_R i \not{D}' U_R + \overline{D}_R i \not{D}' D_R + \overline{E}_R i \not{D}' E_R + \overline{N}_R i \not{D}' N_R$$

$$\mathcal{L}_Y = -\overline{Q}_L Y_u \tilde{H} U_R - \overline{Q}_L Y_d H D_R - \overline{\ell}_L Y_l H E_R - \overline{\ell}_L Y_\nu \tilde{H} N_R + \text{h.c.}$$

**$\nu$ 's Dirac mass**

The strategy of diagnosis:

- ♣ **Flavor mixing:** transform the **flavor eigenstates** of fermions to their **mass eigenstates**, to see whether a kind of “**mismatch**” can occur.
- ♣ **CP violation:** given proper **CP** transformations of gauge, Higgs and fermion fields, one may prove that **1<sup>st</sup>**, **2<sup>nd</sup>** and **3<sup>rd</sup>** terms are formally invariant, and the **4<sup>th</sup>** term can be invariant only if the corresponding **Yukawa coupling matrices** are real. Otherwise, CP violation occurs.

# CP transformations

**Gauge fields:**

$$[B_\mu, W_\mu^1, W_\mu^2, W_\mu^3] \xrightarrow{\text{CP}} [-B^\mu, -W^{1\mu}, +W^{2\mu}, -W^{3\mu}]$$

$$[B_{\mu\nu}, W_{\mu\nu}^1, W_{\mu\nu}^2, W_{\mu\nu}^3] \xrightarrow{\text{CP}} [-B^{\mu\nu}, -W^{1\mu\nu}, +W^{2\mu\nu}, -W^{3\mu\nu}]$$

**Higgs fields:**

$$H(t, \mathbf{x}) = \begin{pmatrix} \phi^+ \\ \phi^0 \end{pmatrix} \xrightarrow{\text{CP}} H^*(t, -\mathbf{x}) = \begin{pmatrix} \phi^- \\ \phi^{0*} \end{pmatrix}$$

**Lepton or quark fields:**

$$\overline{\psi}_1 \gamma_\mu (1 \pm \gamma_5) \psi_2 \xrightarrow{\text{CP}} -\overline{\psi}_2 \gamma^\mu (1 \pm \gamma_5) \psi_1$$

$$\overline{\psi}_1 \gamma_\mu (1 \pm \gamma_5) \partial^\mu \psi_2 \xrightarrow{\text{CP}} \overline{\psi}_2 \gamma^\mu (1 \pm \gamma_5) \partial_\mu \psi_1$$

**Spinor bilinears:**

$\mathcal{L}_G$

$\mathcal{L}_H$

$\mathcal{L}_F$

*Formally invariant under CP*

	$\overline{\psi}_1 \psi_2$	$i\overline{\psi}_1 \gamma_5 \psi_2$	$\overline{\psi}_1 \gamma_\mu \psi_2$	$\overline{\psi}_1 \gamma_\mu \gamma_5 \psi_2$	$\overline{\psi}_1 \sigma_{\mu\nu} \psi_2$
C	$\overline{\psi}_2 \psi_1$	$i\overline{\psi}_2 \gamma_5 \psi_1$	$-\overline{\psi}_2 \gamma_\mu \psi_1$	$\overline{\psi}_2 \gamma_\mu \gamma_5 \psi_1$	$-\overline{\psi}_2 \sigma_{\mu\nu} \psi_1$
P	$\overline{\psi}_1 \psi_2$	$-i\overline{\psi}_1 \gamma_5 \psi_2$	$\overline{\psi}_1 \gamma^\mu \psi_2$	$-\overline{\psi}_1 \gamma^\mu \gamma_5 \psi_2$	$\overline{\psi}_1 \sigma^{\mu\nu} \psi_2$
T	$\overline{\psi}_1 \psi_2$	$-i\overline{\psi}_1 \gamma_5 \psi_2$	$\overline{\psi}_1 \gamma^\mu \psi_2$	$\overline{\psi}_1 \gamma^\mu \gamma_5 \psi_2$	$-\overline{\psi}_1 \sigma^{\mu\nu} \psi_2$
CP	$\overline{\psi}_2 \psi_1$	$-i\overline{\psi}_2 \gamma_5 \psi_1$	$-\overline{\psi}_2 \gamma^\mu \psi_1$	$-\overline{\psi}_2 \gamma^\mu \gamma_5 \psi_1$	$-\overline{\psi}_2 \sigma^{\mu\nu} \psi_1$
CPT	$\overline{\psi}_2 \psi_1$	$i\overline{\psi}_2 \gamma_5 \psi_1$	$-\overline{\psi}_2 \gamma_\mu \psi_1$	$-\overline{\psi}_2 \gamma_\mu \gamma_5 \psi_1$	$\overline{\psi}_2 \sigma_{\mu\nu} \psi_1$

# The source

10

The **Yukawa** interactions of fermions are **formally invariant** under **CP** if and only if

$$Y_u = Y_u^*, \quad Y_d = Y_d^* \\ Y_l = Y_l^*, \quad Y_\nu = Y_\nu^*$$

If the effective **Majorana** mass term is added into the SM, then the **Yukawa** interactions of leptons can be **formally invariant** under **CP** if

$$M_L = M_L^*, \quad Y_l = Y_l^*$$

If the **flavor eigenstates** are transformed into the **mass eigenstates**, flavor mixing and **CP** violation will show up in the **CC** interactions:

**quarks**

$$\mathcal{L}_{cc} = \frac{g}{\sqrt{2}} \overline{(u \ c \ t)}_L \gamma^\mu V \begin{pmatrix} d \\ s \\ b \end{pmatrix}_L W_\mu^+ + \text{h.c.}$$

**leptons**

$$\mathcal{L}_{cc} = \frac{g}{\sqrt{2}} \overline{(e \ \mu \ \tau)}_L \gamma^\mu U \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}_L W_\mu^- + \text{h.c.}$$

**Comment A:** flavor mixing and **CP** violation take place since fermions interact with both the **gauge bosons** and the **Higgs boson**.

**Comment B:** both the **CC** and **Yukawa** interactions have been verified.

**Comment C:** the **CKM** matrix **V** is unitary, the **PMNS** matrix **U** is too?

# Parameter counting

The **3×3** unitary matrix **V** can always be parametrized as a product of **3** unitary rotation matrices in the complex planes:

$$\begin{aligned} O_1(\theta_1, \alpha_1, \beta_1, \gamma_1) &= \begin{pmatrix} c_1 e^{i\alpha_1} & s_1 e^{-i\beta_1} & 0 \\ -s_1 e^{i\beta_1} & c_1 e^{-i\alpha_1} & 0 \\ 0 & 0 & e^{i\gamma_1} \end{pmatrix} \\ O_2(\theta_2, \alpha_2, \beta_2, \gamma_2) &= \begin{pmatrix} e^{i\gamma_2} & 0 & 0 \\ 0 & c_2 e^{i\alpha_2} & s_2 e^{-i\beta_2} \\ 0 & -s_2 e^{i\beta_2} & c_2 e^{-i\alpha_2} \end{pmatrix} \\ O_3(\theta_3, \alpha_3, \beta_3, \gamma_3) &= \begin{pmatrix} c_3 e^{i\alpha_3} & 0 & s_3 e^{-i\beta_3} \\ 0 & e^{i\gamma_3} & 0 \\ -s_3 e^{i\beta_3} & 0 & c_3 e^{-i\alpha_3} \end{pmatrix} \end{aligned}$$

where  $s_i \equiv \sin \theta_i$  and  $c_i \equiv \cos \theta_i$  (for  $i = 1, 2, 3$ )

**Category A: 3 possibilities**

$$V = O_i O_j O_i \quad (i \neq j)$$

**Category B: 6 possibilities**

$$V = O_i O_j O_k \quad (i \neq j \neq k)$$

For instance, the standard parametrization is given below:

**V**

$$\begin{aligned}
 &= \begin{pmatrix} e^{i\gamma_2} & 0 & 0 \\ 0 & c_2 e^{i\alpha_2} & s_2 e^{-i\beta_2} \\ 0 & -s_2 e^{i\beta_2} & c_2 e^{-i\alpha_2} \end{pmatrix} \begin{pmatrix} c_3 e^{i\alpha_3} & 0 & s_3 e^{-i\beta_3} \\ 0 & e^{i\gamma_3} & 0 \\ -s_3 e^{i\beta_3} & 0 & c_3 e^{-i\alpha_3} \end{pmatrix} \begin{pmatrix} c_1 e^{i\alpha_1} & s_1 e^{-i\beta_1} & 0 \\ -s_1 e^{i\beta_1} & c_1 e^{-i\alpha_1} & 0 \\ 0 & 0 & e^{i\gamma_1} \end{pmatrix} \\
 &= \begin{pmatrix} c_1 c_3 e^{i(\alpha_1 + \gamma_2 + \alpha_3)} & s_1 c_3 e^{i(-\beta_1 + \gamma_2 + \alpha_3)} & s_3 e^{i(\gamma_1 + \gamma_2 - \beta_3)} \\ -s_1 c_2 e^{i(\beta_1 + \alpha_2 + \gamma_3)} - c_1 s_2 s_3 e^{i(\alpha_1 - \beta_2 + \beta_3)} & c_1 c_2 e^{i(-\alpha_1 + \alpha_2 + \gamma_3)} - s_1 s_2 s_3 e^{i(-\beta_1 - \beta_2 + \beta_3)} & s_2 c_3 e^{i(\gamma_1 - \beta_2 - \alpha_3)} \\ s_1 s_2 e^{i(\beta_1 + \beta_2 + \gamma_3)} - c_1 c_2 s_3 e^{i(\alpha_1 - \alpha_2 + \beta_3)} & -c_1 s_2 e^{i(-\alpha_1 + \beta_2 + \gamma_3)} - s_1 c_2 s_3 e^{i(-\beta_1 - \alpha_2 + \beta_3)} & c_2 c_3 e^{i(\gamma_1 - \alpha_2 - \alpha_3)} \end{pmatrix} \\
 &= \begin{pmatrix} e^{ia} & 0 & 0 \\ 0 & e^{ib} & 0 \\ 0 & 0 & e^{ic} \end{pmatrix} \begin{pmatrix} c_1 c_3 & s_1 c_3 & s_3 e^{-i\delta} \\ -s_1 c_2 - c_1 s_2 s_3 e^{i\delta} & c_1 c_2 - s_1 s_2 s_3 e^{i\delta} & s_2 c_3 \\ s_1 s_2 - c_1 c_2 s_3 e^{i\delta} & -c_1 s_2 - s_1 c_2 s_3 e^{i\delta} & c_2 c_3 \end{pmatrix} \begin{pmatrix} e^{ix} & 0 & 0 \\ 0 & e^{iy} & 0 \\ 0 & 0 & e^{iz} \end{pmatrix}
 \end{aligned}$$

$$a = (\alpha_1 - \beta_1) - (\alpha_2 + \beta_2 - \gamma_2) - \gamma_3, \quad b = -\beta_2 - \alpha_3, \quad c = -\alpha_2 - \alpha_3;$$

$$x = \beta_1 + (\alpha_2 + \beta_2) + (\alpha_3 + \gamma_3), \quad y = -\alpha_1 + (\alpha_2 + \beta_2) + (\alpha_3 + \gamma_3), \quad z = \gamma_1.$$

$$\delta = \beta_3 - \gamma_1 - \gamma_2$$



If neutrinos are **Dirac** particles, the phases **x**, **y** and **z** can be removed. Then the neutrino mixing matrix is

## Dirac neutrino mixing matrix

$$V = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix}$$

If neutrinos are **Majorana** particles, left- and right-handed fields are correlated. Hence only a common phase of three left-handed fields can be redefined (e.g., **z = 0**). Then

## Majorana neutrino mixing matrix

$$V = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix} \begin{pmatrix} e^{i\rho} & 0 & 0 \\ 0 & e^{i\sigma} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

## F. Capozzi et al (1703.04471): a global fit of current $\nu$ -oscillation data

Parameter	Ordering	Best fit	$1\sigma$ range	$2\sigma$ range	$3\sigma$ range
$\delta m^2 / 10^{-5} \text{ eV}^2$	NO, IO, any	7.37	7.21–7.54	7.07–7.73	6.93–7.96
$\sin^2 \theta_{12} / 10^{-1}$	NO, IO, any	2.97	2.81–3.14	2.65–3.34	2.50–3.54
$ \Delta m^2  / 10^{-3} \text{ eV}^2$	NO	2.525	2.495–2.567	2.454–2.606	2.411–2.646
	IO	2.505	2.473–2.539	2.430–2.582	2.390–2.624
	Any	2.525	2.495–2.567	2.454–2.606	2.411–2.646
$\sin^2 \theta_{13} / 10^{-2}$	NO	2.15	2.08–2.22	1.99–2.31	1.90–2.40
	IO	2.16	2.07–2.24	1.98–2.33	1.90–2.42
	Any	2.15	2.08–2.22	1.99–2.31	1.90–2.40
$\sin^2 \theta_{23} / 10^{-1}$	NO	4.25	4.10–4.46	3.95–4.70	3.81–6.15
	IO	5.89	4.17–4.48 $\oplus$ 5.67–6.05	3.99–4.83 $\oplus$ 5.33–6.21	3.84–6.36
	Any	4.25	4.10–4.46	3.95–4.70 $\oplus$ 5.75–6.00	3.81–6.26
$\delta / \pi$	NO	1.38	1.18–1.61	1.00–1.90	0–0.17 $\oplus$ 0.76–2
	IO	1.31	1.12–1.62	0.92–1.88	0–0.15 $\oplus$ 0.69–2
	Any	1.38	1.18–1.61	1.00–1.90	0–0.17 $\oplus$ 0.76–2

“Summarizing, the **SK (+T2K)** official results and ours suggest, at face value, that global **3 $\nu$**  oscillation analysis may have reached an overall  $\sim 2\sigma$  sensitivity to the mass ordering, with a preference for **NO** driven by atmospheric data (= **multi-GeV e-like events**) and corroborated by accelerator data, together with reactor constraints.”

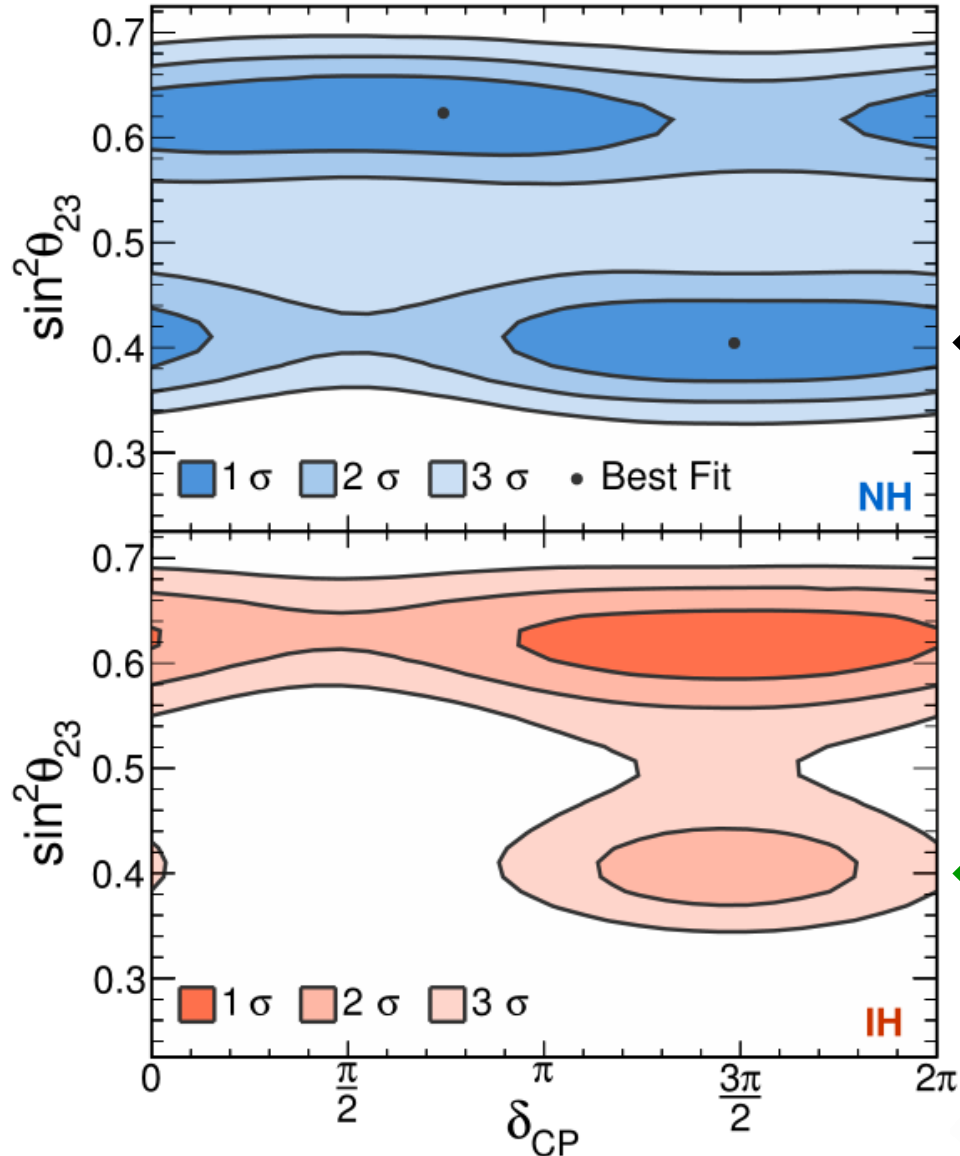
Terrestrial **matter effects** play the crucial role.

**T2K, NO $\nu$ A, SK, PINGU, INO, ...**

$$\Delta m_{31}^2 \mp 2\sqrt{2}G_F N_e E$$

# Normal ordering?

**NO<sub>v</sub>A (arXiv:1703.03328, 9 March 2017):**



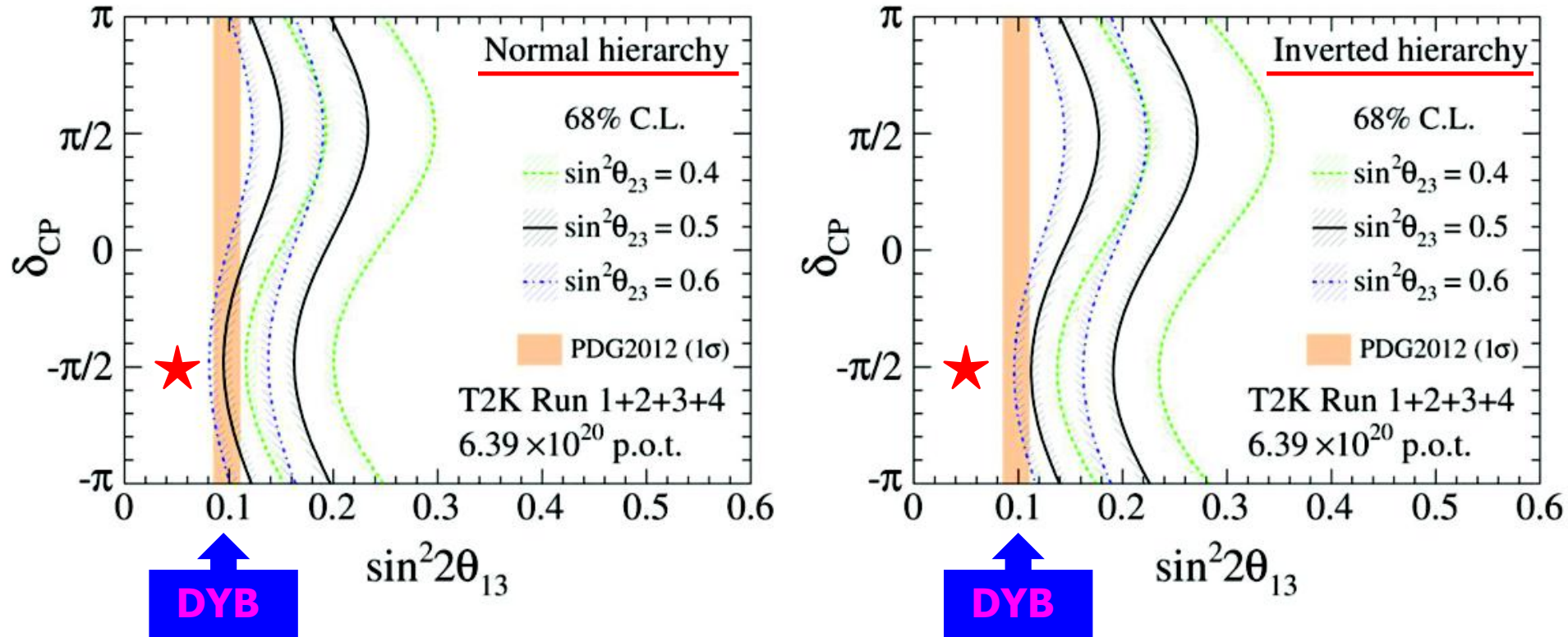
**my bet**

disfavored at greater than 93% C.L. for all values of  $\delta_{CP}$

# Hint for the CP phase

16

The **T2K** observation of a relatively strong  $\nu_\mu \rightarrow \nu_e$  appearance plays a crucial role in the global fit to make  $\theta_{13}$  consistent with the **Daya Bay** result and drive a slight but intriguing preference for  $\delta \sim -\pi/2$ .



**DYB's good news:**  $\theta_{13}$  unsuppressed

**T2K's good news:**  $\delta$  unsuppressed

Life is easier for probing CP violation,  $\nu$  mass hierarchy ....

precision measurements

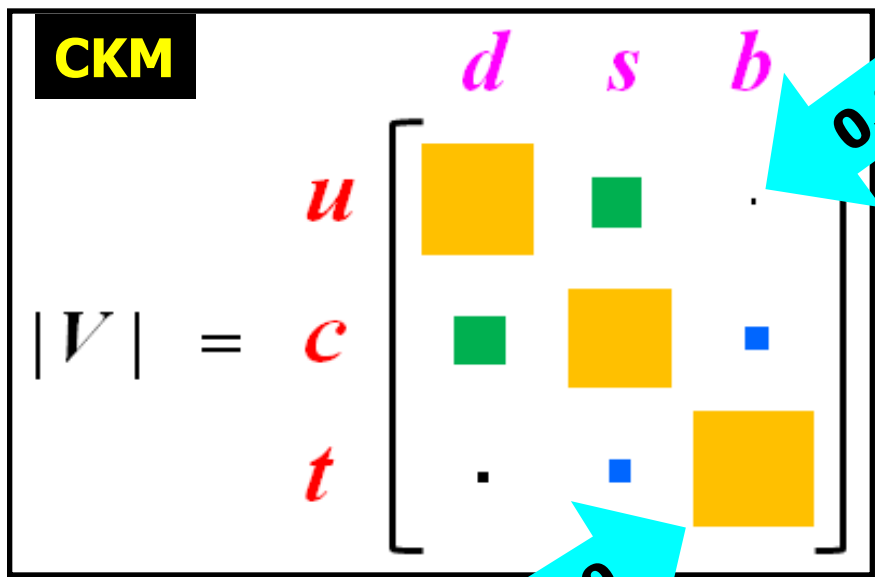
# Flavor mixing puzzle

$$-\mathcal{L}_{cc} = \frac{g}{\sqrt{2}} \left[ \overline{(u \ c \ t)}_L \gamma^\mu \underset{\substack{\uparrow \\ \text{CKM}}}{V} \begin{pmatrix} d \\ s \\ b \end{pmatrix}_L W_\mu^+ + \overline{(e \ \mu \ \tau)}_L \gamma^\mu \underset{\substack{\uparrow \\ \text{PMNS}}}{U} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}_L W_\mu^- \right] + \text{h.c.}$$

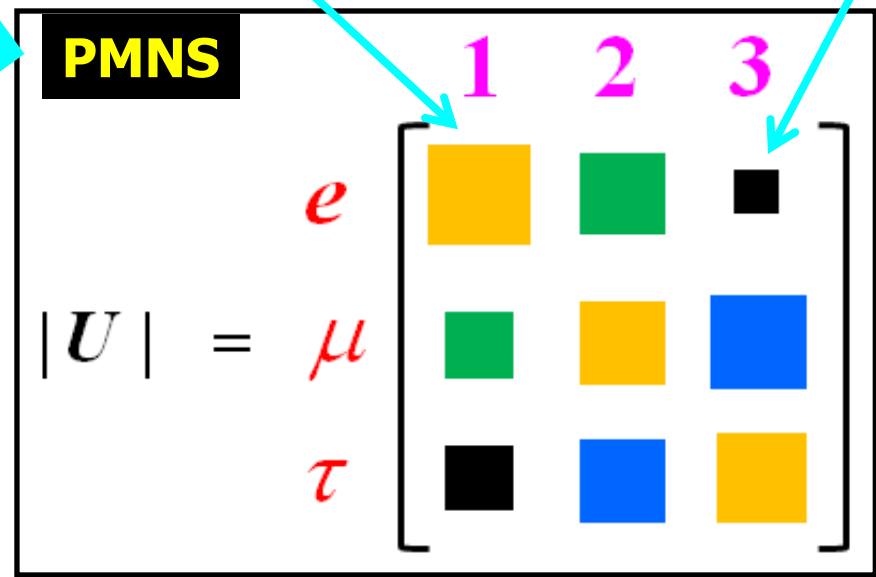
**CKM**

**PMNS**

Quark mixing: **hierarchy!**



4 parameters



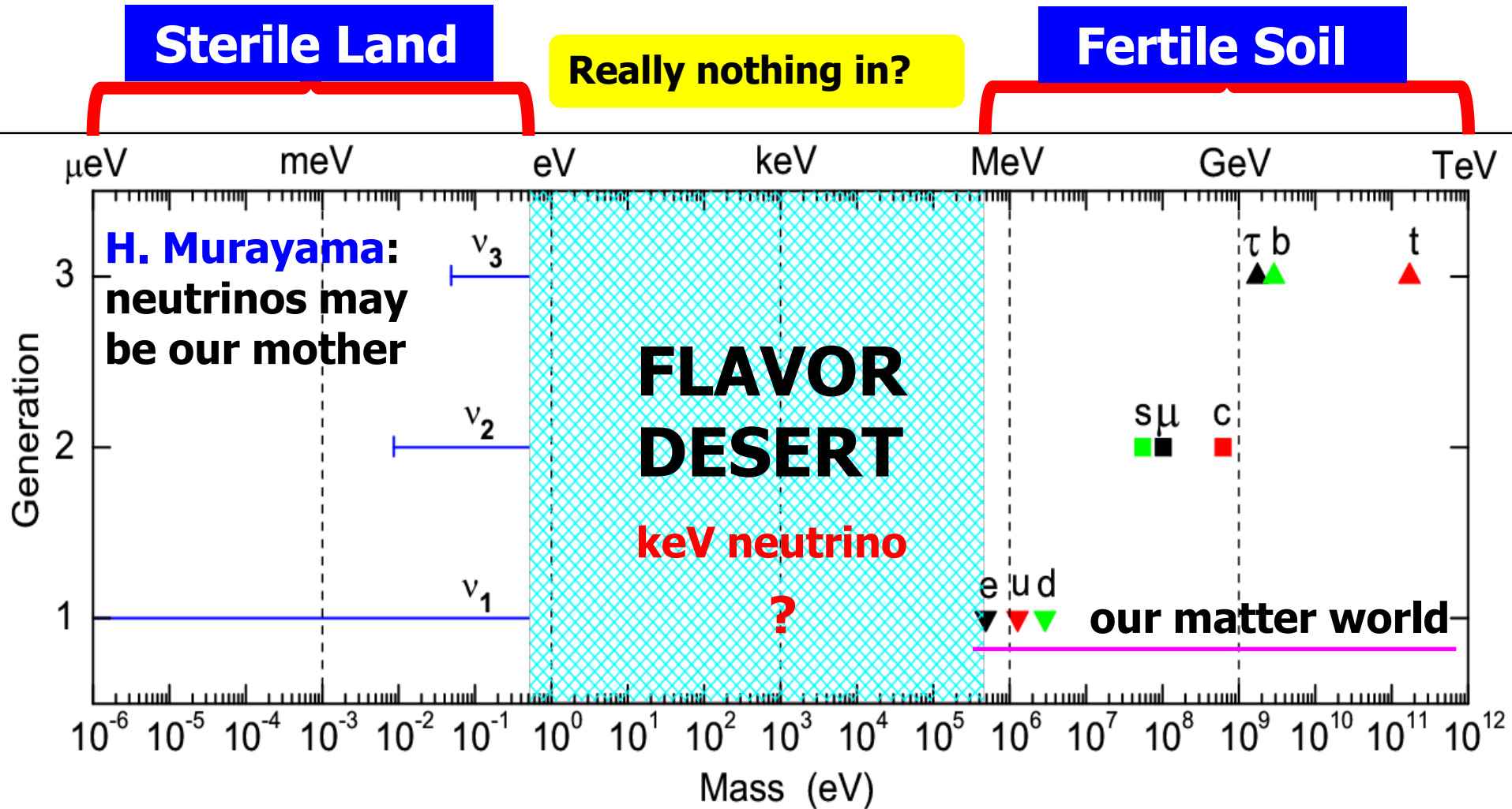
Lepton mixing: **anarchy?**

$\sim 0.8$  4/6 parameters  $\sim 0.15$



# Flavor mass puzzle

18



**Gauge Hierarchy & Desert Puzzles / Flavor Hierarchy & Desert Puzzles**

**Implications of electron mass < u quark mass < d quark mass on .....**

# What is behind?

19

**What** distinguishes different families of fermions?

----- they have the same gauge quantum numbers, yet they are quite different from one another, in their masses, flavor mixing strengths, .....

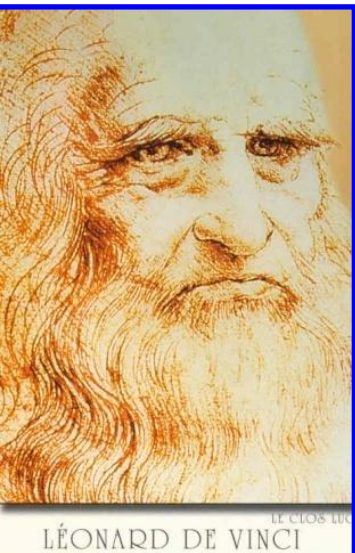


**We are blind today: no convincing predictive flavor theory**

**The structure of flavors should determine their properties**



**Bottom-Up Way**



**Although nature commences with reason and ends in experience, it is necessary for us to do the opposite, that is, to commence with experience and from this to proceed to investigate the reason**

**We will see: the minimal symmetry behind:  $\mu$ - $\tau$  symmetry!**

## Symmetries: crucial for understanding the laws of Nature.

Examples: they help simplify problems, classify complicated systems, fix conservation laws and even determine dynamics of interactions.

- Continuous space-time (translational/rotational) symmetries  
⇒ energy-momentum conservation laws
- Gauge symmetries ⇒ electroweak and strong interactions
- SU(3) quark flavor symmetry ⇒ the quark model ♣♣♣

Symmetries may keep **exact** or be **broken**: both important!

- Continuous space-time symmetries: **exact**
- U(1) electromagnetic gauge symmetry: **exact** (massless photon)
- SU(2) weak gauge symmetry: **broken** (massive  $W, Z, etc$ )
- SU(3) color gauge symmetry: **exact** (massless gluons)
- SU(3) quark flavor symmetry: **broken** ♣♣♣

# What the data tell?

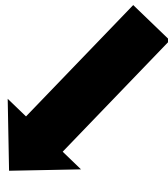
Given the global-fit results at the  $3\sigma$  level, the elements of the PMNS matrix are:

The normal ordering:

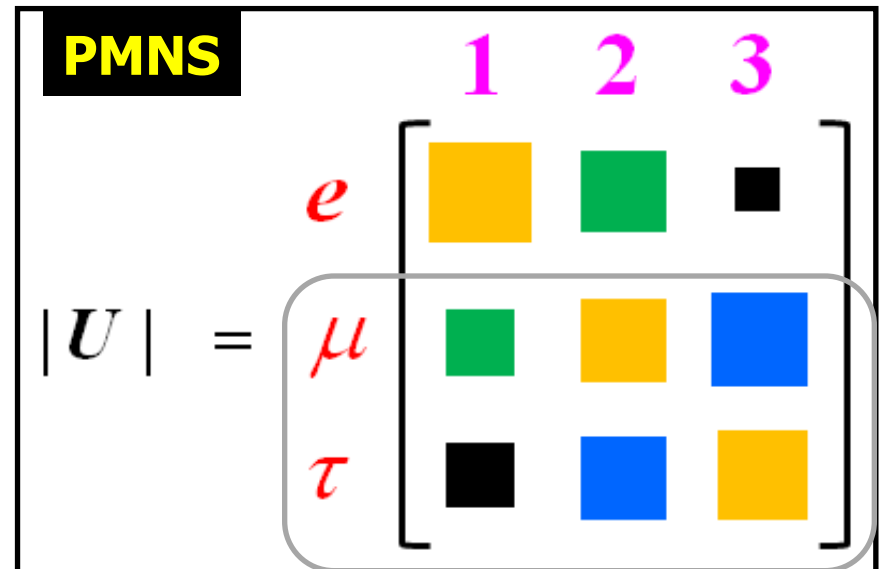
$$|U| \simeq \begin{pmatrix} 0.79 - 0.85 & 0.50 - 0.59 & 0.13 - 0.17 \\ 0.19 - 0.56 & 0.41 - 0.74 & 0.60 - 0.78 \\ 0.19 - 0.56 & 0.41 - 0.74 & 0.60 - 0.78 \end{pmatrix}$$

The inverted ordering:

$$|U| \simeq \begin{pmatrix} 0.89 - 0.85 & 0.50 - 0.59 & 0.13 - 0.17 \\ 0.19 - 0.56 & 0.40 - 0.73 & 0.61 - 0.79 \\ 0.20 - 0.56 & 0.41 - 0.74 & 0.59 - 0.78 \end{pmatrix}$$



$$\begin{array}{l} |U_{\mu 1}| \simeq |U_{\tau 1}| \\ |U_{\mu 2}| \simeq |U_{\tau 2}| \\ |U_{\mu 3}| \simeq |U_{\tau 3}| \end{array}$$



Behind the observed pattern of lepton flavor mixing is an **approximate** (or a **partial**)  $\mu$ - $\tau$  **flavor symmetry!**

$$|U_{\mu 1}| \simeq |U_{\tau 1}|, \quad |U_{\mu 2}| \simeq |U_{\tau 2}|, \quad |U_{\mu 3}| \simeq |U_{\tau 3}|$$



It is very likely that the **PMNS** matrix possesses an **exact**  $\mu$ - $\tau$  **symmetry** at a given energy scale, and this symmetry must be **softly broken** — shed light on flavor structures

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{13}s_{23}e^{i\delta} & c_{12}c_{23} - s_{12}s_{13}s_{23}e^{i\delta} & c_{13}s_{23} \\ s_{12}s_{23} - c_{12}s_{13}c_{23}e^{i\delta} & -c_{12}s_{23} - s_{12}s_{13}c_{23}e^{i\delta} & c_{13}c_{23} \end{pmatrix} P_\nu$$

**Conditions** for the **exact**  $\mu$ - $\tau$  **symmetry** in the **PMNS** matrix:

$$|U_{\mu i}| = |U_{\tau i}| \implies \begin{cases} \theta_{13} = 0 \\ \theta_{23} = \pi/4 \end{cases} \quad \text{or} \quad \begin{cases} \delta = +\pi/2 \\ \theta_{23} = \pi/4 \end{cases} \quad \text{or} \quad \begin{cases} \delta = -\pi/2 \\ \theta_{23} = \pi/4 \end{cases}$$

**Current data:**

**ruled out**

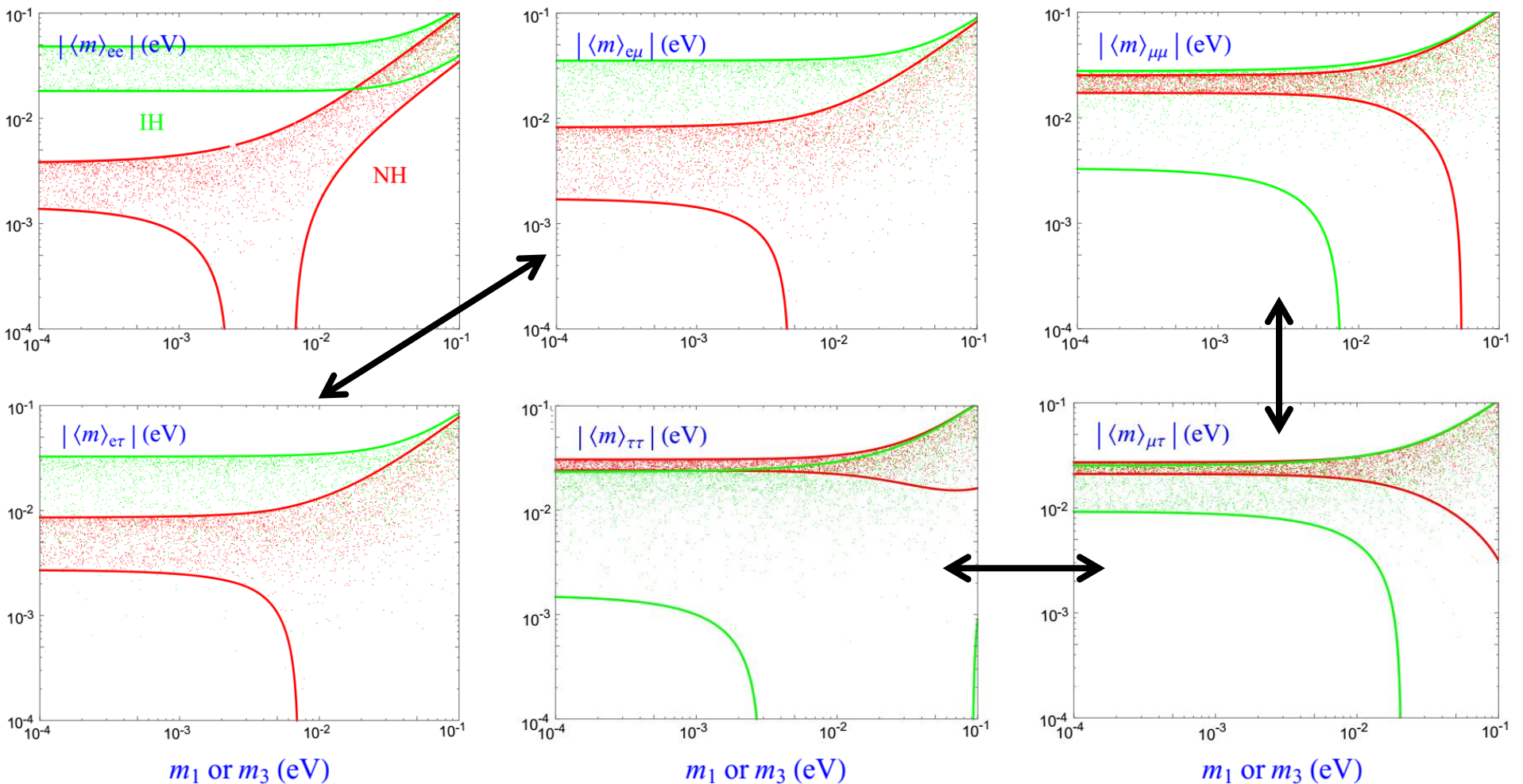
**not sure**

**avored**




# Neutrino mass matrix

$$M_\nu = \begin{pmatrix} M_{ee} & M_{e\mu} & M_{e\tau} \\ M_{e\mu} & M_{\mu\mu} & M_{\mu\tau} \\ M_{e\tau} & M_{\mu\tau} & M_{\tau\tau} \end{pmatrix} = U \begin{pmatrix} m_1 & 0 & 0 \\ 0 & m_2 & 0 \\ 0 & 0 & m_3 \end{pmatrix} U^T = \begin{pmatrix} \langle m \rangle_{ee} & \langle m \rangle_{e\mu} & \langle m \rangle_{e\tau} \\ \langle m \rangle_{e\mu} & \langle m \rangle_{\mu\mu} & \langle m \rangle_{\mu\tau} \\ \langle m \rangle_{e\tau} & \langle m \rangle_{\mu\tau} & \langle m \rangle_{\tau\tau} \end{pmatrix}$$



# $\mu$ - $\tau$ flavor symmetry

In the flavor basis, the **Majorana  $\nu$**  mass matrix can be reconstructed:


$$M_\nu = \begin{pmatrix} M_{ee} & M_{e\mu} & M_{e\tau} \\ M_{e\mu} & M_{\mu\mu} & M_{\mu\tau} \\ M_{e\tau} & M_{\mu\tau} & M_{\tau\tau} \end{pmatrix} = U \begin{pmatrix} m_1 & 0 & 0 \\ 0 & m_2 & 0 \\ 0 & 0 & m_3 \end{pmatrix} U^T$$


$\mu$ - $\tau$  permutation symmetry

$\mu$ - $\tau$  reflection symmetry

$$M_\nu = \begin{pmatrix} C & D & D \\ D & A & B \\ D & B & A \end{pmatrix}$$

$\nu_e \quad \nu_{\mu \leftrightarrow \tau}$




$$\begin{cases} \theta_{13} = 0 \\ \theta_{23} = \pi/4 \end{cases}$$

$$\frac{1}{2} \overline{(\nu_e \ \nu_\mu \ \nu_\tau)_L} M_\nu \begin{pmatrix} \nu_e^c \\ \nu_\mu^c \\ \nu_\tau^c \end{pmatrix}_R$$

$$M_\nu = \begin{pmatrix} C & D & D^* \\ D & A & B \\ D^* & B & A^* \end{pmatrix}$$

$\nu_e \quad \nu_{\mu \leftrightarrow \tau^c}$



$$\begin{cases} \delta = \pm\pi/2 \\ \theta_{23} = \pi/4 \end{cases}$$

Current data



**T. Fukuyama, H. Nishiura**  
hep-ph/9702253

**K. Babu, E. Ma, J. Valle**  
hep-ph/0206292

**Bimaximal, Tribimaximal ...**

**TM1, Tetramaximal ...**

**Larger**



$\mu$ - $\tau$  symmetry breaking



**Softer**

# A proof: permutation

A generic (symmetric) Majorana neutrino mass term reads as follows:

$$\begin{aligned} -\mathcal{L}_{\text{mass}} = & M_{ee}\overline{\nu_{eL}}(\nu_{eL})^c + M_{e\mu}\overline{\nu_{eL}}(\nu_{\mu L})^c + M_{e\tau}\overline{\nu_{eL}}(\nu_{\tau L})^c \\ & + M_{e\mu}\overline{\nu_{\mu L}}(\nu_{eL})^c + \underline{M_{\mu\mu}\overline{\nu_{\mu L}}(\nu_{\mu L})^c} + M_{\mu\tau}\overline{\nu_{\mu L}}(\nu_{\tau L})^c \\ & + M_{e\tau}\overline{\nu_{\tau L}}(\nu_{eL})^c + \underline{M_{\mu\tau}\overline{\nu_{\tau L}}(\nu_{\mu L})^c} + M_{\tau\tau}\overline{\nu_{\tau L}}(\nu_{\tau L})^c + \text{h.c.} \end{aligned}$$

Under  $\mu$ - $\tau$  permutation, the above term changes to

$$\begin{aligned} -\mathcal{L}_{\text{mass}} = & M_{ee}\overline{\nu_{eL}}(\nu_{eL})^c + M_{e\mu}\overline{\nu_{eL}}(\nu_{\tau L})^c + \underline{M_{e\tau}\overline{\nu_{eL}}(\nu_{\mu L})^c} \\ & + M_{e\mu}\overline{\nu_{\tau L}}(\nu_{eL})^c + M_{\mu\mu}\overline{\nu_{\tau L}}(\nu_{\tau L})^c + M_{\mu\tau}\overline{\nu_{\tau L}}(\nu_{\mu L})^c \\ & + M_{e\tau}\overline{\nu_{\mu L}}(\nu_{eL})^c + M_{\mu\tau}\overline{\nu_{\mu L}}(\nu_{\tau L})^c + \underline{M_{\tau\tau}\overline{\nu_{\mu L}}(\nu_{\mu L})^c} + \text{h.c.} \end{aligned}$$

$$\nu_{\mu L} \leftrightarrow \nu_{\tau L}$$

Invariance of this transformation requires:  $\underline{M_{e\mu} = M_{e\tau}}$  and  $\underline{M_{\mu\mu} = M_{\tau\tau}}$



$$M_\nu = \begin{pmatrix} C & D & D \\ D & A & B \\ D & B & A \end{pmatrix}$$

$\nu_e \quad \nu_\mu \leftrightarrow \nu_\tau$



$$\begin{cases} \theta_{13} = 0 \\ \theta_{23} = \pi/4 \end{cases}$$

# reflection

A generic Majorana neutrino mass term reads as follows:

Under  $\mu$ - $\tau$  reflection, the mass term is

$$\begin{aligned} \nu_{eL} &\leftrightarrow (\nu_{eL})^c \\ \nu_{\mu L} &\leftrightarrow (\nu_{\tau L})^c \\ \nu_{\tau L} &\leftrightarrow (\nu_{\mu L})^c \end{aligned}$$

Invariance of this transformation:

$$\begin{aligned} M_{ee} &= M_{ee}^* \\ M_{\mu\tau} &= M_{\mu\tau}^* \\ M_{e\mu} &= M_{e\tau}^* \\ M_{\mu\mu} &= M_{\tau\tau}^* \end{aligned}$$

$$\begin{aligned} -\mathcal{L}_{\text{mass}} &= \underline{M_{ee}\overline{\nu_{eL}}(\nu_{eL})^c} + \underline{M_{e\mu}\overline{\nu_{eL}}(\nu_{\mu L})^c} + M_{e\tau}\overline{\nu_{eL}}(\nu_{\tau L})^c \\ &+ M_{e\mu}\overline{\nu_{\mu L}}(\nu_{eL})^c + \underline{M_{\mu\mu}\overline{\nu_{\mu L}}(\nu_{\mu L})^c} + \underline{M_{\mu\tau}\overline{\nu_{\mu L}}(\nu_{\tau L})^c} \\ &+ M_{e\tau}\overline{\nu_{\tau L}}(\nu_{eL})^c + \underline{M_{\mu\tau}\overline{\nu_{\tau L}}(\nu_{\mu L})^c} + \underline{M_{\tau\tau}\overline{\nu_{\tau L}}(\nu_{\tau L})^c} \\ &+ M_{ee}^*\overline{(\nu_{eL})^c}\nu_{eL} + M_{e\mu}^*\overline{(\nu_{\mu L})^c}\nu_{eL} + M_{e\tau}^*\overline{(\nu_{\tau L})^c}\nu_{eL} \\ &+ M_{e\mu}^*\overline{(\nu_{eL})^c}\nu_{\mu L} + M_{\mu\mu}^*\overline{(\nu_{\mu L})^c}\nu_{\mu L} + M_{\mu\tau}^*\overline{(\nu_{\tau L})^c}\nu_{\mu L} \\ &+ M_{e\tau}^*\overline{(\nu_{eL})^c}\nu_{\tau L} + M_{\mu\tau}^*\overline{(\nu_{\mu L})^c}\nu_{\tau L} + M_{\tau\tau}^*\overline{(\nu_{\tau L})^c}\nu_{\tau L} \end{aligned}$$

$$\begin{aligned} -\mathcal{L}_{\text{mass}} &= M_{ee}\overline{(\nu_{eL})^c}\nu_{eL} + M_{e\mu}\overline{(\nu_{eL})^c}\nu_{\tau L} + M_{e\tau}\overline{(\nu_{eL})^c}\nu_{\mu L} \\ &+ M_{e\mu}\overline{(\nu_{\tau L})^c}\nu_{eL} + M_{\mu\mu}\overline{(\nu_{\tau L})^c}\nu_{\tau L} + M_{\mu\tau}\overline{(\nu_{\tau L})^c}\nu_{\mu L} \\ &+ M_{e\tau}\overline{(\nu_{\mu L})^c}\nu_{eL} + M_{\mu\tau}\overline{(\nu_{\mu L})^c}\nu_{\tau L} + M_{\tau\tau}\overline{(\nu_{\mu L})^c}\nu_{\mu L} \\ &+ \underline{M_{ee}^*\overline{\nu_{eL}}(\nu_{eL})^c} + M_{e\mu}^*\overline{\nu_{\tau L}}(\nu_{eL})^c + M_{e\tau}^*\overline{\nu_{\mu L}}(\nu_{eL})^c \\ &+ M_{e\mu}^*\overline{\nu_{eL}}(\nu_{\tau L})^c + M_{\mu\mu}^*\overline{\nu_{\tau L}}(\nu_{\tau L})^c + \underline{M_{\mu\tau}^*\overline{\nu_{\mu L}}(\nu_{\tau L})^c} \\ &+ \underline{M_{e\tau}^*\overline{\nu_{eL}}(\nu_{\mu L})^c} + M_{\mu\tau}^*\overline{\nu_{\tau L}}(\nu_{\mu L})^c + \underline{M_{\tau\tau}^*\overline{\nu_{\mu L}}(\nu_{\mu L})^c} \end{aligned}$$

$$M_\nu = \begin{pmatrix} C & D & D^* \\ D & A & B \\ D^* & B & A^* \end{pmatrix} \begin{matrix} \nu_e \\ \nu_\mu \leftrightarrow \nu_\tau^c \end{matrix} \longrightarrow \begin{cases} \delta = \pm\pi/2 \\ \theta_{23} = \pi/4 \end{cases}$$

The **flavor symmetry** is a powerful **guiding principle** of model building.

The **flavor symmetry** could be

- ♣ Abelian or non-Abelian
- ♣ Continuous or discrete
- ♣ Local or global
- ♣ Spontaneously or explicitly broken

$S_3 / S_4 / A_4 / Z_2 /$   
 $U(1)_F / SU(2)_F / \dots$



**Advantages** of choosing a **global + discrete** flavor symmetry group  $G_F$ .

- ♣ No Goldstone bosons
- ♣ No additional gauge bosons mediating harmful FCNC processes
- ♣ No family-dependent D-terms contributing to sfermion masses
- ♣ Discrete  $G_F$  could come from some string compactifications
- ♣ Discrete  $G_F$  could be embedded in a continuous symmetry group

**SUSY**



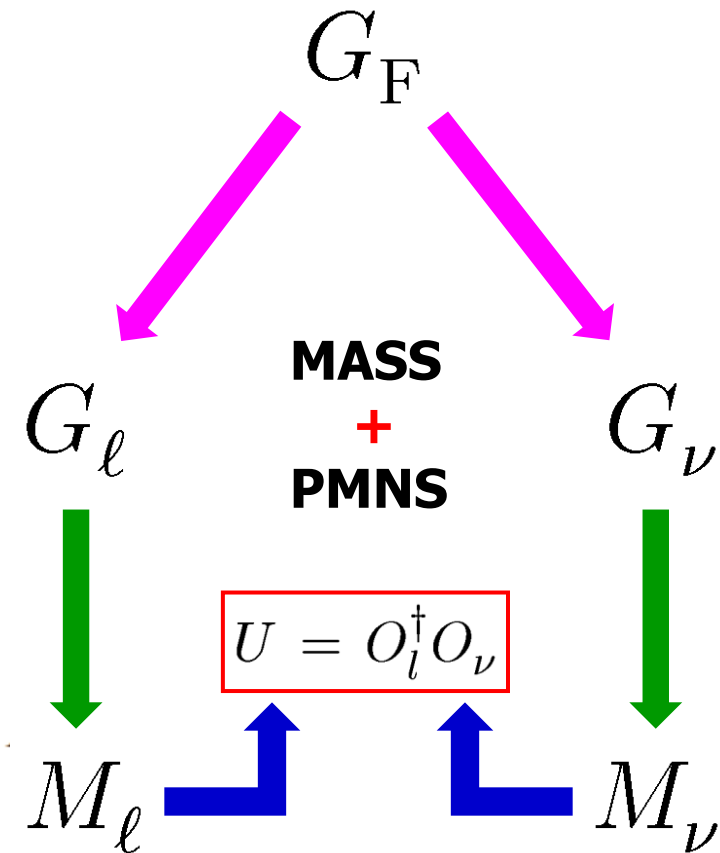
# Flavor symmetry groups

28

Some small **discrete groups** for model building (Altarelli, Feruglio **2010**).

Group	$d$	Irreducible representation
$D_3 \sim S_3$	6	$1, 1', 2$
$D_4$	8	$1_1, \dots, 1_4, 2$
$D_7$	14	$1, 1', 2, 2', 2''$
$A_4$	12	$1, 1', 1'', 3$
$A_5 \sim PSL_2(5)$	60	$1, 3, 3', 4, 5$
$T'$	24	$1, 1', 1'', 2, 2', 2'', 3$
$S_4$	24	$1, 1', 2, 3, 3'$
$\Delta(27) \sim Z_3 \rtimes Z_3$	27	$1_1, 1_9, 3, \bar{3}$
$PSL_2(7)$	168	$1, 3, \bar{3}, 6, 7, 8$
$T_7 \sim Z_7 \rtimes Z_3$	21	$1, 1', \bar{1}', 3, \bar{3}$

Too many possibilities, but the  $\mu$ - $\tau$  symmetry inclusive



**Generalized CP** combined with flavor symmetry to predict the phase  $\delta$ .

**Matter effects:** the behavior of neutrino oscillations is modified due to the coherent forward scattering induced by the weak charged-current interactions. The effective Hamiltonian for neutrino propagation:

$$\tilde{\mathcal{H}}_{\text{eff}} = \frac{1}{2E} \left[ \tilde{U} \begin{pmatrix} \tilde{m}_1^2 & 0 & 0 \\ 0 & \tilde{m}_2^2 & 0 \\ 0 & 0 & \tilde{m}_3^2 \end{pmatrix} \tilde{U}^\dagger \right] = \frac{1}{2E} \left[ U \begin{pmatrix} m_1^2 & 0 & 0 \\ 0 & m_2^2 & 0 \\ 0 & 0 & m_3^2 \end{pmatrix} U^\dagger + \begin{pmatrix} A & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix} \right]$$

**in matter** **in vacuum** **correction**

**Sum rules between matter and vacuum:**

$$A = 2\sqrt{2} G_F N_e E$$

$$\sum_{i=1}^3 \tilde{m}_i^2 \tilde{U}_{\alpha i} \tilde{U}_{\beta i}^* = \sum_{i=1}^3 m_i^2 U_{\alpha i} U_{\beta i}^* + \underline{A \delta_{\alpha e} \delta_{e\beta}}$$

$$\sum_{i=1}^3 \tilde{m}_i^4 \tilde{U}_{\alpha i} \tilde{U}_{\beta i}^* = \sum_{i=1}^3 m_i^2 \left[ m_i^2 + \underline{A (\delta_{\alpha e} + \delta_{e\beta})} \right] U_{\alpha i} U_{\beta i}^* + \underline{A^2 \delta_{\alpha e} \delta_{e\beta}}$$

$$\sum_{i=1}^3 \tilde{U}_{\alpha i} \tilde{U}_{\beta i}^* = \sum_{i=1}^3 U_{\alpha i} U_{\beta i}^* = \delta_{\alpha\beta}$$

**disappear  
when  $\alpha, \beta$   
=  $\mu, \tau$**

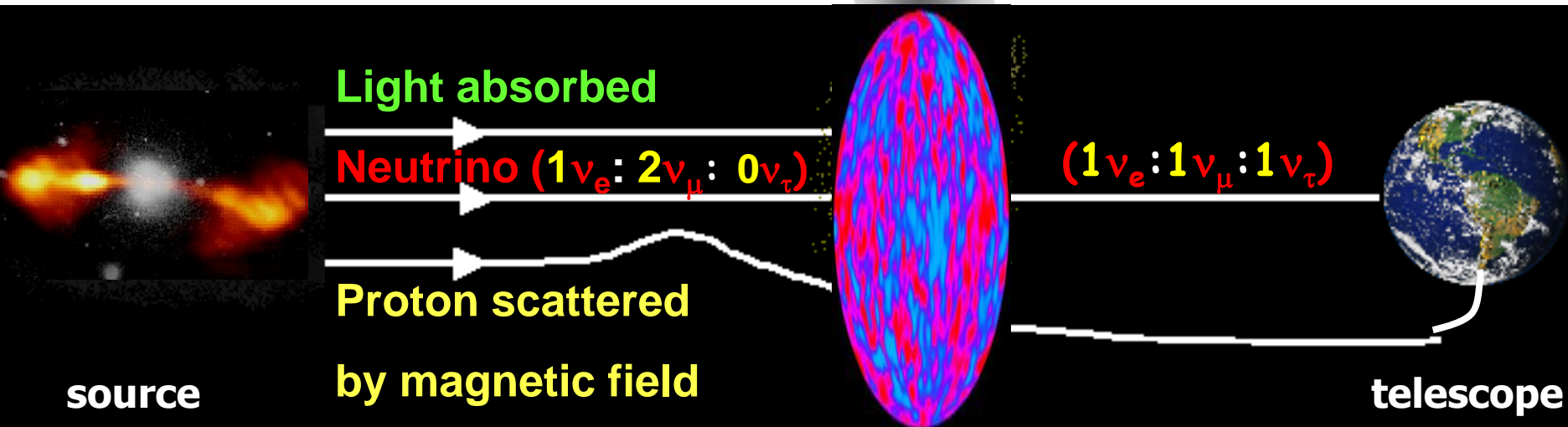
**A proper phase convention leads us to  $|\tilde{U}_{\mu i}| = |\tilde{U}_{\tau i}|$  from  $|U_{\mu i}| = |U_{\tau i}|$ .**

**Namely, matter effects (a constant profile) respect the  $\mu$ - $\tau$  symmetry.**

# Phenomenology (2)

30

**Ultrahigh-energy cosmic neutrinos** from distant astrophysical sources



**A conventional UHE cosmic neutrino source ( $p + p$  or  $p + \gamma$  collisions)**

$$\Phi_\mu^T - \Phi_\tau^T = \frac{\Phi_0}{3} \sum_i (|U_{\mu i}|^2 - |U_{\tau i}|^2)^2$$

$$\Phi_e^T : \Phi_\mu^T : \Phi_\tau^T = (1 + D_e) : (1 + D_\mu) : (1 + D_\tau)$$

with  $D_e = -2\Delta$ ,  $D_\mu = \Delta + \bar{\Delta}$  and  $D_\tau = \Delta - \bar{\Delta}$

$$\Delta \simeq \frac{1}{2} \sin^2 2\theta_{12} \sin \varepsilon - \frac{1}{4} \sin 4\theta_{12} \sin \theta_{13} \cos \delta$$

$$\varepsilon \equiv \theta_{23} - \pi/4$$

$$\bar{\Delta} \simeq (4 - \sin^2 2\theta_{12}) \sin^2 \varepsilon + \sin^2 2\theta_{12} \sin^2 \theta_{13} \cos^2 \delta + \sin 4\theta_{12} \sin \varepsilon \sin \theta_{13} \cos \delta$$

sensitive to  $\mu$ - $\tau$  flavor symmetry

# Summary

**Z.Z.X., Z.H. Zhao (1512.04207)**

**— A review of mu-tau flavor symmetry in neutrino physics**

**Report on Progress in Physics**

**79 (2016) 076201**



**C.S. Wu: It is easy to do the right thing once you have the **right ideas**.**

**I.I. Rabi: Physics needs new ideas. But to have a **new idea** is a very difficult task.... (Berezhiani's talk)**

**L.C. Pauling: The best way to have a **good idea** is to have a lot of ideas.**

## 1 Introduction

- 1.1 A brief history of the neutrino families . . . . .
- 1.2 The  $\mu$ - $\tau$  flavor symmetry stands out . . . . .

## 2 Behind the lepton flavor mixing pattern

- 2.1 Lepton flavor mixing and neutrino oscillations . . . . .
- 2.2 Current neutrino oscillation experiments . . . . .
- 2.3 The observed pattern of the PMNS matrix . . . . .

## 3 An overview of the $\mu$ - $\tau$ flavor symmetry

- 3.1 The  $\mu$ - $\tau$  permutation symmetry . . . . .
- 3.2 The  $\mu$ - $\tau$  reflection symmetry . . . . .
- 3.3 Breaking of the  $\mu$ - $\tau$  permutation symmetry . . . . .
- 3.4 Breaking of the  $\mu$ - $\tau$  reflection symmetry . . . . .
- 3.5 RGE-induced  $\mu$ - $\tau$  symmetry breaking effects . . . . .
- 3.6 Flavor mixing from the charged-lepton sector . . . . .

## 4 Larger flavor symmetry groups

- 4.1 Neutrino mixing and flavor symmetries . . . . .
- 4.2 Model building with discrete flavor symmetries . . . . .
- 4.3 Generalized CP and spontaneous CP violation . . . . .

## 5 Realization of the $\mu$ - $\tau$ flavor symmetry

- 5.1 Models with the  $\mu$ - $\tau$  permutation symmetry . . . . .
- 5.2 Models with the  $\mu$ - $\tau$  reflection symmetry . . . . .
- 5.3 On the TM1 and TM2 neutrino mixing patterns . . . . .
- 5.4 When the sterile neutrinos are concerned . . . . .

## 6 Some consequences of the $\mu$ - $\tau$ symmetry

- 6.1 Neutrino oscillations in matter . . . . .
- 6.2 Flavor distributions of UHE cosmic neutrinos . . . . .
- 6.3 Matter-antimatter asymmetry via leptogenesis . . . . .
- 6.4 Fermion mass matrices with the  $Z_2$  symmetry . . . . .

## 7 Summary and outlook

# Lecture A5

- ★ **Lepton number violation**
- ★ **neutrinoless double-beta decays**
- ★ **Possible new physics effects**

# 1935: $2\nu 2\beta$ decays

33

$2\nu 2\beta$  decay: certain **even-even** nuclei have an opportunity to decay to the 2nd nearest neighbors via 2 simultaneous  $\beta$  decays (equivalent to the decays of two neutrons).

necessary conditions:

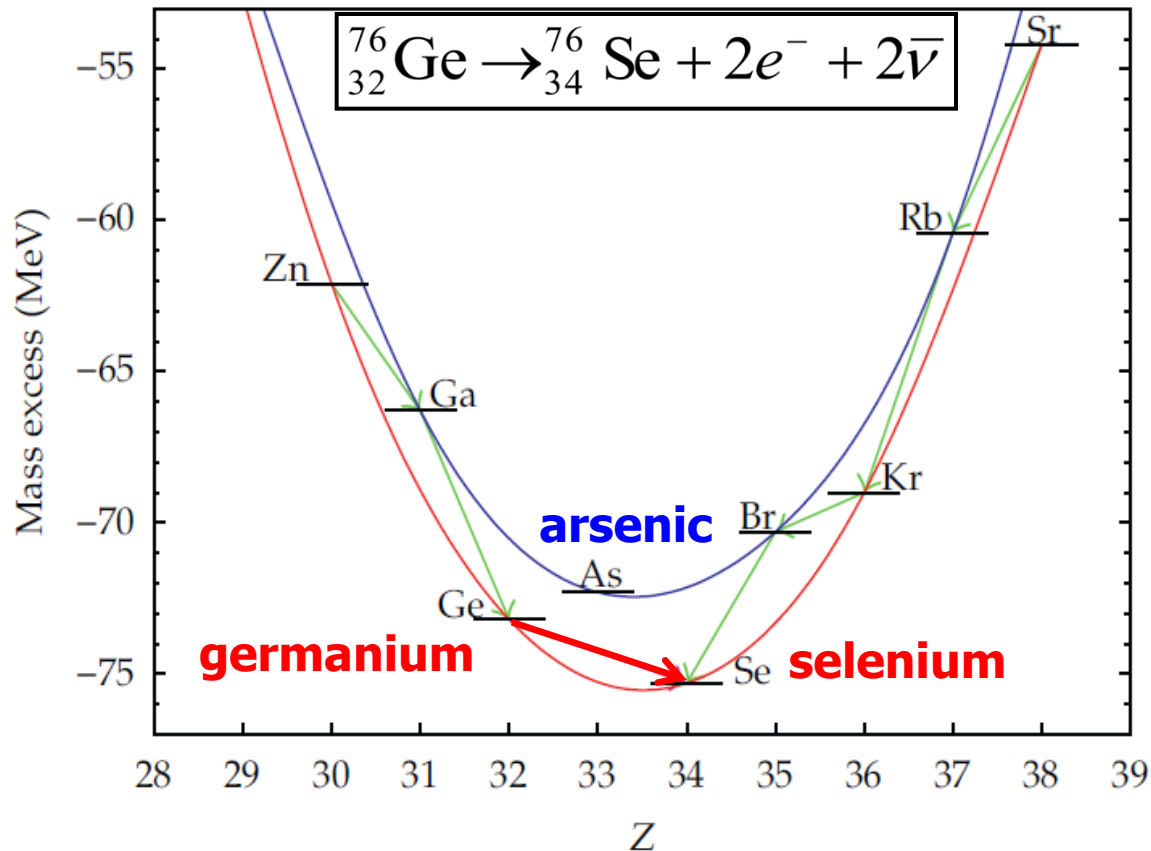
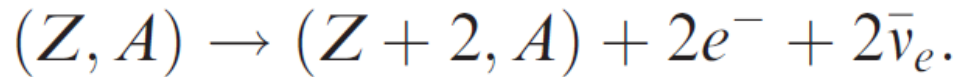
$$m(Z, A) > m(Z + 2, A)$$

$$m(Z, A) < m(Z + 1, A)$$



1935

Maria Goeppert Mayer





## ★ Theory of the Symmetry of Electrons and Positrons

**Ettore Majorana** *Nuovo Cim. 14 (1937) 171*

“...there is now no need to assume the existence of antineutron or antineutrinos. The latter particles are indeed introduced in the theory of positive beta-ray emission; the theory, however, can be obviously modified so that the beta-emission, both positive and negative, is always accompanied by the emission of a neutrino.”

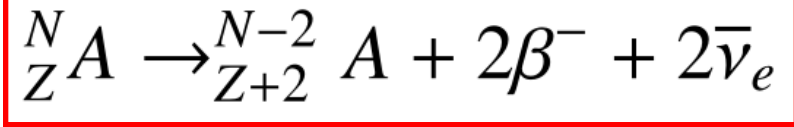
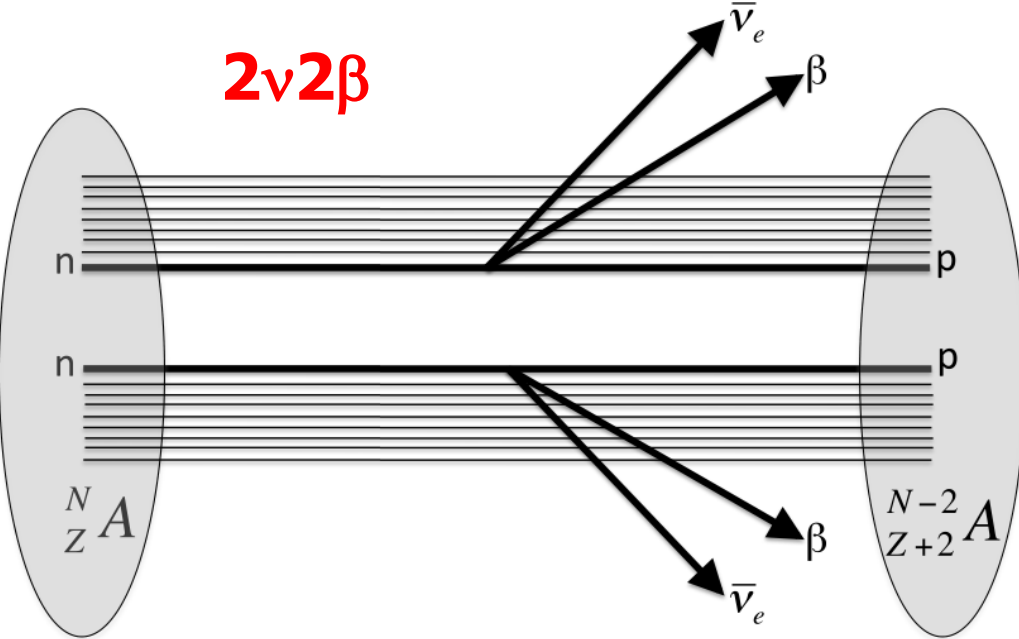
**1938年**：自杀；逃往阿根廷，并在那里隐姓埋名地生活了二十几年；遁入空门；遭到绑架或杀害，以阻止他加入制造原子弹的项目；沦为乞丐；.....

**Enrico Fermi (1938):**

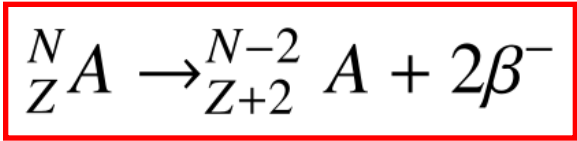
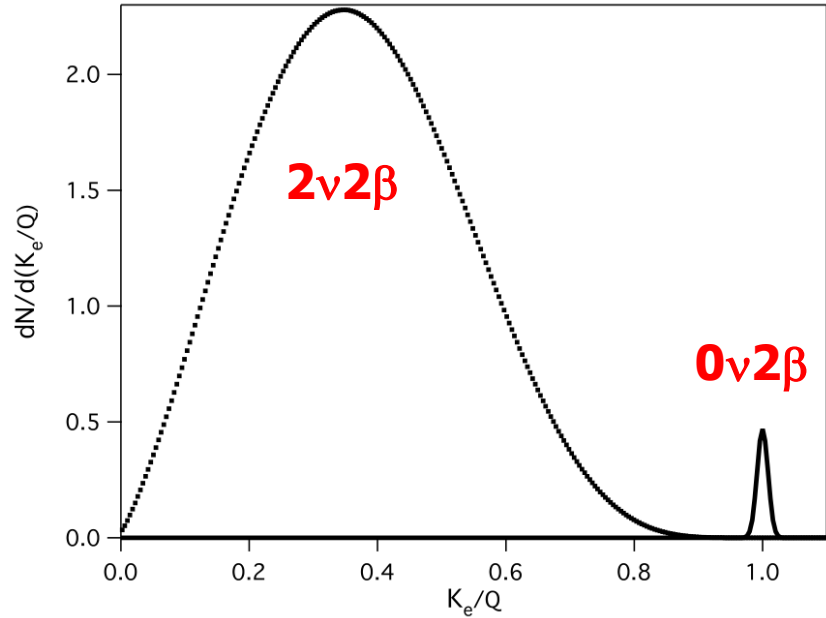
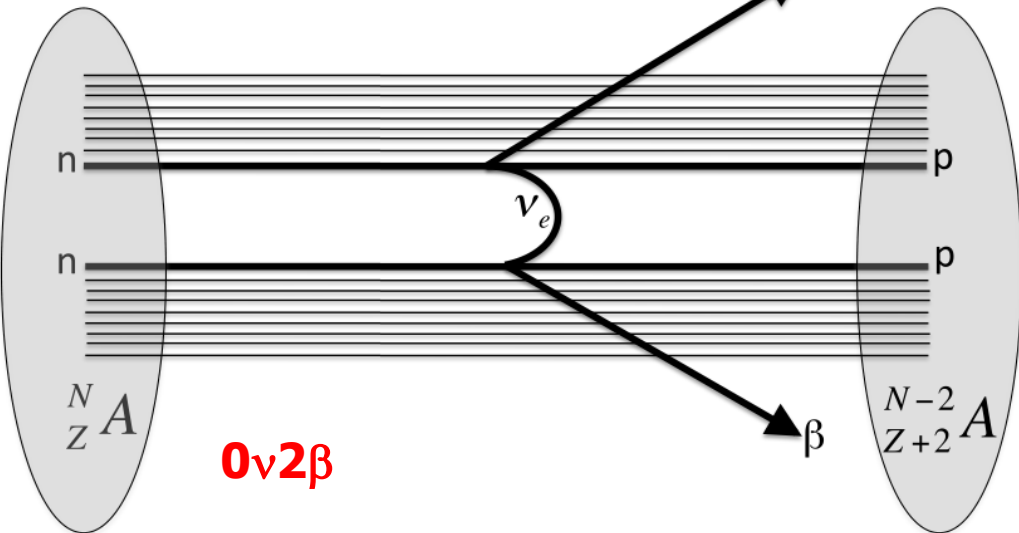
“There are various kind of scientists in the world. The second- and third-rate ones do their best but do not get very far. There are also first-rate people who make very important discoveries which are of capital importance for the development of the science. Then there are genius like Galileo and Newton. Ettore Majorana was one of these. Majorana had greater gifts than anyone else in the world; unfortunately he lacked one quality which other men generally have: plain common sense”

# If this is the case, ...

**2ν2β**



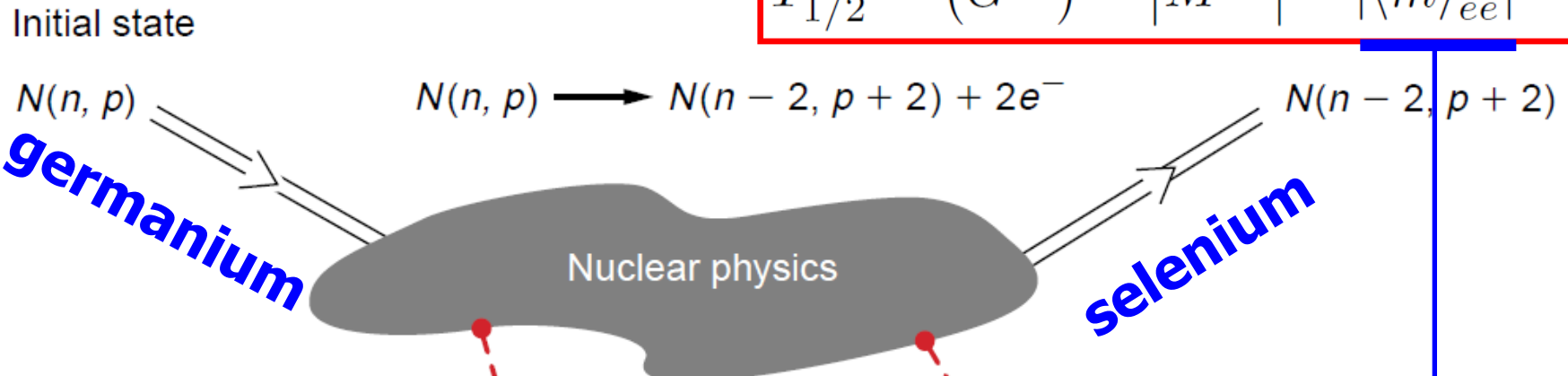
**0ν2β**



# 1939: $0\nu 2\beta$ decays

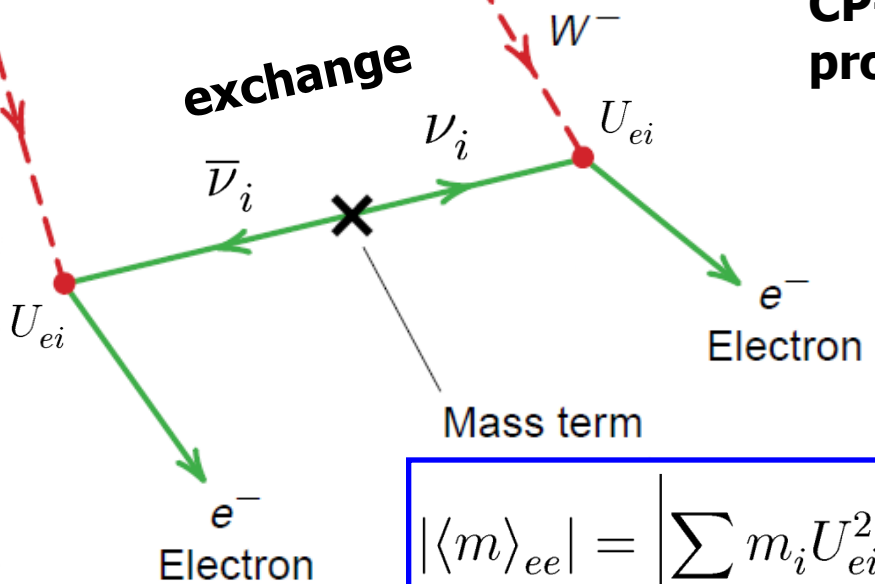
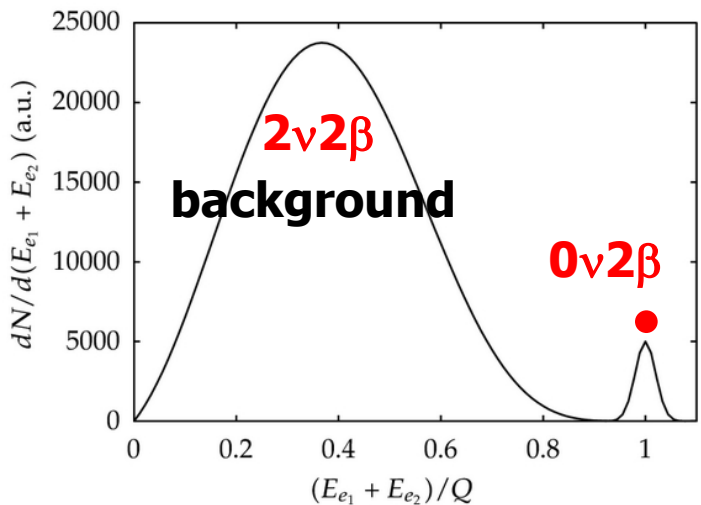
A  $0\nu 2\beta$  decay can happen if massive  $\nu$ 's have the Majorana nature (Wendell Furry 1939)

$$T_{1/2}^{0\nu} = (G^{0\nu})^{-1} |M^{0\nu}|^{-2} |\langle m \rangle_{ee}|^{-2}$$



Lepton number violation  $\longrightarrow$

CP-conserving process  $\longleftarrow$

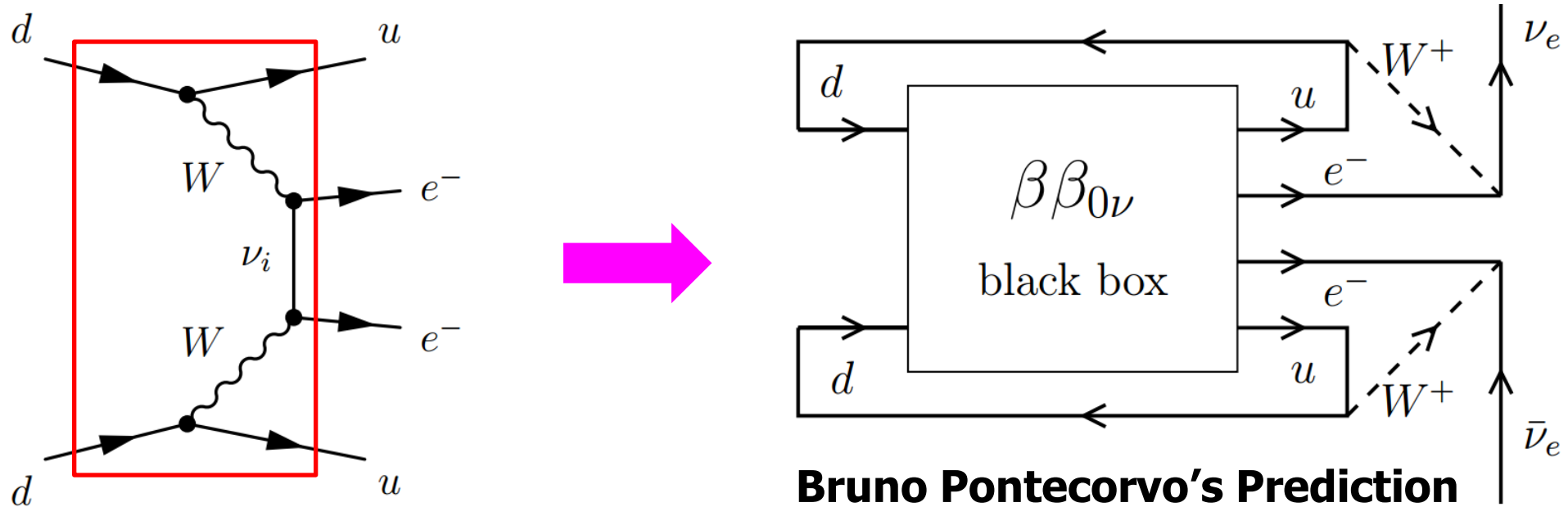


$$|\langle m \rangle_{ee}| = \left| \sum_i m_i U_{ei}^2 \right|$$

# Schechter-Valle theorem

37

**THEOREM (1982):** if a  $0\nu 2\beta$  decay happens, there must be an effective **Majorana** mass term.



**Bruno Pontecorvo's Prediction**

**That is why we want to see  $0\nu 2\beta$**

**Four-loop  $\nu$  mass:**

$$\delta m_\nu = \mathcal{O}(10^{-24} \text{ eV}) \quad (\text{Duerr, Lindner, Merle, 2011})$$

**Note:** The **black box** can in principle have many different processes (new physics). Only in the simplest case, which is most interesting, it's likely to constrain neutrino masses

**GERDA** has killed the **Heidelberg-Moscow's** claim on  $0\nu 2\beta$ .

PRL 111, 122503 (2013)

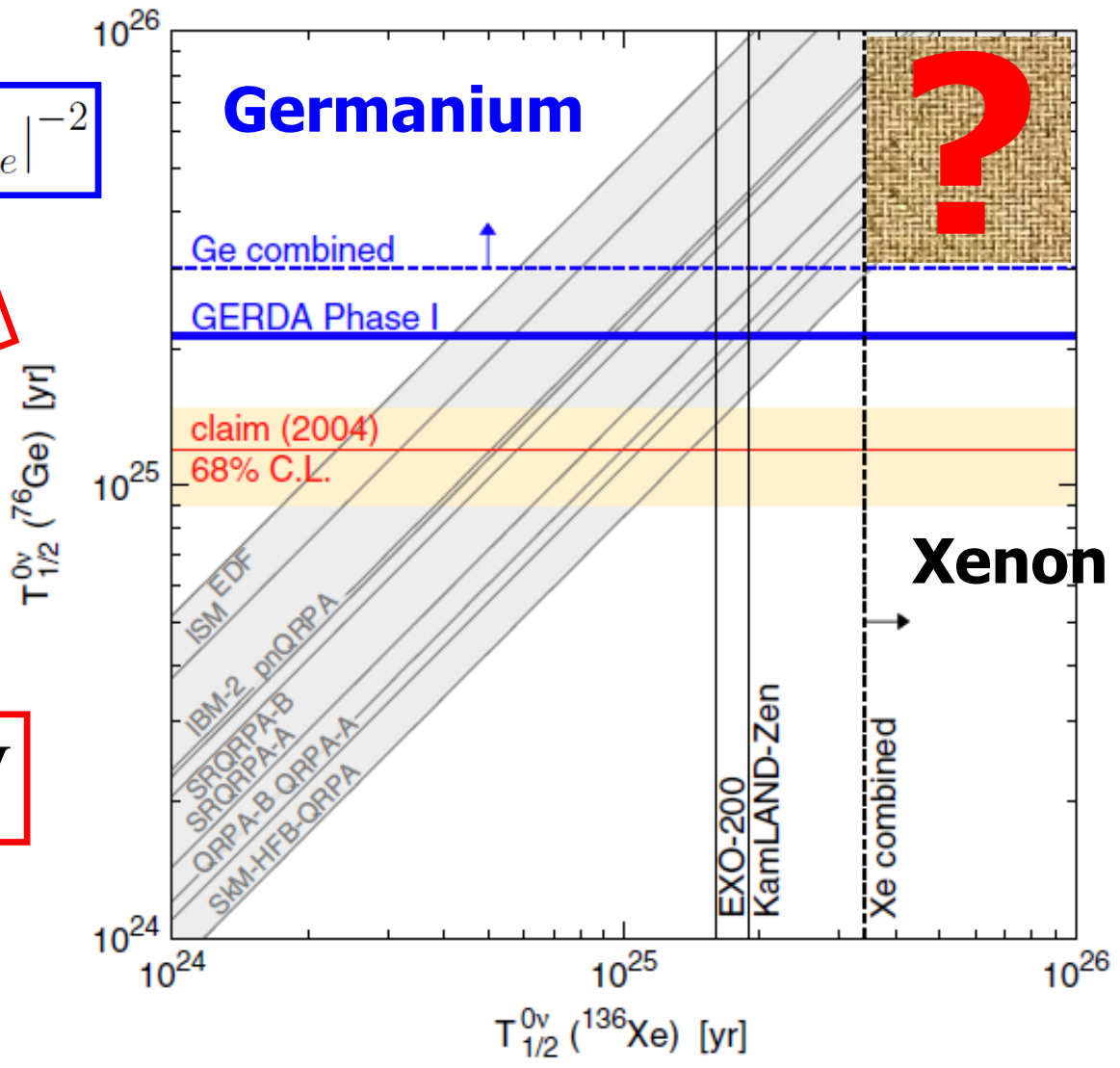
$$T_{1/2}^{0\nu} = (G^{0\nu})^{-1} |M^{0\nu}|^{-2} |\langle m \rangle_{ee}|^{-2}$$

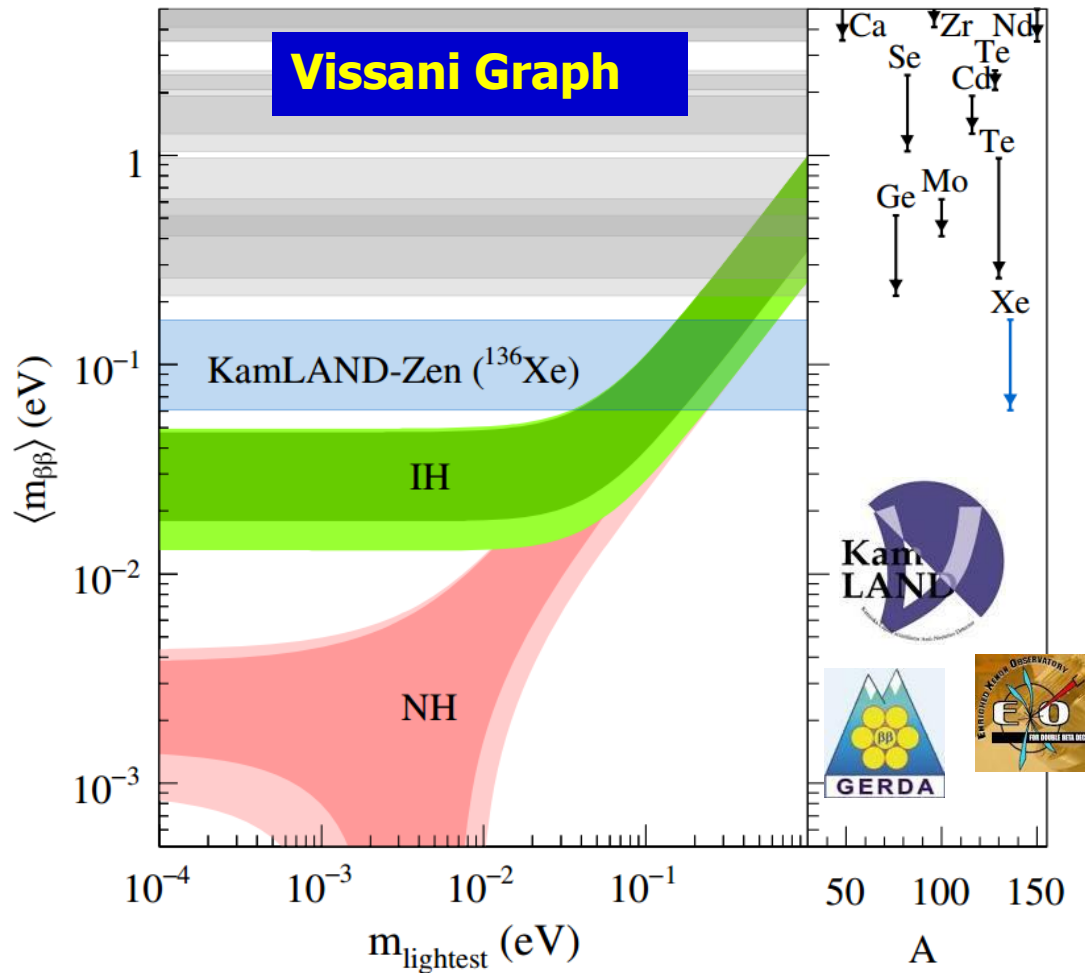
$T_{1/2}^{0\nu} > 3.0 \times 10^{25}$  yr (90% C.L.)



$$|\langle m \rangle_{ee}| < 0.2 \rightarrow 0.4 \text{ eV}$$

$$|\langle m \rangle_{ee}| = \left| \sum_i m_i U_{ei}^2 \right|$$





## The effective mass

$$|\langle m \rangle_{ee}| = \left| \sum_i m_i U_{ei}^2 \right|$$

Maury Goodman asks:  
**Is IH**  
 an intelligent design?



If it is **inverted**, why do not we reorder it?

Then how about the PMNS matrix?

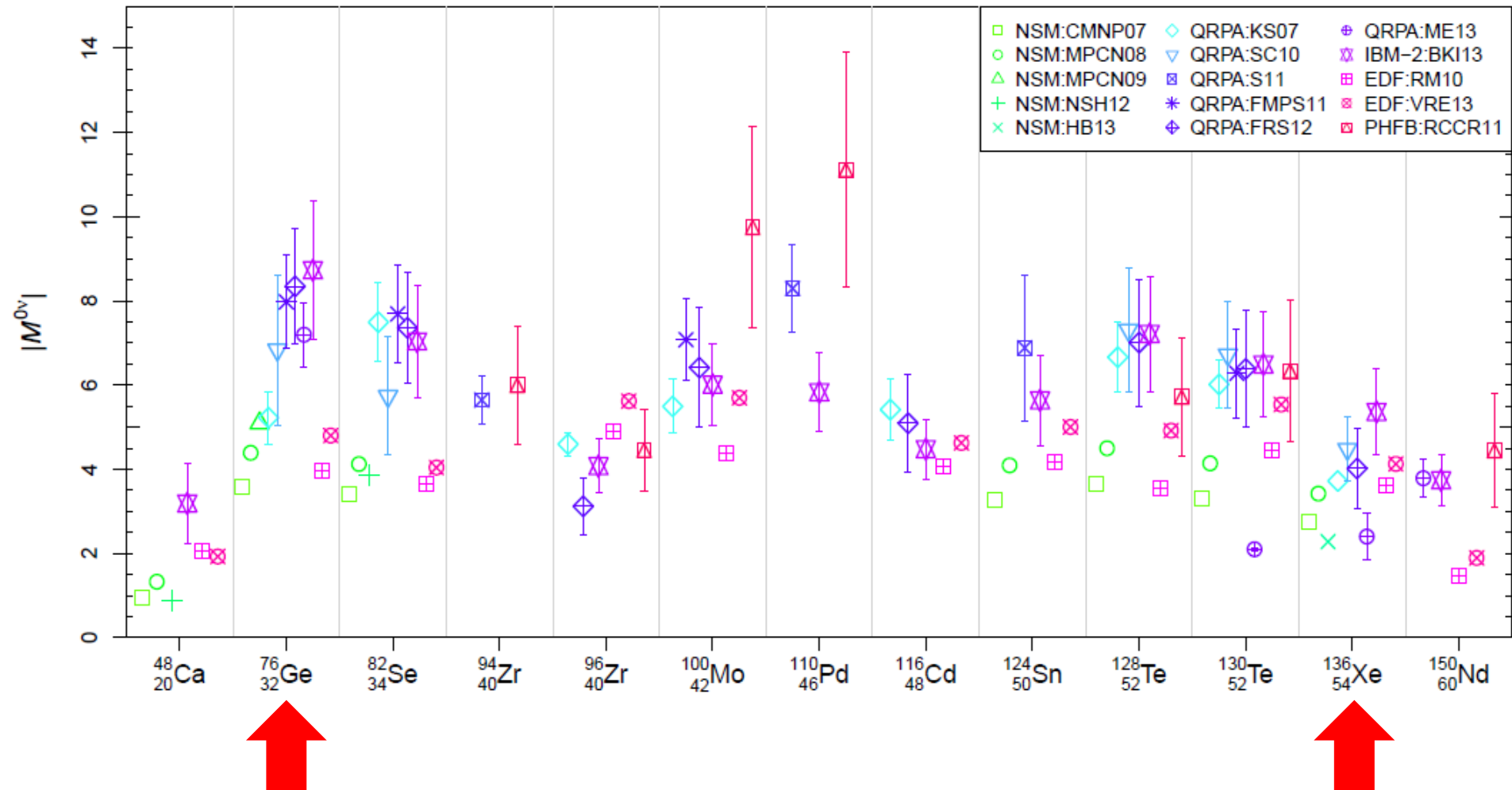
If **normal**, how possible to fall into the well?

- 1) The structure of the well?
- 2) Role of Majorana phases?

# Nuclear matrix elements

40

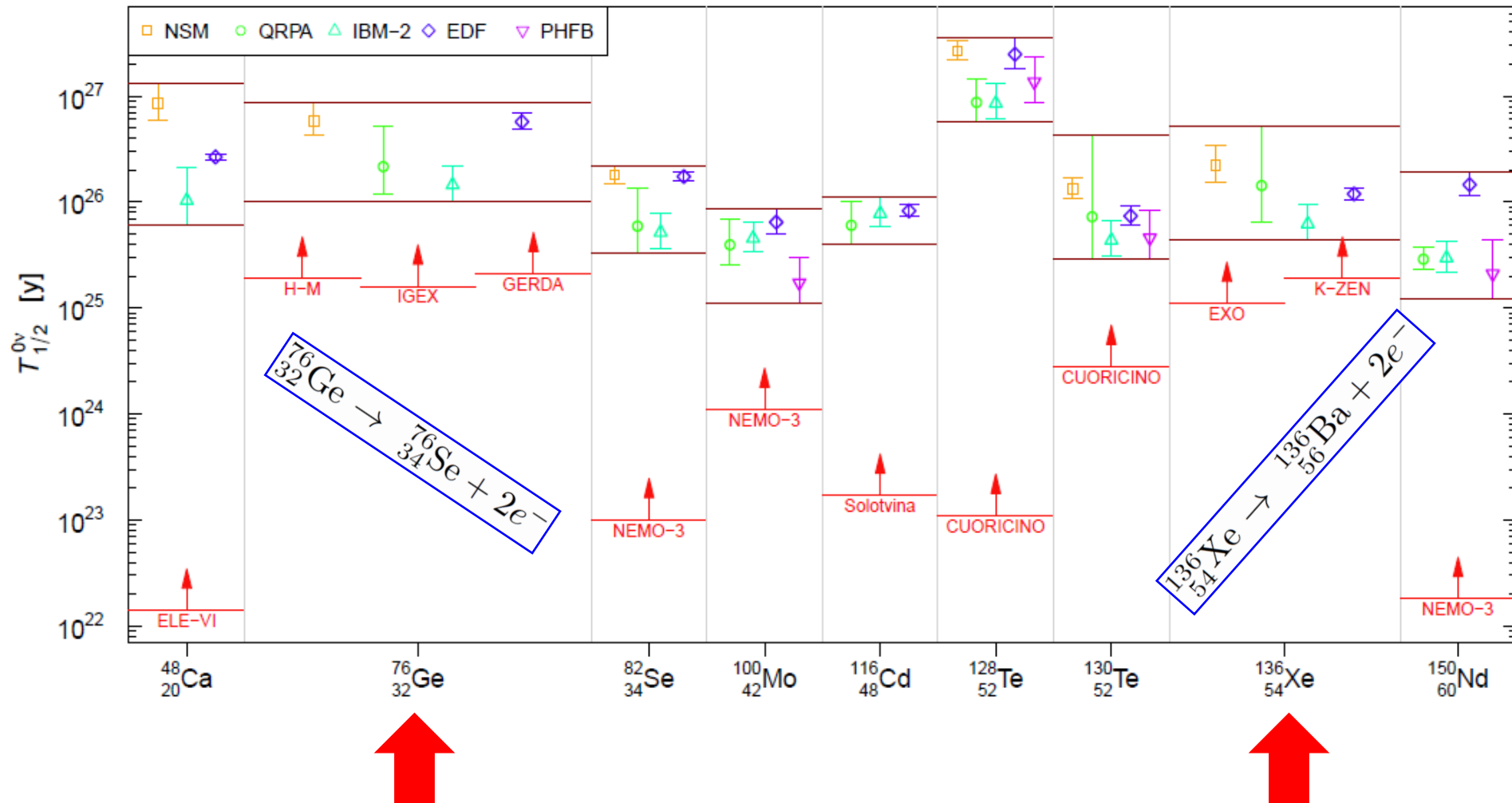
Unfortunately, nuclear matrix elements can be calculated only based on some models which describe many-body interactions of nucleons in nuclei. Since different models focus on different aspects of nuclear physics, **large uncertainties (a factor of 2 or 3)** are unavoidable.





# Half-life

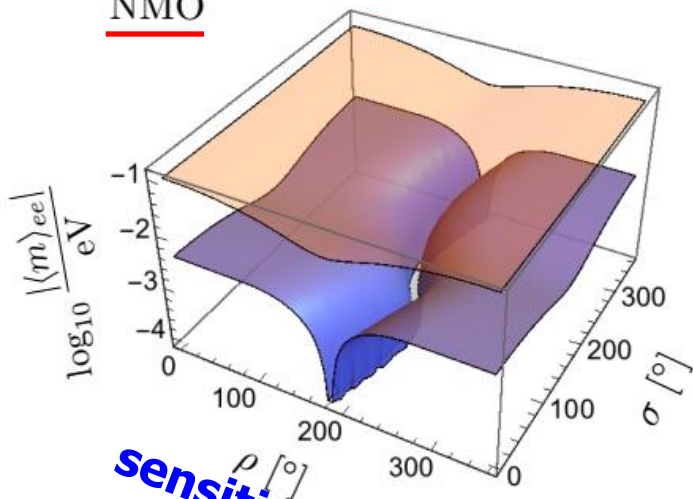
Comparing the 90% C.L. experimental lower limits on the half-life of a  $0\nu 2\beta$ -decaying nuclide with the corresponding range of theoretical prediction, given a value of  $0.1 \text{ eV}$  for the effective Majorana neutrino mass term (Bilenky and Giunti, 1411.4791).



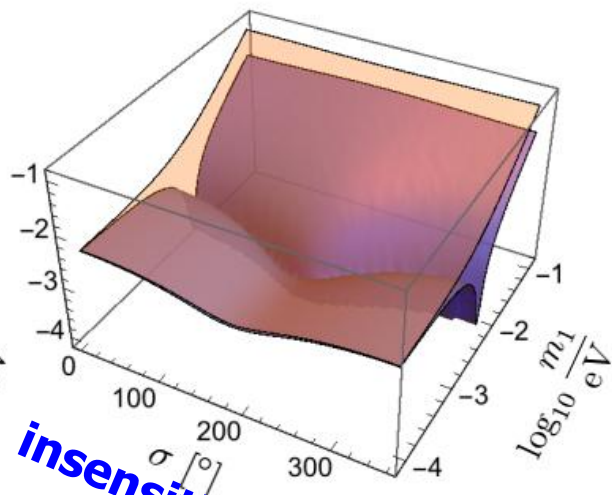
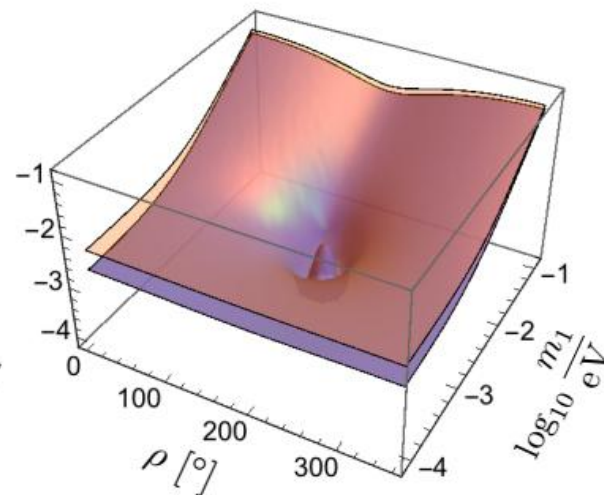
# Effective mass term

42

NMO



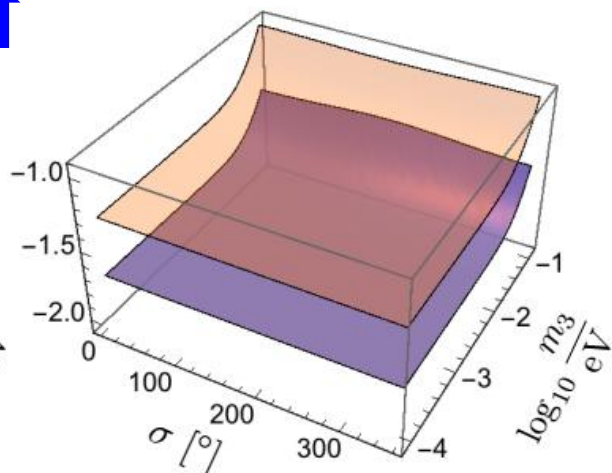
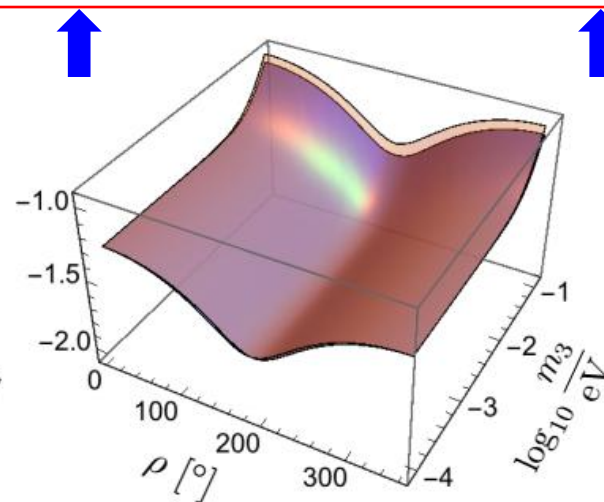
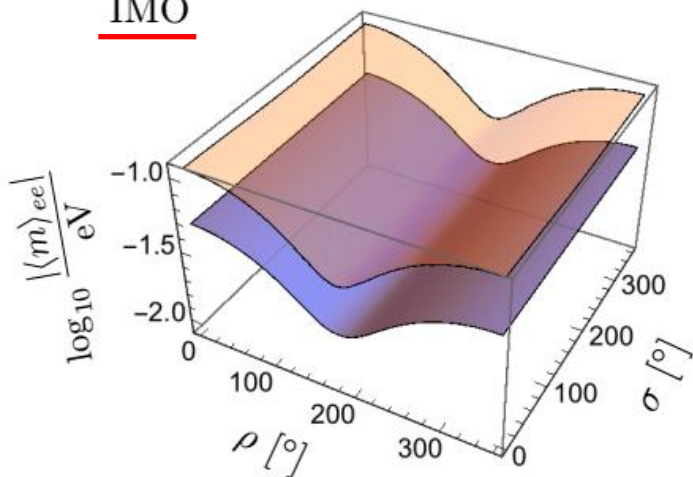
sensitive



insensitive

$$\langle m \rangle_{ee} = m_1 |U_{e1}|^2 e^{i\rho} + m_2 |U_{e2}|^2 + m_3 |U_{e3}|^2 e^{i\sigma}$$

IMO



Lower/upper bound: blue/light orange.  $3\sigma$  inputs of  $\nu$ -oscillation data with a new phase convention (Xing, Zhao, Zhou, 1504.05820)

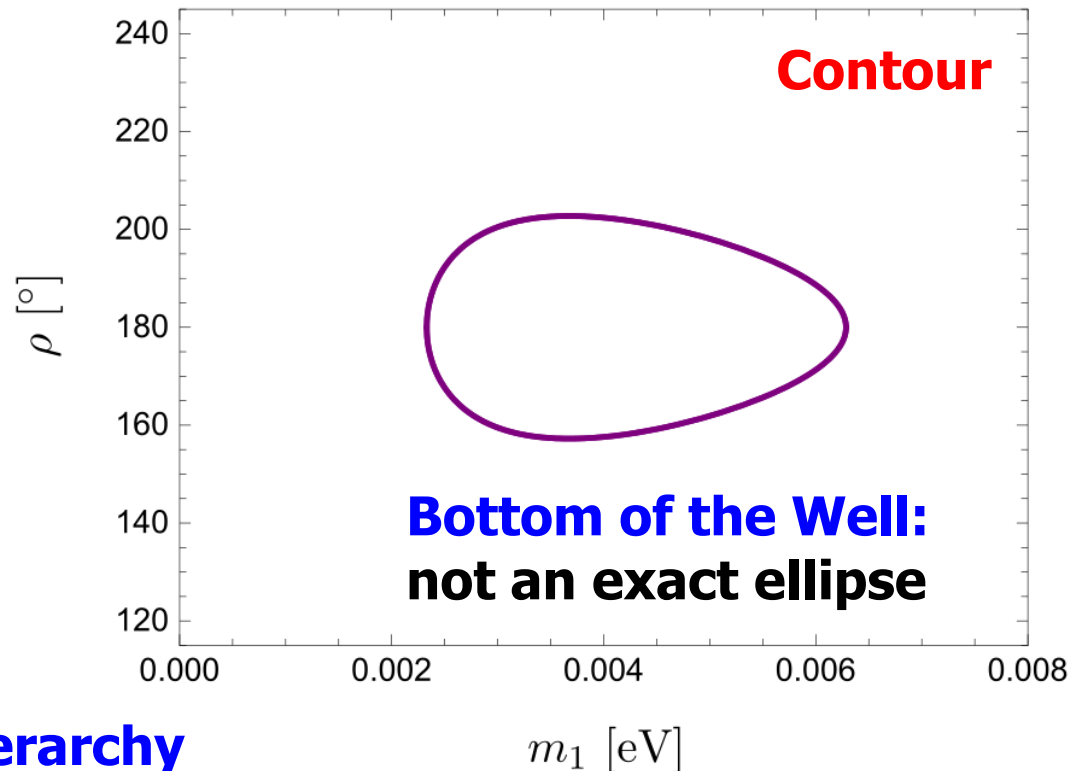
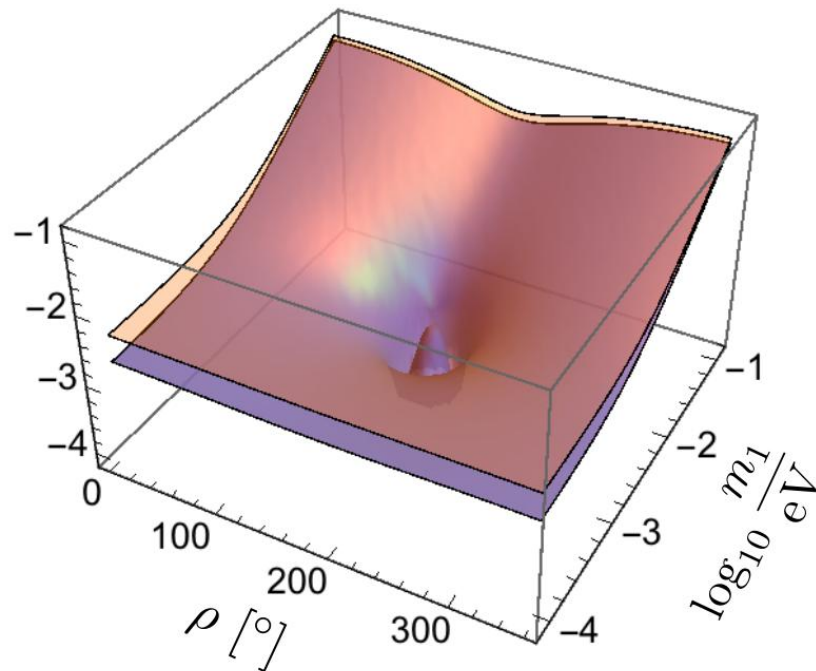
# Contour of the bottom

43

Let us understand the champagne-bottle profile of the effective  $0\nu 2\beta$  mass term in the **normal hierarchy** case:

$$\langle m \rangle_{ee} = m_1 c_{12}^2 c_{13}^2 e^{i\rho} + m_2 s_{12}^2 c_{13}^2 + m_3 s_{13}^2 e^{i\sigma} = 0$$

$$m_1^2 c_{12}^4 c_{13}^4 + 2m_1 m_2 c_{12}^2 s_{12}^2 c_{13}^4 \cos \rho + m_2^2 s_{12}^4 c_{13}^4 = m_3^2 s_{13}^4$$



The dark well in the **normal hierarchy**

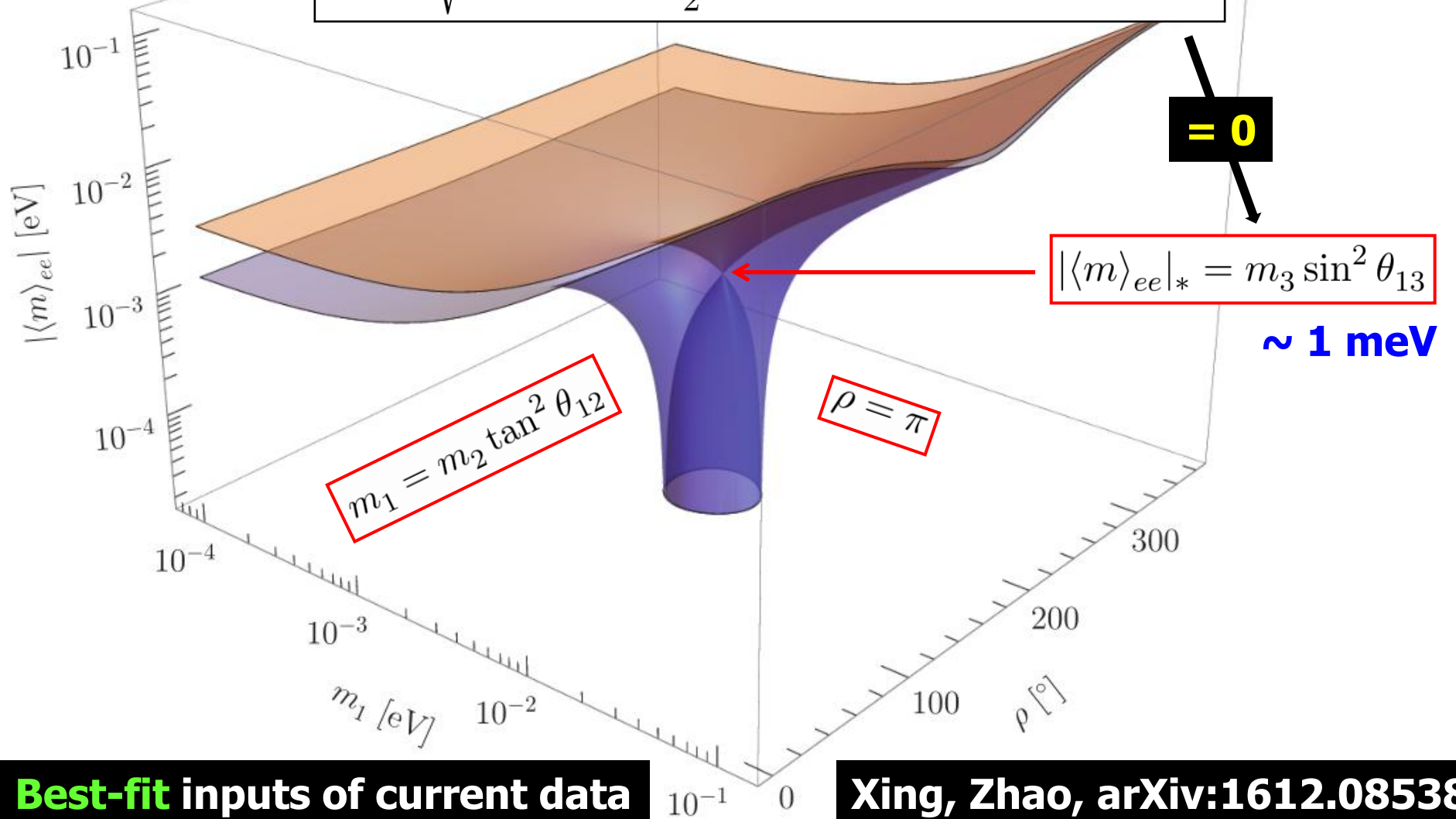
$m_1$  [eV]

# A bullet structure

**EXTREMUM**

$$|\langle m \rangle_{ee}|_{U,L} = |\bar{m}_{12} \cos^2 \theta_{13} \pm m_3 \sin^2 \theta_{13}|$$

$$\bar{m}_{12} \equiv \sqrt{m_1^2 \cos^4 \theta_{12} + \frac{1}{2} m_1 m_2 \sin^2 2\theta_{12} \cos \rho + m_2^2 \sin^4 \theta_{12}}$$

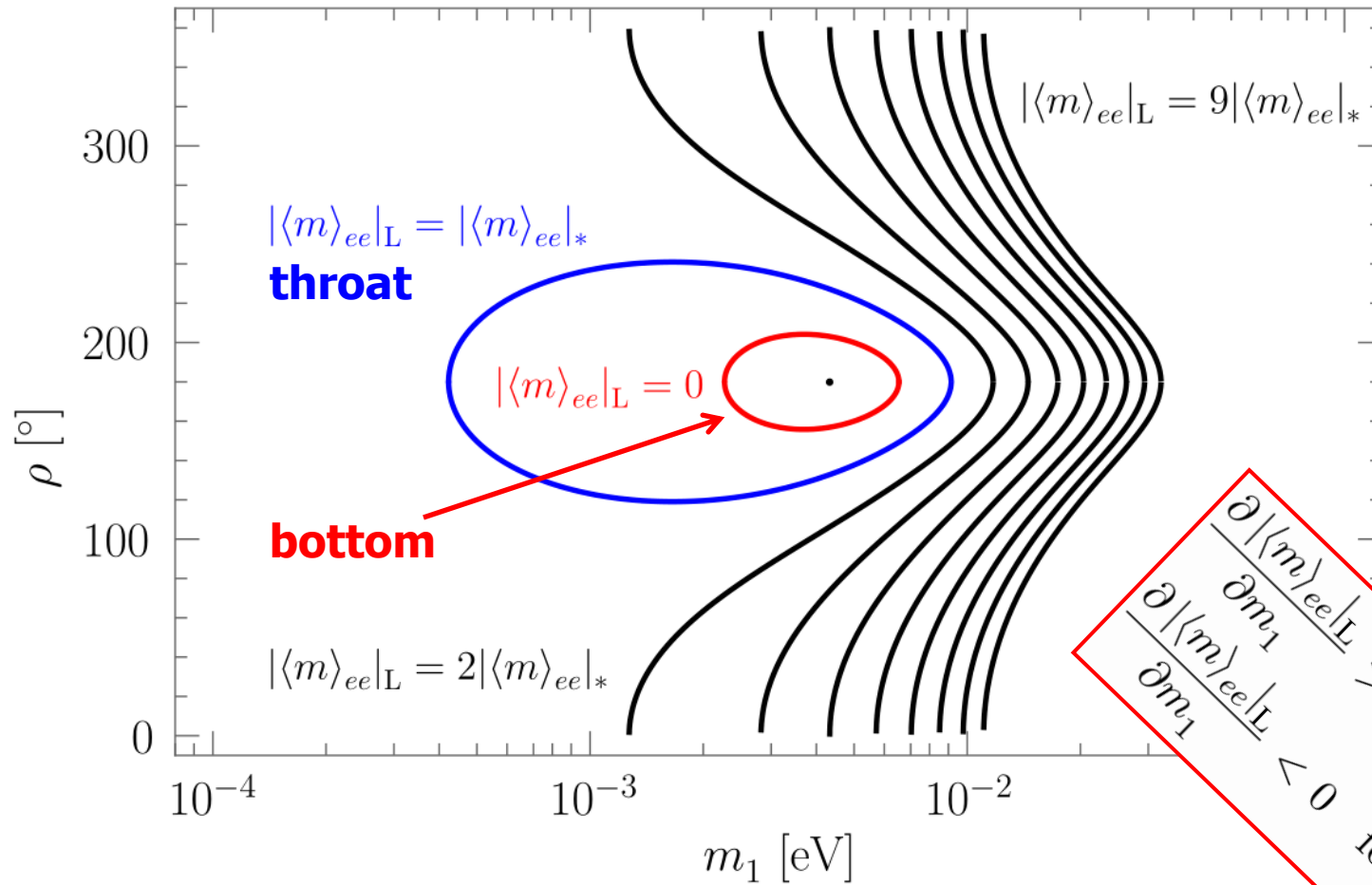


**= 0**

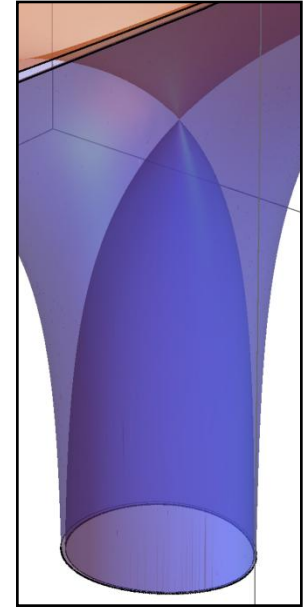
$$|\langle m \rangle_{ee}|_* = m_3 \sin^2 \theta_{13}$$

**~ 1 meV**

# The throat



## EXTREMUM



$$\frac{\partial |\langle m \rangle_{ee}|_L}{\partial m_1} > 0 \text{ for } m_1 < m_2 \tan^2 \theta_{12}$$

$$\frac{\partial |\langle m \rangle_{ee}|_L}{\partial m_1} < 0 \text{ for } m_1 > m_2 \tan^2 \theta_{12}$$

$$\frac{\partial |\langle m \rangle_{ee}|_L}{\partial \rho} = \frac{m_1 m_2 \sin^2 2\theta_{12} \cos^2 \theta_{13}}{4\bar{m}_{12}} \sin \rho = 0$$

$$\frac{\partial |\langle m \rangle_{ee}|_L}{\partial m_1} = \frac{m_1}{m_3} \sin^2 \theta_{13} \pm \left( \cos^2 \theta_{12} - \frac{m_1}{m_2} \sin^2 \theta_{12} \right) \cos^2 \theta_{13} \neq 0$$

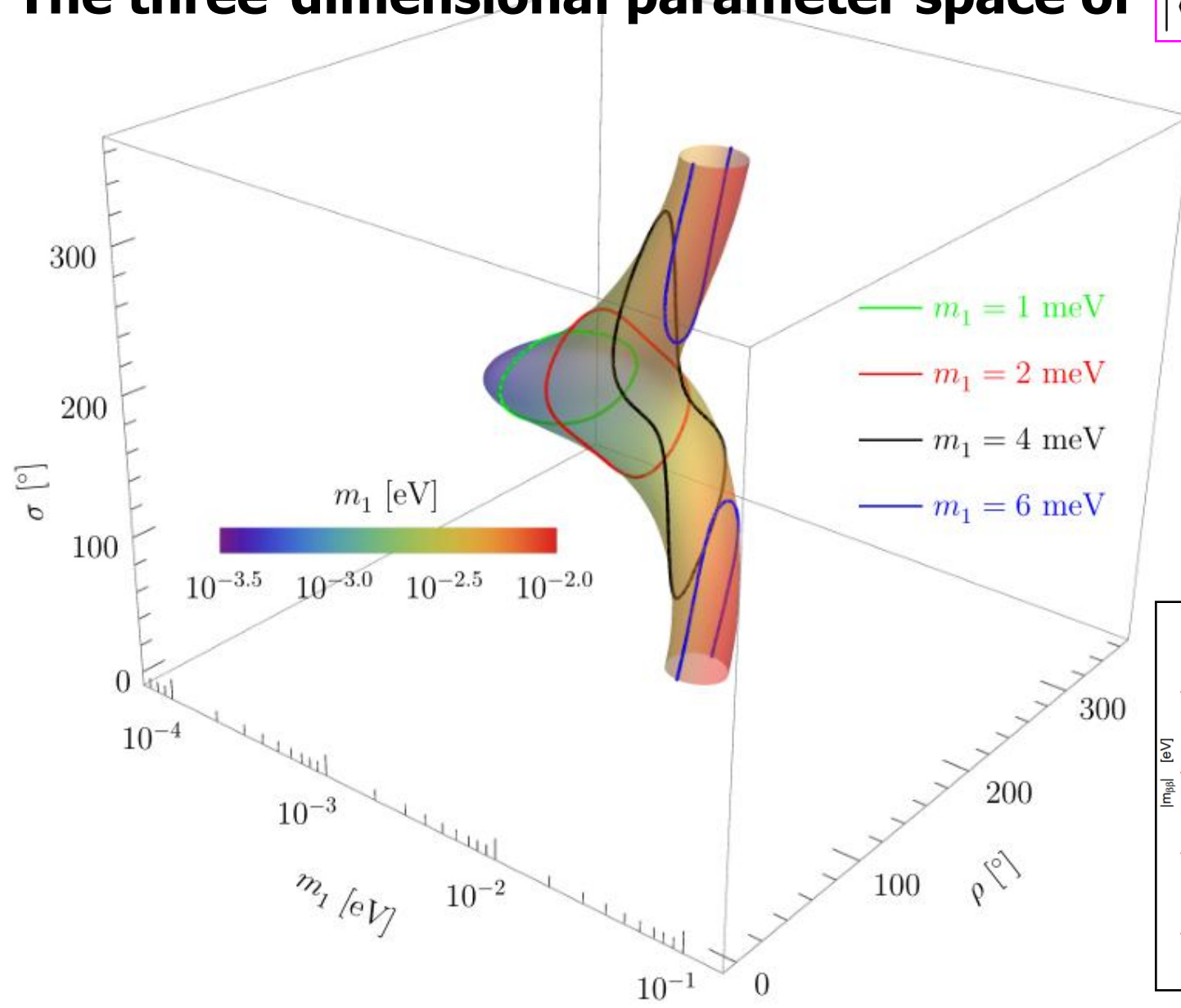


# To fall into the well

46

The three-dimensional parameter space of

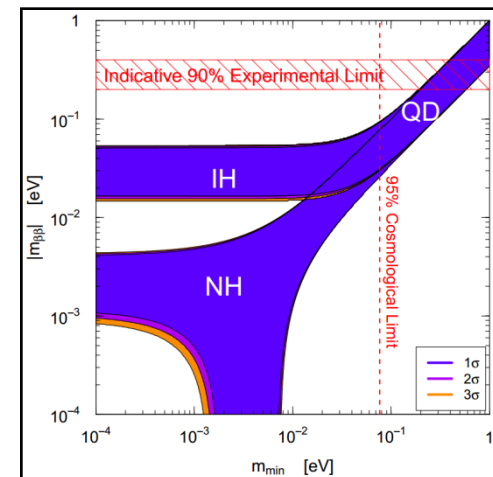
$$|\langle m \rangle_{ee}| < |\langle m \rangle_{ee}|_*$$



**Take it easy!**

**It is difficult  
to fall into  
the well!**

**Vissani Graph**



# Model building?

47

Why the relationship  $\tan \theta_{12} = \sqrt{m_1/m_2}$  is reasonable?

Remember  $\tan \theta_C \simeq \sqrt{m_d/m_s}$  in the quark sector as done by



S. Weinberg

H. Fritzsch

F. Wilczek + A. Zee

1977

The effective Majorana neutrino mass matrix (Xing, Zhao, 1612.08538)

$$M_\nu = \begin{pmatrix} 0 & A & A \\ A & B & C \\ A & C & B \end{pmatrix} - m_3 \frac{\sin \theta_{13}}{\sqrt{2}} \begin{pmatrix} \sqrt{2} \sin \theta_{13} & +i & -i \\ +i & 0 & 0 \\ -i & 0 & 0 \end{pmatrix}$$

$$M_\nu = \begin{pmatrix} C & D & D^* \\ D & A & B \\ D^* & B & A^* \end{pmatrix}$$

$\nu_e \quad \nu_\mu \leftrightarrow \nu_\tau^c$

Predictions, thanks to the  $\mu$ - $\tau$  reflection symmetry:

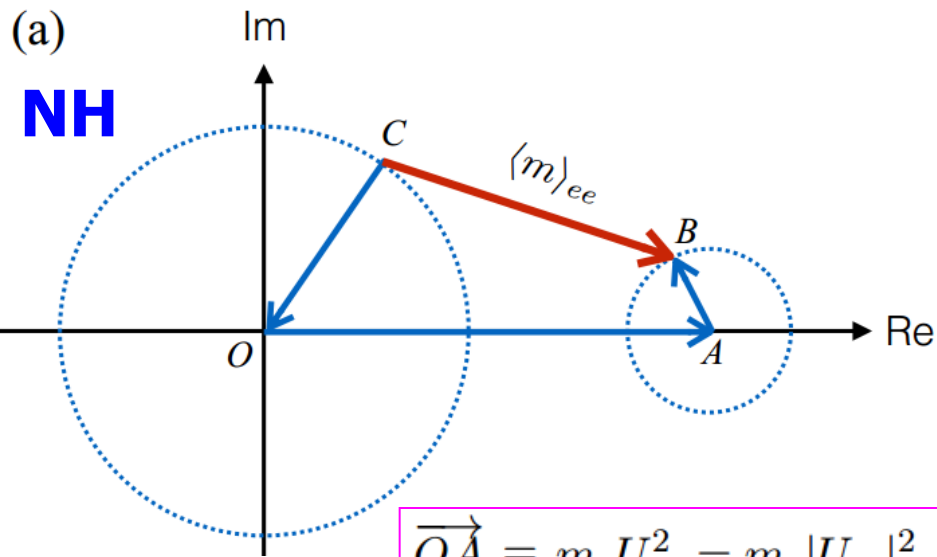
$$|\langle m \rangle_{ee}| = m_3 \sin^2 \theta_{13} \text{ and } \tan \theta_{12} = \sqrt{m_1/m_2}$$

$$\theta_{23} = \pi/4, \delta = -\pi/2, \rho = \pi \text{ and } \sigma = 0$$

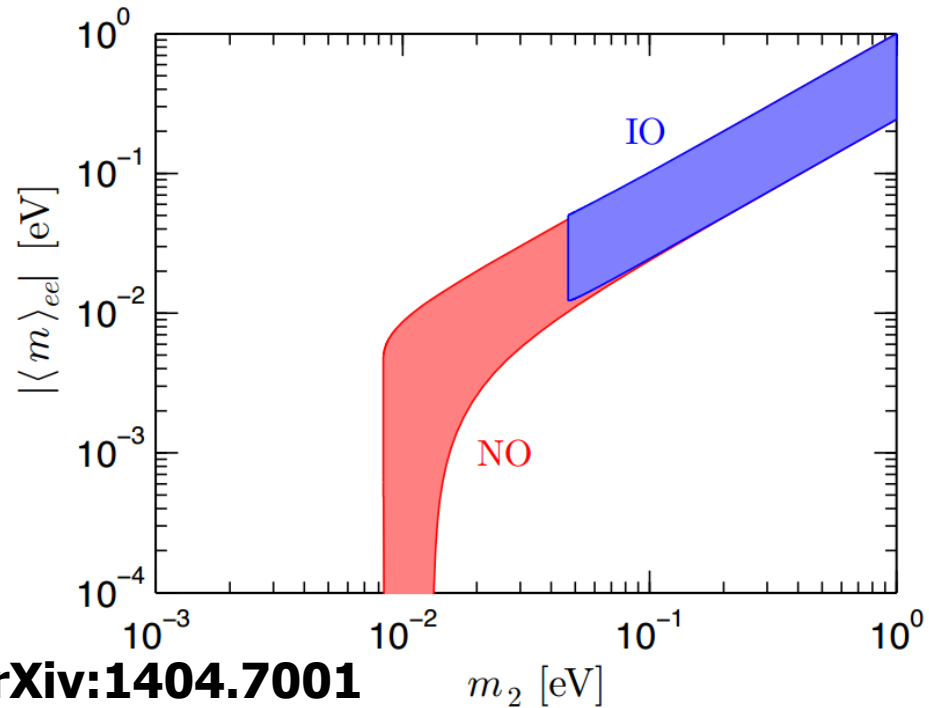
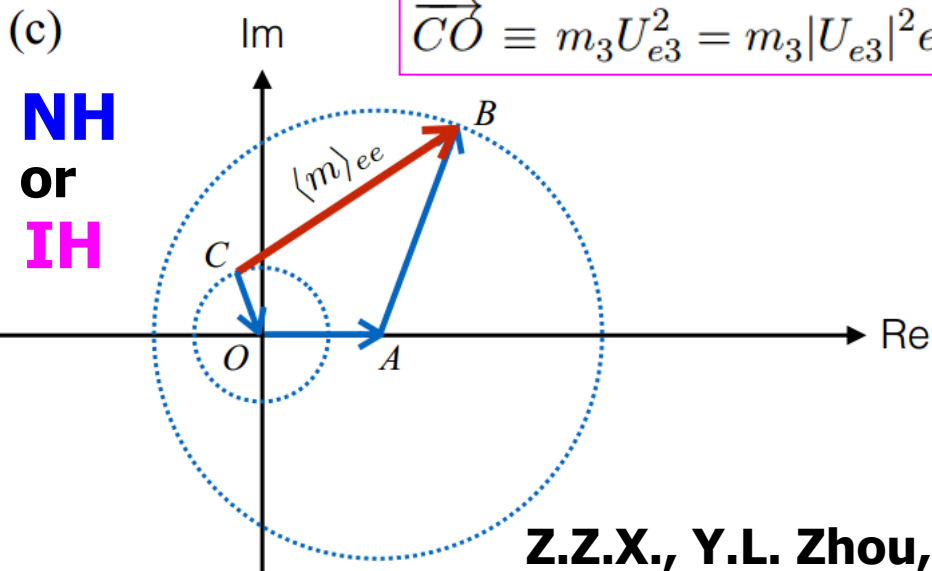
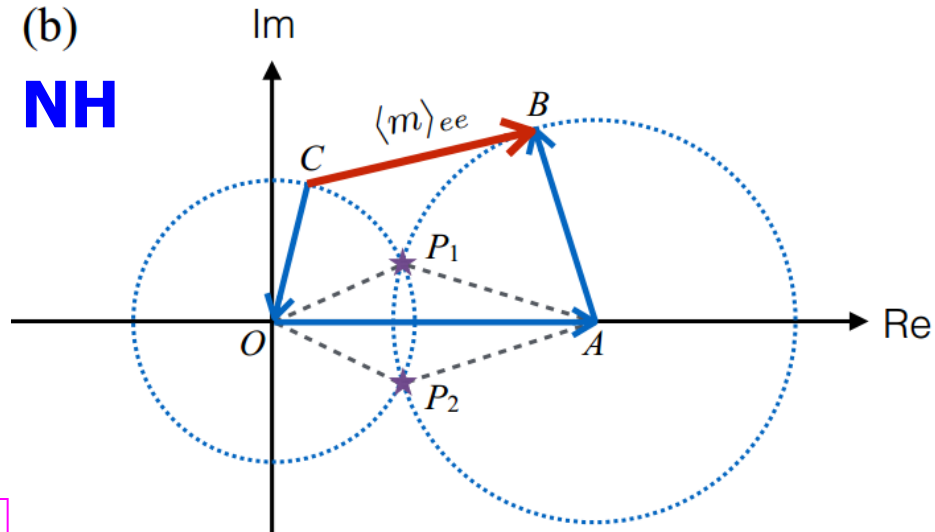
consistent with  
current data!



# Coupling-rod diagram

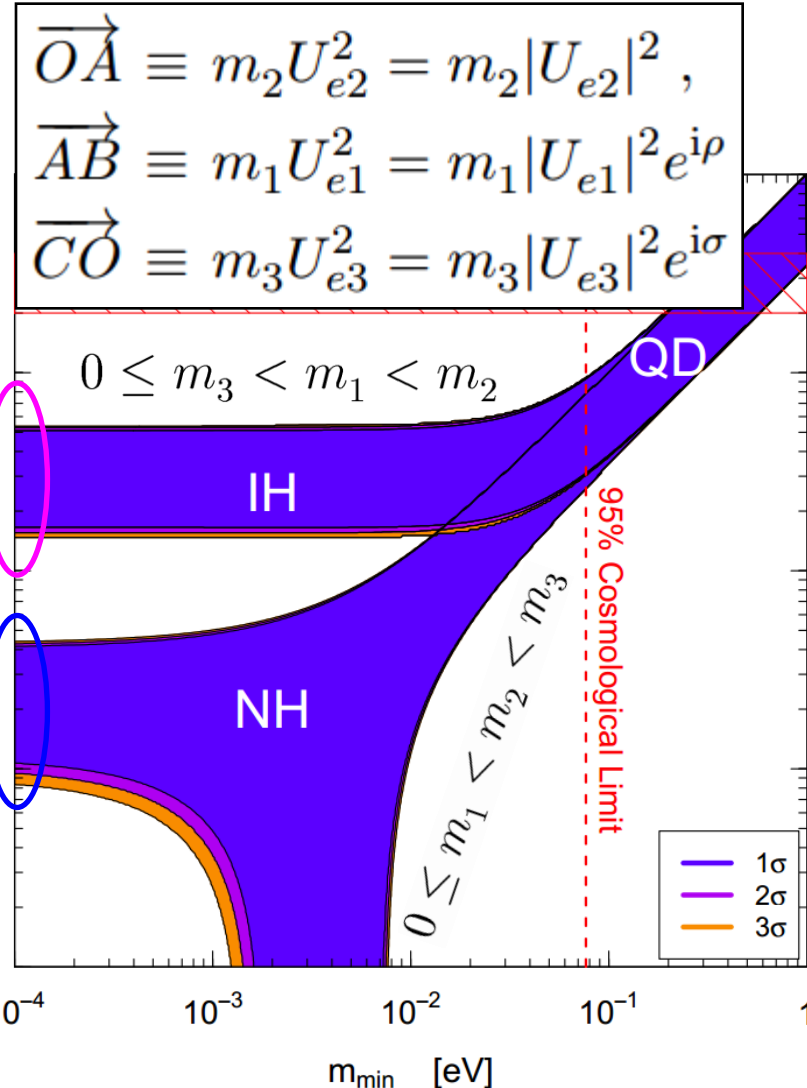
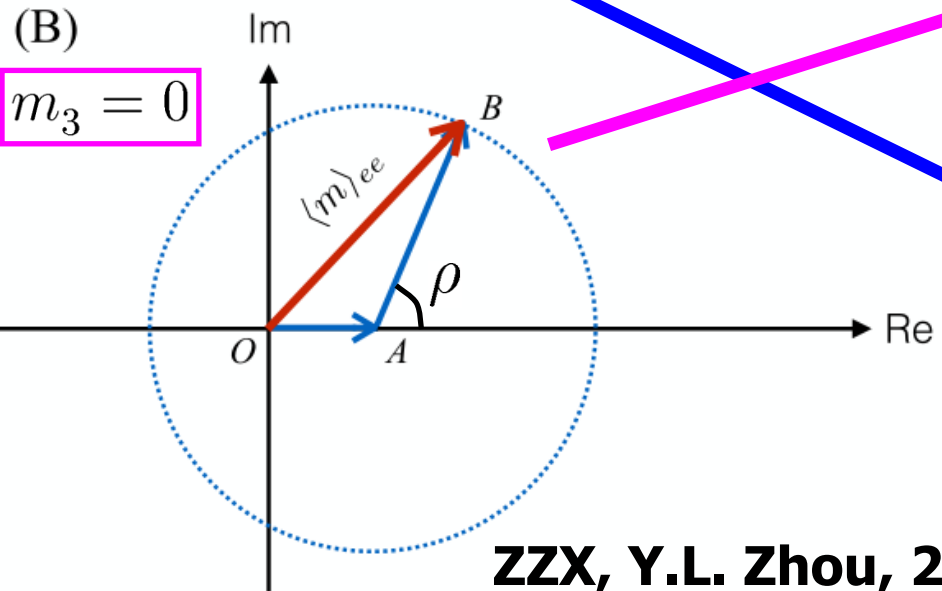
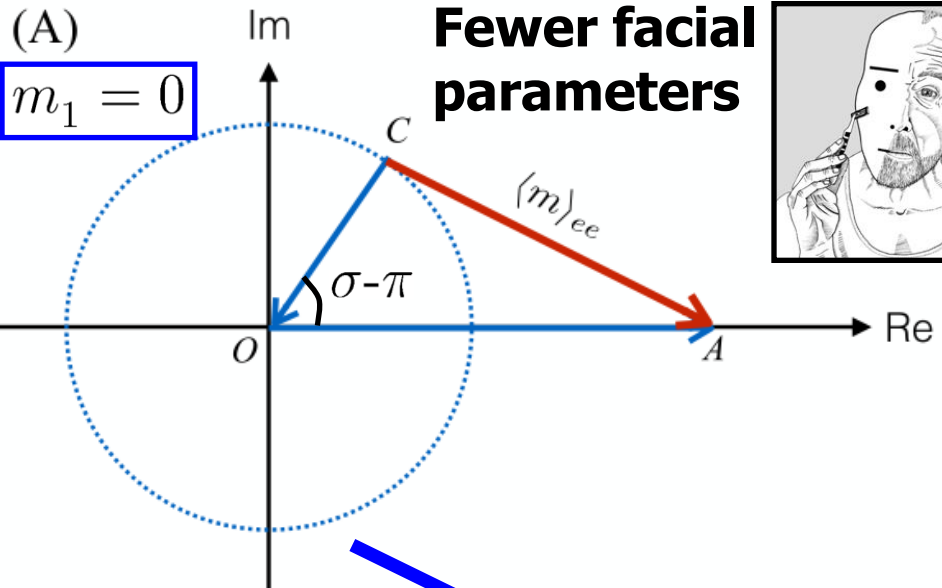


$$\begin{aligned} \vec{OA} &\equiv m_2 U_{e2}^2 = m_2 |U_{e2}|^2, \\ \vec{AB} &\equiv m_1 U_{e1}^2 = m_1 |U_{e1}|^2 e^{i\rho} \\ \vec{CO} &\equiv m_3 U_{e3}^2 = m_3 |U_{e3}|^2 e^{i\sigma} \end{aligned}$$



# Occam's razor: $0\nu 2\beta$

**Entities must not be multiplied beyond necessity.**

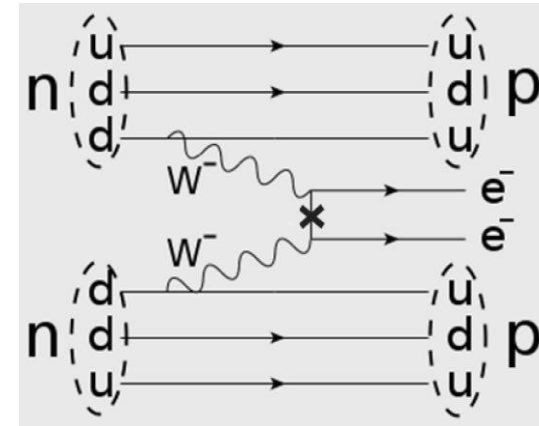


**Type (A): NP directly related to extra species of neutrinos.**

**Example 1: heavy Majorana neutrinos from type-I seesaw**

$$-\mathcal{L}_{\text{lepton}} = \bar{l}_L Y_l H E_R + \bar{l}_L Y_\nu \tilde{H} N_R + \frac{1}{2} \overline{N_R^c} M_R N_R + \text{h.c.}$$

$$\Gamma_{0\nu\beta\beta} \propto \left| \sum_{i=1}^3 m_i U_{ei}^2 - \sum_{k=1}^n \frac{R_{ek}^2}{M_k} M_A^2 \mathcal{F}(A, M_k) \right|^2$$



**In most cases the heavy contribution is negligible**

**Example 2: light sterile neutrinos from LSND etc**

$$\langle m \rangle'_{ee} \equiv \sum_{i=1}^6 m_i U_{ei}^2 = \underline{\langle m \rangle_{ee}} (c_{14} c_{15} c_{16})^2 + \underline{m_4 (\hat{s}_{14}^* c_{15} c_{16})^2} + m_5 (\hat{s}_{15}^* c_{16})^2 + m_6 (\hat{s}_{16}^*)^2$$

**In this case the new contribution might be constructive or destructive**

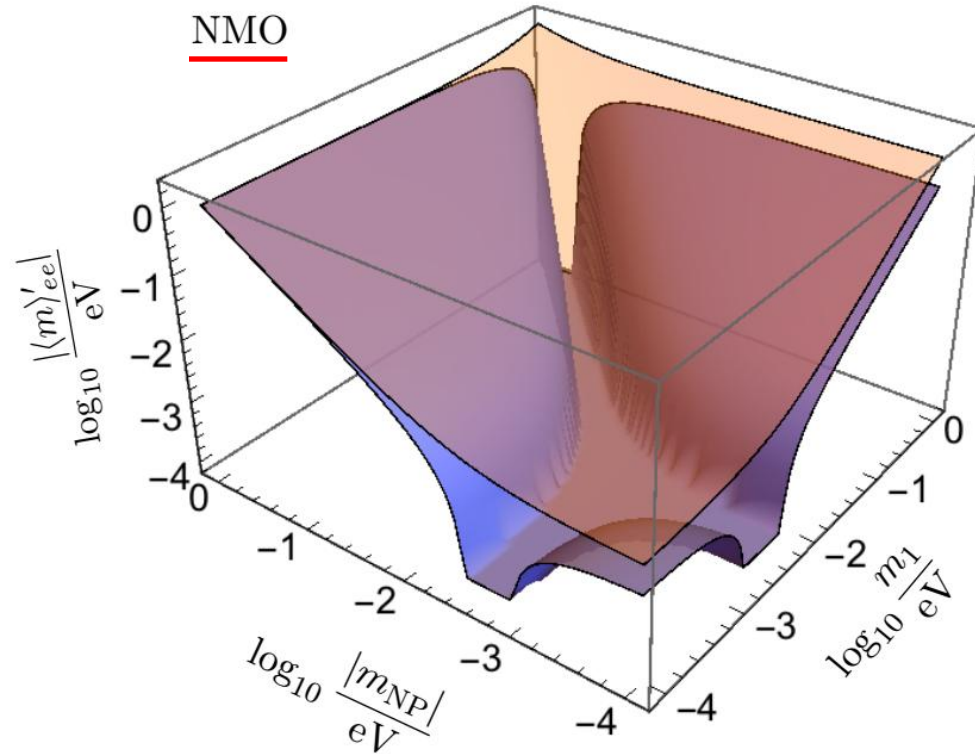
**Type (B): NP has little to do with the neutrino mass issue.**

**SUSY, Left-right, and some others that I don't understand**

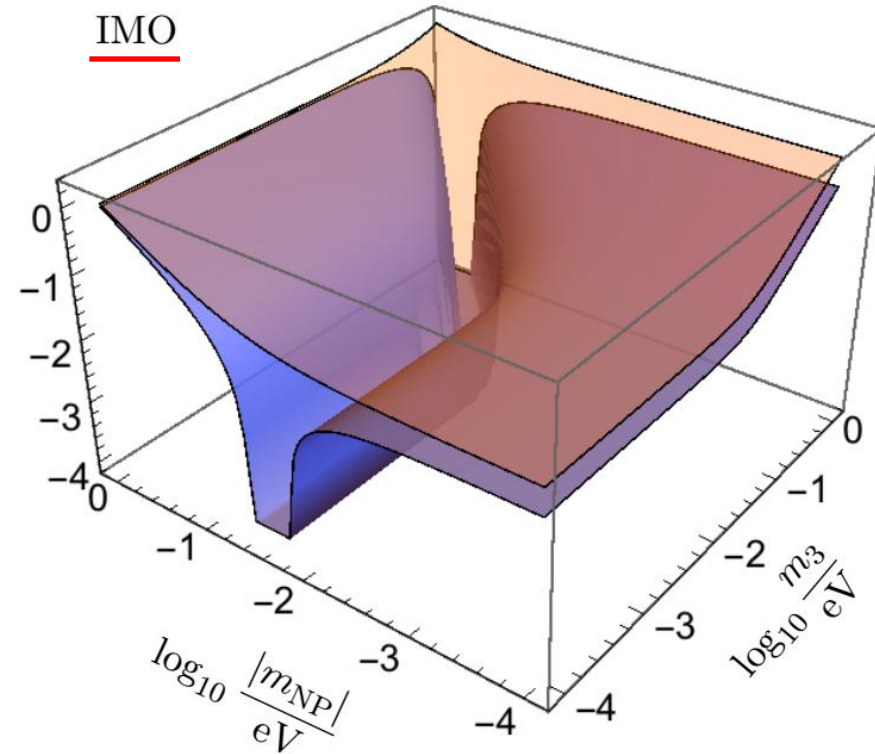
## New physics effects:

$$\langle m \rangle'_{ee} = m_1 U_{e1}^2 + m_2 U_{e2}^2 + m_3 U_{e3}^2 + m_{\text{NP}}$$

NMO



IMO



**Lower bound: blue; upper bound: light orange.** Clearer sensitivities to mass and phase parameters (Xing, Zhao, Zhou, arXiv:1504.05820)

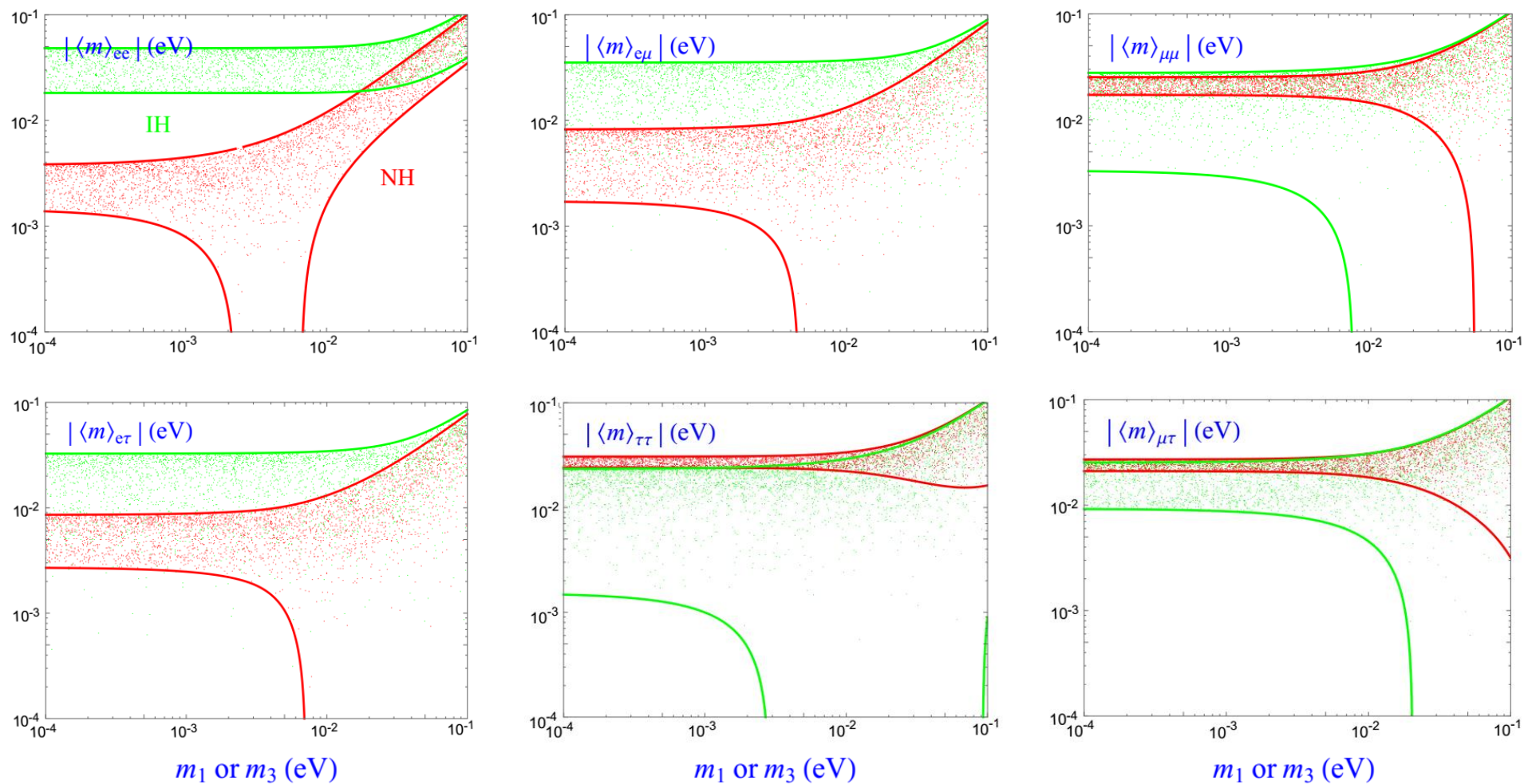
$$|\langle m \rangle'_{ee}|_{\text{upper}} = m_1 |U_{e1}|^2 + m_2 |U_{e2}|^2 + m_3 |U_{e3}|^2 + |m_{\text{NP}}|,$$

$$|\langle m \rangle'_{ee}|_{\text{lower}} = \max \left\{ 0, 2m_i |U_{ei}|^2 - |\langle m \rangle'_{ee}|_{\text{upper}}, 2|m_{\text{NP}}| - |\langle m \rangle'_{ee}|_{\text{upper}} \right\}$$

**It is hard to tell much**

# More LNV processes

52



To identify the Majorana nature, CP-violating phases and new physics it is imperative to observe the  $0\nu 2\beta$  decays and other lepton-number-violating processes (e.g., neutrino-antineutrino oscillations, the relic neutrino background, doubly-charged Higgs decays). **None is realistic**

# Lecture A6

- ★ **How to Generate Neutrino Mass**
- ★ **3 Typical Seesaw Mechanisms**
- ★ **Active-sterile neutrino mixing**

# Hybrid mass term (1)

A **hybrid** mass term can be written out in terms of the left- and right-handed neutrino fields and their charge-conjugate counterparts:

$$\begin{aligned}
 -\mathcal{L}'_{\text{hybrid}} &= \overline{\nu}_L M_D N_R + \frac{1}{2} \overline{\nu}_L M_L (\nu_L)^c + \frac{1}{2} \overline{(N_R)^c} M_R N_R + \text{h.c.} \\
 &= \frac{1}{2} \begin{bmatrix} \overline{\nu}_L & \overline{(N_R)^c} \end{bmatrix} \begin{pmatrix} M_L & M_D \\ M_D^T & M_R \end{pmatrix} \begin{bmatrix} (\nu_L)^c \\ N_R \end{bmatrix} + \text{h.c.},
 \end{aligned}$$

← **type-(I+II) seesaw**

Here we have used

**Diagonalization** by means of a **6×6 unitary** matrix:

$$\overline{(N_R)^c} M_D^T (\nu_L)^c = [(\nu_L)^T \mathcal{C} M_D^T \mathcal{C} \overline{\nu}_L^T]^T = \overline{\nu}_L M_D N_R$$

$$\begin{pmatrix} V & R \\ S & U \end{pmatrix}^\dagger \begin{pmatrix} M_L & M_D \\ M_D^T & M_R \end{pmatrix} \begin{pmatrix} V & R \\ S & U \end{pmatrix}^* = \begin{pmatrix} \widehat{M}_\nu & \mathbf{0} \\ \mathbf{0} & \widehat{M}_N \end{pmatrix}$$

$$\widehat{M}_\nu \equiv \text{Diag}\{m_1, m_2, m_3\}, \quad \widehat{M}_N \equiv \text{Diag}\{M_1, M_2, M_3\}$$

**Majorana mass states**

$$\nu' = \begin{bmatrix} \nu'_L \\ (N'_R)^c \end{bmatrix} + \begin{bmatrix} (\nu'_L)^c \\ N'_R \end{bmatrix} = \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \\ N_1 \\ N_2 \\ N_3 \end{pmatrix}$$

$$(\nu')^c = \nu'$$

It is actually a **Majorana** mass term!

$$-\mathcal{L}'_{\text{hybrid}} = \frac{1}{2} \begin{bmatrix} \overline{\nu}'_L & \overline{(N'_R)^c} \end{bmatrix} \begin{pmatrix} \widehat{M}_\nu & \mathbf{0} \\ \mathbf{0} & \widehat{M}_N \end{pmatrix} \begin{bmatrix} (\nu'_L)^c \\ N'_R \end{bmatrix} + \text{h.c.}$$

$$\nu'_L = V^\dagger \nu_L + S^\dagger (N_R)^c$$

$$N'_R = R^T (\nu_L)^c + U^T N_R$$



# Hybrid mass term (2)

**Physical mass term:**

$$-\mathcal{L}'_{\text{hybrid}} = \frac{1}{2} \bar{\nu}' \begin{pmatrix} \widehat{M}_\nu & \mathbf{0} \\ \mathbf{0} & \widehat{M}_N \end{pmatrix} \nu' = \frac{1}{2} \sum_{i=1}^3 (m_i \bar{\nu}'_i \nu'_i + M_i \bar{N}'_i N'_i)$$

**Kinetic term:**

$$\begin{aligned} \mathcal{L}_{\text{kinetic}} &= i \bar{\nu}'_L \gamma_\mu \partial^\mu \nu'_L + i \bar{N}'_R \gamma_\mu \partial^\mu N'_R \\ &= \frac{i}{2} \left[ \bar{\nu}'_L \quad \overline{(N'_R)^c} \right] \gamma_\mu \partial^\mu \begin{bmatrix} \nu'_L \\ (N'_R)^c \end{bmatrix} + \frac{i}{2} \left[ (\nu'_L)^c \quad \bar{N}'_R \right] \gamma_\mu \partial^\mu \begin{bmatrix} (\nu'_L)^c \\ N'_R \end{bmatrix} \\ &= \frac{i}{2} \left[ \bar{\nu}'_L \quad \overline{(N'_R)^c} \right] \gamma_\mu \partial^\mu \begin{pmatrix} V & R \\ S & U \end{pmatrix}^\dagger \begin{pmatrix} V & R \\ S & U \end{pmatrix} \begin{bmatrix} \nu'_L \\ (N'_R)^c \end{bmatrix} \\ &\quad + \frac{i}{2} \left[ (\nu'_L)^c \quad \bar{N}'_R \right] \gamma_\mu \partial^\mu \begin{pmatrix} V & R \\ S & U \end{pmatrix}^T \begin{pmatrix} V & R \\ S & U \end{pmatrix}^* \begin{bmatrix} (\nu'_L)^c \\ N'_R \end{bmatrix} \\ &= \frac{i}{2} \left[ \bar{\nu}'_L \quad \overline{(N'_R)^c} \right] \gamma_\mu \partial^\mu \begin{bmatrix} \nu'_L \\ (N'_R)^c \end{bmatrix} + \frac{i}{2} \left[ (\nu'_L)^c \quad \bar{N}'_R \right] \gamma_\mu \partial^\mu \begin{bmatrix} (\nu'_L)^c \\ N'_R \end{bmatrix} \\ &= i \bar{\nu}'_L \gamma_\mu \partial^\mu \nu'_L + i \bar{N}'_R \gamma_\mu \partial^\mu N'_R \\ &= \frac{i}{2} \bar{\nu}' \gamma_\mu \partial^\mu \nu' = \frac{i}{2} \sum_{k=1}^3 (\bar{\nu}'_k \gamma_\mu \partial^\mu \nu'_k + \bar{N}'_k \gamma_\mu \partial^\mu N'_k) \end{aligned}$$

# Non-unitary flavor mixing

56

Weak charged-current interactions of leptons:

In the flavor basis

$$\mathcal{L}_{\text{cc}} = \frac{g}{\sqrt{2}} \overline{(e \ \mu \ \tau)}_{\text{L}} \gamma^{\mu} \begin{pmatrix} \nu_e \\ \nu_{\mu} \\ \nu_{\tau} \end{pmatrix}_{\text{L}} W_{\mu}^{-} + \text{h.c.}$$

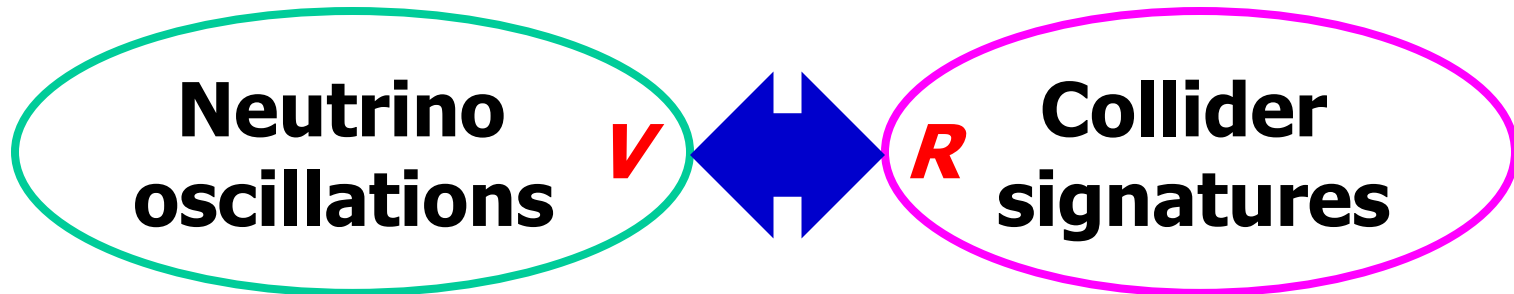
$$\nu_{\text{L}} = V \nu'_{\text{L}} + R (N'_{\text{R}})^c$$

In the mass basis

$$\mathcal{L}_{\text{cc}} = \frac{g}{\sqrt{2}} \overline{(e \ \mu \ \tau)}_{\text{L}} \gamma^{\mu} \left[ V \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}_{\text{L}} + R \begin{pmatrix} N_1 \\ N_2 \\ N_3 \end{pmatrix}_{\text{L}} \right] W_{\mu}^{-} + \text{h.c.}$$

$V$  = non-unitary light neutrino mixing (PMNS) matrix  $VV^{\dagger} + RR^{\dagger} = 1$

$R$  = light-heavy neutrino mixing (CC interactions of heavy neutrinos)



TeV seesaws may bridge the gap between neutrino & collider physics.

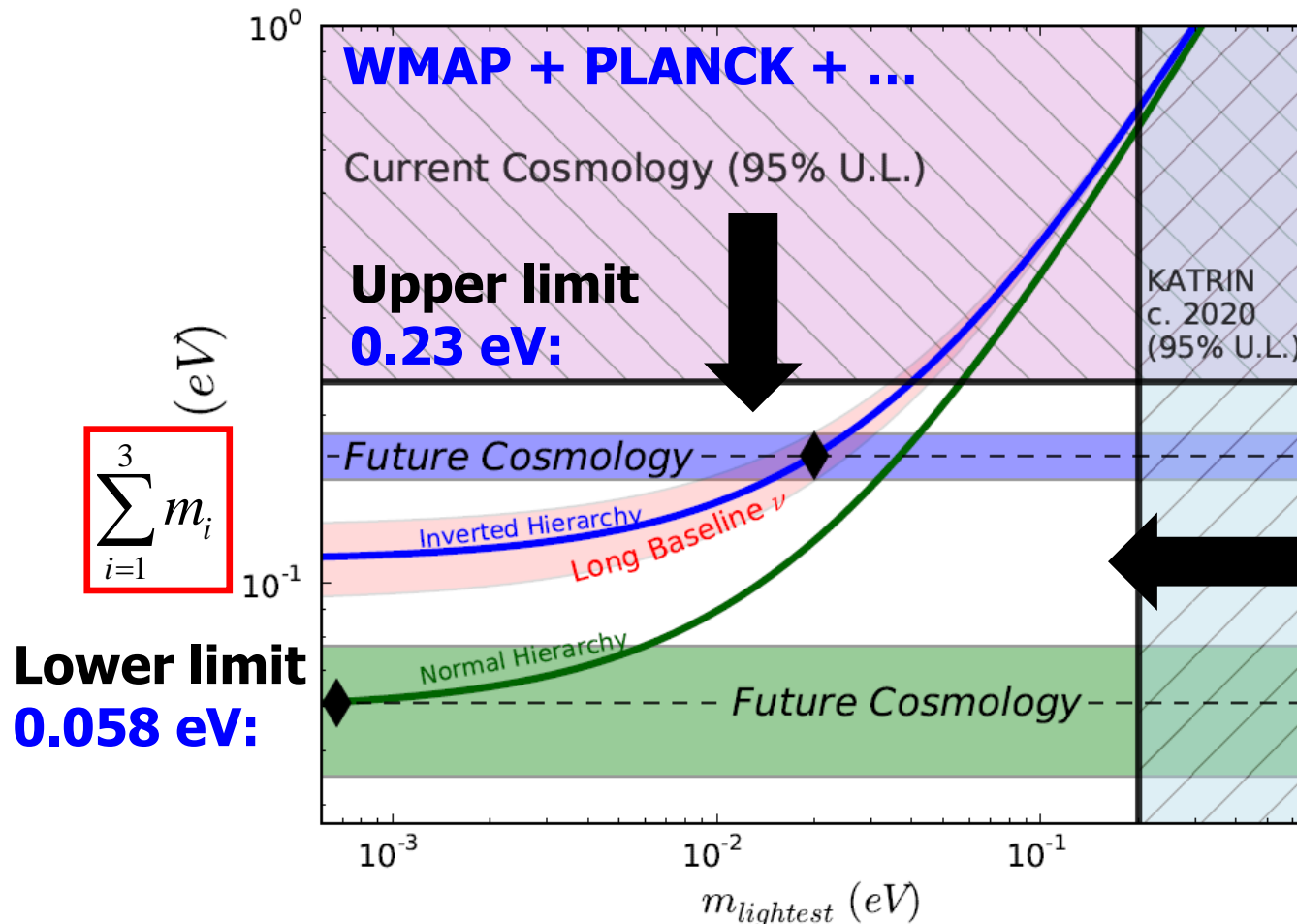
# Neutrino mass scale

Three ways: the  $\beta$  decay, the  $0\nu\beta\beta$  decay, and cosmology (CMB + LSS).

$$\langle m \rangle_e^2 = \sum_{i=1}^3 m_i^2 |U_{ei}|^2$$

$$|\langle m \rangle_{ee}| = \left| \sum_{i=1}^3 m_i U_{ei}^2 \right|$$

$$\sum_{i=1}^3 m_i$$



mass scale  
 $\leq 0(0.1) \text{ eV}$

Why so tiny?

arXiv:1309.5383

Stage-4 CMB

$$\sigma \left( \sum m_\nu \right) = 16 \text{ meV}$$

$$\sigma (N_{\text{eff}}) = 0.020 .$$

A **hybrid** mass term may have three distinct components:

$$\begin{aligned} -\mathcal{L}'_{\text{hybrid}} &= \overline{\nu}_L M_D N_R + \frac{1}{2} \overline{\nu}_L M_L (\nu_L)^c + \frac{1}{2} \overline{(N_R)^c} M_R N_R + \text{h.c.} \\ &= \frac{1}{2} \begin{bmatrix} \overline{\nu}_L & \overline{(N_R)^c} \end{bmatrix} \begin{pmatrix} M_L & M_D \\ M_D^T & M_R \end{pmatrix} \begin{bmatrix} (\nu_L)^c \\ N_R \end{bmatrix} + \text{h.c.}, \end{aligned}$$

- ♣ **Normal Dirac mass term**, proportional to the scale of electroweak symmetry breaking ( $\sim$  **174 GeV**);
- ♣ **Light Majorana mass term**, violating the SM gauge symmetry and much lower than **174 GeV** ('t Hooft's naturalness criterion);
- ♣ **Heavy Majorana mass term**, originating from the SU(2)<sub>L</sub> singlet and having a scale much higher than **174 GeV**.

A strong hierarchy of **3** mass scales allows us to make approximation

$$\begin{pmatrix} V & R \\ S & U \end{pmatrix}^\dagger \begin{pmatrix} M_L & M_D \\ M_D^T & M_R \end{pmatrix} \begin{pmatrix} V & R \\ S & U \end{pmatrix}^* = \begin{pmatrix} \widehat{M}_\nu & \mathbf{0} \\ \mathbf{0} & \widehat{M}_N \end{pmatrix}$$

# Seesaw mechanisms (2)

59

The above **unitary** transformation leads to the following relationships:

$$\begin{aligned} R\widehat{M}_N &= M_L R^* + M_D U^* \\ S\widehat{M}_\nu &= M_D^T V^* + M_R S^* \end{aligned}$$

$$M_R \gg M_D \gg M_L$$

$$R \sim S \sim \mathcal{O}(M_D/M_R)$$

$$\begin{aligned} U\widehat{M}_N &= M_R U^* + M_D^T R^* \\ V\widehat{M}_\nu &= M_L V^* + M_D S^* \end{aligned}$$

$$\begin{aligned} U\widehat{M}_N U^T &= M_R (UU^\dagger)^T + M_D^T (R^* U^T) \approx M_R, \\ V\widehat{M}_\nu V^T &= M_L (VV^\dagger)^T + M_D (S^* V^T) \approx M_L + M_D (S^* V^T) \end{aligned}$$

$$S^* V^T = M_R^{-1} S\widehat{M}_\nu V^T - M_R^{-1} M_D^T (VV^\dagger)^T \approx -M_R^{-1} M_D^T$$

Then we arrive at the **type-(I+II)** seesaw formula:

$$M_\nu \equiv V\widehat{M}_\nu V^T \approx M_L - M_D M_R^{-1} M_D^T$$

**Type-I** seesaw limit:

$$M_\nu \approx -M_D M_R^{-1} M_D^T \quad (\text{Fritzsch, Gell-Mann, Minkowski, 1975; Minkowski, 1977; ...})$$

**Type-II** seesaw limit:

$$M_\nu = M_L \quad (\text{Konetschny, Kummer, 1977; ...})$$

# History of type-I seesaw

60

The **seesaw** idea **originally** appeared in a paper's **footnote**.



## **Seesaw**—**A Footnote Idea:**

H. Fritzsch, M. Gell-Mann,

P. Minkowski, PLB 59 (**1975**) 256

This idea was very clearly elaborated by **Minkowski** in Phys. Lett. B 67 (**1977**) 421 ---- but it was unjustly forgotten until **2004**.



The idea was later on embedded into the **GUT** frameworks in **1979** and **1980**:

- T. Yanagida **1979**
- M. Gell-Mann, P. Ramond, R. Slansky **1979**
- S. Glashow **1979**
- R. Mohapatra, G. Senjanovic **1980**

It was **Yanagida** who named this mechanism as "**seesaw**".

# What is History?

**History is a set of lies agreed upon**





# Summary of 3 seesaws

62

**Type-I seesaw: SM + right-handed neutrinos + L violation**  
(Minkowski 77; Yanagida 79; Glashow 79; Gell-Mann, Ramond, Slansky 79; Mohapatra, Senjanovic 80)

$$-\mathcal{L}_{\text{lepton}} = \bar{l}_L Y_l H E_R + \bar{l}_L Y_\nu \tilde{H} N_R + \frac{1}{2} \bar{N}_R^c M_R N_R + \text{h.c.}$$

**Type-II seesaw: SM + 1 Higgs triplet + L violation**  
(Konetschny, Kummer 77; Magg, Wetterich 80; Schechter, Valle 80; Cheng, Li 80; Lazarides et al 80)

$$-\mathcal{L}_{\text{lepton}} = \bar{l}_L Y_l H E_R + \frac{1}{2} \bar{l}_L Y_\Delta \Delta i\sigma_2 l_L^c - \lambda_\Delta M_\Delta H^T i\sigma_2 \Delta H + \text{h.c.}$$

**Type-III seesaw: SM + 3 triplet fermions + L violation**  
(Foot, Lew, He, Joshi 1989)

$$-\mathcal{L}_{\text{lepton}} = \bar{l}_L Y_l H E_R + \bar{l}_L \sqrt{2} Y_\Sigma \Sigma^c \tilde{H} + \frac{1}{2} \text{Tr} (\bar{\Sigma} M_\Sigma \Sigma^c) + \text{h.c.}$$

# Effective mass term

**Weinberg (1979):** the unique **dimension-five** operator of  **$\nu$ -masses** after integrating out heavy degrees of freedom.

$$\frac{\mathcal{L}_{d=5}}{\Lambda} = \begin{cases} \frac{1}{2} (Y_\nu M_R^{-1} Y_\nu^T)_{\alpha\beta} \bar{l}_{\alpha L} \tilde{H} \tilde{H}^T l_{\beta L}^c + \text{h.c.} & \text{(Type 1)} \\ -\frac{\lambda_\Delta}{M_\Delta} (Y_\Delta)_{\alpha\beta} \bar{l}_{\alpha L} \tilde{H} \tilde{H}^T l_{\beta L}^c + \text{h.c.} & \text{(Type 2)} \\ \frac{1}{2} (Y_\Sigma M_\Sigma^{-1} Y_\Sigma^T)_{\alpha\beta} \bar{l}_{\alpha L} \tilde{H} \tilde{H}^T l_{\beta L}^c + \text{h.c.} & \text{(Type 3)} \end{cases}$$

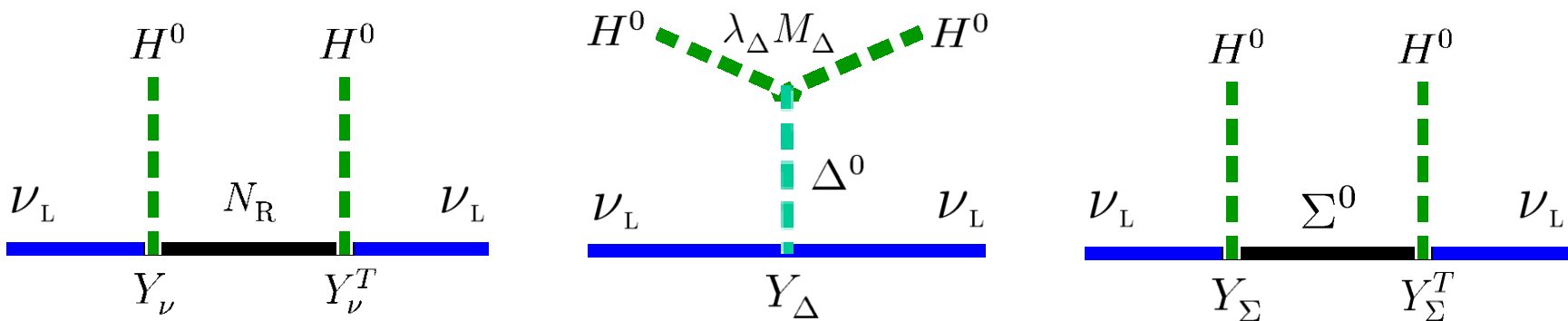
$$M_\nu = \begin{cases} -\frac{1}{2} Y_\nu \frac{v^2}{M_R} Y_\nu^T & \text{(Type 1)} \\ \lambda_\Delta Y_\Delta \frac{v^2}{M_\Delta} & \text{(Type 2)} \\ -\frac{1}{2} Y_\Sigma \frac{v^2}{M_\Sigma} Y_\Sigma^T & \text{(Type 3)} \end{cases}$$



**After SSB, a Majorana mass term is**

$$-\mathcal{L}_{\text{mass}} = \frac{1}{2} \bar{\nu}_L M_\nu \nu_L^c + \text{h.c.}$$

$$\langle \tilde{H} \rangle = v/\sqrt{2}$$

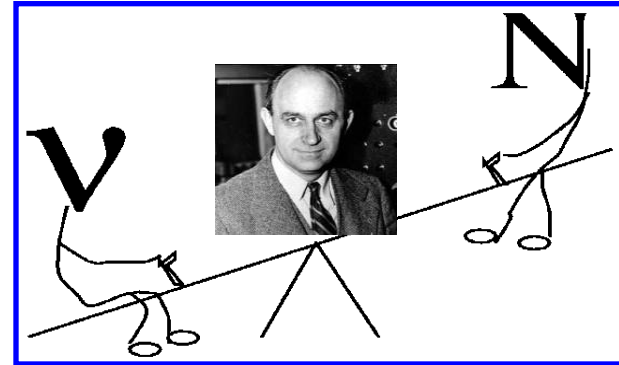


# Seesaw scale?

What is the scale at which the **seesaw** mechanism works?



← **Planck**



← **GUT**

to unify strong, weak & electromagnetic forces

**Conventional Seesaws: heavy degrees of freedom near **GUT****

This appears to be rather reasonable, since one often expects **new physics** to appear around a fundamental scale

← **Fermi**

**Naturalness** ✓

**Testability** ✗

**Uniqueness** ✗

**Hierarchy** ✗

# TeV Neutrino Physics?

to discover the SM Higgs boson



to verify Yukawa interactions



Why

to pin down heavy seesaw particles

to test seesaw mechanism(s)

Not

to measure low-energy effects

Try



LHC

TeV

# Real + Hypothetical $\nu$ 's

**sub-eV**

active  
neutrinos

**sub-eV**

sterile  
neutrinos

**keV**

sterile  
neutrinos

**TeV**

Majorana  
neutrinos

**$\geq$  EeV**

Majorana  
neutrinos

LSND + MiniBooNE + reactor  
anomalies CMB + BBN hints

LHC  
motivated

standard  
weak  
interaction

oscillation

cosmic  
messenger

warm  
dark  
matter

classical seesaws + GUTs

# (3+3) flavor mixing

**active  
flavor**

 $\nu_e$  $\nu_\mu$  $\nu_\tau$ 

**sterile  
flavor**

 $\nu_x$  $\nu_y$  $\nu_z$  $= \mathcal{U}$  $\nu_1$  $\nu_2$  $\nu_3$  $\nu_4$  $\nu_5$  $\nu_6$ 

**mass  
state**

# A full parametrization

$$\mathcal{U} = \underbrace{\begin{pmatrix} 1 & \mathbf{0} \\ \mathbf{0} & U_0 \end{pmatrix}}_{\text{sterile part}} \underbrace{\begin{pmatrix} A & R \\ S & B \end{pmatrix}}_{\text{interplay}} \underbrace{\begin{pmatrix} V_0 & \mathbf{0} \\ \mathbf{0} & 1 \end{pmatrix}}_{\text{active part}}$$

$$\begin{pmatrix} V_0 & \mathbf{0} \\ \mathbf{0} & 1 \end{pmatrix} = O_{23} O_{13} O_{12} ,$$

**Full parametrization:**

**15** rotation angles

$$\begin{pmatrix} 1 & \mathbf{0} \\ \mathbf{0} & U_0 \end{pmatrix} = O_{56} O_{46} O_{45} ,$$

**15** phase phases

Xing, arXiv:1110.0083

$$\begin{pmatrix} A & R \\ S & B \end{pmatrix} = O_{36} O_{26} O_{16} O_{35} O_{25} O_{15} O_{34} O_{24} O_{14}$$



# Questions

69

- 1) Do you feel **happy** / **painful** / **sorry** to introduce sterile neutrinos into the SM (remember Weinberg's theorem)?
- 2) How many species of sterile neutrinos should be taken into account for your this or that purpose? **1?** **2?** **3?** ....?
- 3) If all the current experimental and observational hints disappear, will the **sterile neutrino physics** still survive?
- 4) Do you like to throw many stones to only kill few birds or just the opposite? **And is this a very stupid question?**

**Weinberg's 3rd law of progress in theoretical physics (83):**

You may use any degrees of freedom you like to describe a physical system, but if you use the wrong ones, you will be sorry ..... **What could be better?**

