

# Dark Matter, Dark Energy & Neutrino Mass

## 暗物质，暗能量和中微子质量

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理论物理前沿暑期讲习班——暗物质，中微子与粒子物理前沿

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中山大學  
SUN YAT-SEN UNIVERSITY

**Lecture 1: Introduction to Particle Physics and Cosmology**

**Lecture 2: Some Basic Backgrounds of the Standard Model of Particle Physics and Cosmology**

**Lecture 3: Neutrino Mass Generation**

**Lecture 4: Theoretical Understanding of Dark Matter Detections**

**Lecture 5: Dark Energy and Gravitational Waves**

# Lecture 3: Neutrino Mass Generation

## Outline

- Introduction
- A brief overview of neutrino mass generation
- **A special class of models to generate  $M_\nu$** 
  - Neutrino mass generation
  - **$0\nu\beta\beta$  decays**
  - **Other physics**

# ● Introduction

Weak eigenstate  
( $\alpha = e, \mu, \tau$ )

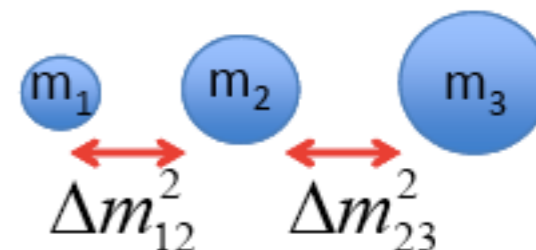
$$|\nu_\alpha\rangle = \sum_i U_{\alpha i} |\nu_i\rangle$$

Mass eigenstate  
( $i = 1, 2, 3$ )

## • PMNS (Pontecorvo-Maki-Nakagawa-Sakata) matrix

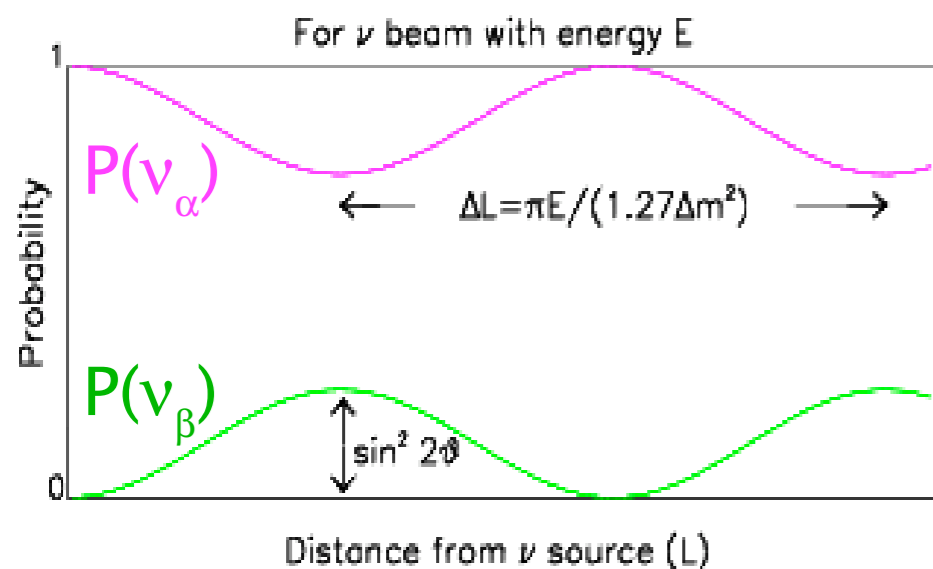
$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$U_{\text{PMNS}} = \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix}}_{\text{Atmospheric}} \underbrace{\begin{pmatrix} \cos\theta_{13} & 0 & e^{-i\delta}\sin\theta_{13} \\ 0 & 1 & 0 \\ -e^{i\delta}\sin\theta_{13} & 0 & \cos\theta_{13} \end{pmatrix}}_{\text{Reactor}} \underbrace{\begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\text{Solar}}$$



Two-neutrino case:

$$P(\nu_\alpha \rightarrow \nu_\beta) = |\langle \nu_\beta | \nu(t) \rangle|^2 = \sin^2 2\theta \sin^2 \left( \frac{1.27 \Delta m^2 [\text{eV}^2] L [\text{km}]}{E [\text{GeV}]} \right)$$



$$P(\nu_\alpha \rightarrow \nu_\alpha) = 1 - P(\nu_\alpha \rightarrow \nu_\beta)$$

disappearance of  $\nu_\alpha$

appearance of  $\nu_\beta$

$$P(\nu_\alpha \Rightarrow \nu_\beta) = \sin^2(2\theta) \sin^2 \left( \frac{\Delta m_{12}^2 L}{4E} \right)$$

$\theta$	: mixing angle
$\Delta m^2$	: mass squared difference
$L$ [km]	: the distance traveled
$E$ (GeV)	: the energy of neutrino

# 中微子振盪

如果中微子有質量，則不同類中微子間會產生振盪現象。

舉  $\nu_\mu$  及  $\nu_\tau$  為例：
$$\begin{pmatrix} \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

$\nu_\mu, \nu_\tau$  : 弱作用本徵態  
 $\nu_1, \nu_2$  : 質量本徵態

在  $t=0$  時， $\nu_\mu$  在大氣中產生，則

$$\begin{aligned} |\nu_\mu(0)\rangle &\equiv |\nu_\mu\rangle = \cos \theta |\nu_1(0)\rangle + \sin \theta |\nu_2(0)\rangle \\ |\nu_\tau\rangle &= -\sin \theta |\nu_1(0)\rangle + \cos \theta |\nu_2(0)\rangle \end{aligned}$$

到了時間  $t$ ，上述狀態演變為

$$|\nu_\mu(t)\rangle = \exp(-iE_1 t / \hbar) \cos \theta |\nu_1(0)\rangle + \exp(-iE_2 t / \hbar) \sin \theta |\nu_2(0)\rangle,$$

Using  $t \approx L$  and  $E - P_i = E - \sqrt{E^2 - m_i^2} \approx m_i^2 / 2E$ ,

如果  $m_1 \neq m_2$ ，則  $E_1 \neq E_2$ ，因此  $|\nu_\mu(t)\rangle$  不再垂直於  $|\nu_\tau\rangle$ !

$$\Delta m_{21}^2 = m_2^2 - m_1^2$$

$$P(\nu_\alpha \Rightarrow \nu_\beta) = \sin^2(2\theta) \sin^2\left(\frac{\Delta m_{12}^2 L}{4E}\right)$$



$$m_1 \neq m_2$$



# Neutrino Oscillations



Neutrinos have masses and mix with each other.

$$P(\nu_l \rightarrow \nu_{l'}) = \sum_j |U_{lj}|^2 |U_{l'j}|^2 + 2 \sum_{j>k} |U_{l'j} U_{lj}^* U_{lk} U_{l'k}^*|^2 \cos\left(\frac{\Delta m_{jk}^2}{2p} L - \varphi_{l'l;jk}\right)$$

$$\Delta m_{jk}^2 = m_j^2 - m_k^2$$

$$\varphi_{l'l;jk} = \arg(U_{l'j} U_{lj}^* U_{lk} U_{l'k}^*)$$

$\nu_{\text{atm}}$

SK UP-DOWN ASYMMETRY

$\theta_{23}$ -, L/E- dependences of  $\mu$ -like events

Dominant  $\nu_{\mu} \rightarrow \nu_{\tau}$  **K2K, MINOS, T2K; CNGS (OPERA)**  $\longrightarrow |\Delta m_{32}^2|, \sin^2\theta_{23}$

$\nu_{\odot}$

**Homestake, Kamiokande, SAGE, GALLEX/GNO  
Super-Kamiokande, SNO, BOREXINO; KamLAND**

$\longrightarrow \Delta m_{21}^2, \sin^2\theta_{12}$

$\bar{\nu}_e$

(from reactors): **Daya Bay, RENO, Double Chooz**

Dominant  $\nu_e \rightarrow \nu_{\mu,\tau}$   $\longrightarrow \sin^2\theta_{13}$

**T2K, MINOS** ( $\nu_{\mu}$  from accelerators):  $\nu_{\mu} \rightarrow \nu_e$

# Experiments on solar neutrinos

$$\Delta m_{21}^2 = \left(7.54_{-0.22}^{+0.26}\right) \times 10^{-5} \text{ eV}^2$$

$$\Delta m_{sun}^2 = |m_2|^2 - |m_1|^2 > 0$$

# Neutrinos born in Cosmic ray collisions and on earth

$$|\Delta m_{31}^2| = \begin{cases} (2.43 \pm 0.06) \times 10^{-3} \text{ eV}^2 & \text{normal hierarchy,} \\ (2.38 \pm 0.06) \times 10^{-3} \text{ eV}^2 & \text{inverted hierarchy,} \end{cases}$$

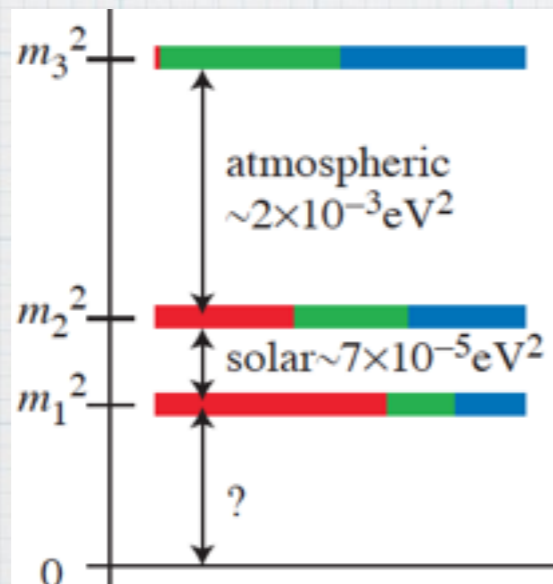
$$\Delta m_{atm}^2 = |\Delta m_{31}^2|$$

$$\Delta m_{31}^2 = |m_3|^2 - |m_1|^2$$

## Normal hierarchy

$$|m_{\nu_1}| < |m_{\nu_2}| \ll |m_{\nu_3}|$$

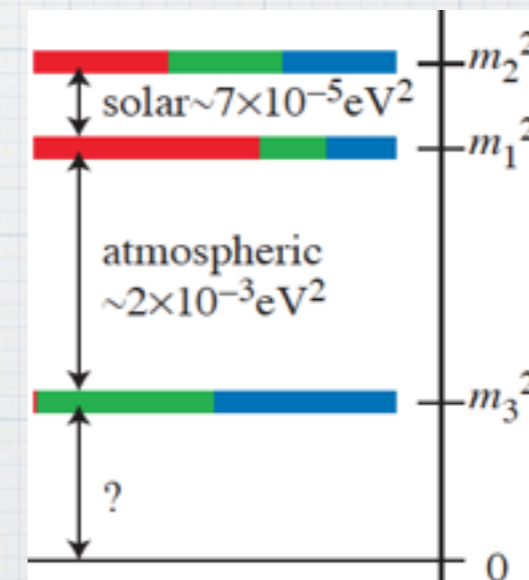
$$m_1 \simeq 0, m_2^2 \simeq \Delta m_{\odot}^2, \text{ and } m_3^2 \simeq \Delta m_{atm}^2$$



## Inverted hierarchy

$$|m_{\nu_1}| \simeq |m_{\nu_2}| \gg |m_{\nu_3}|$$

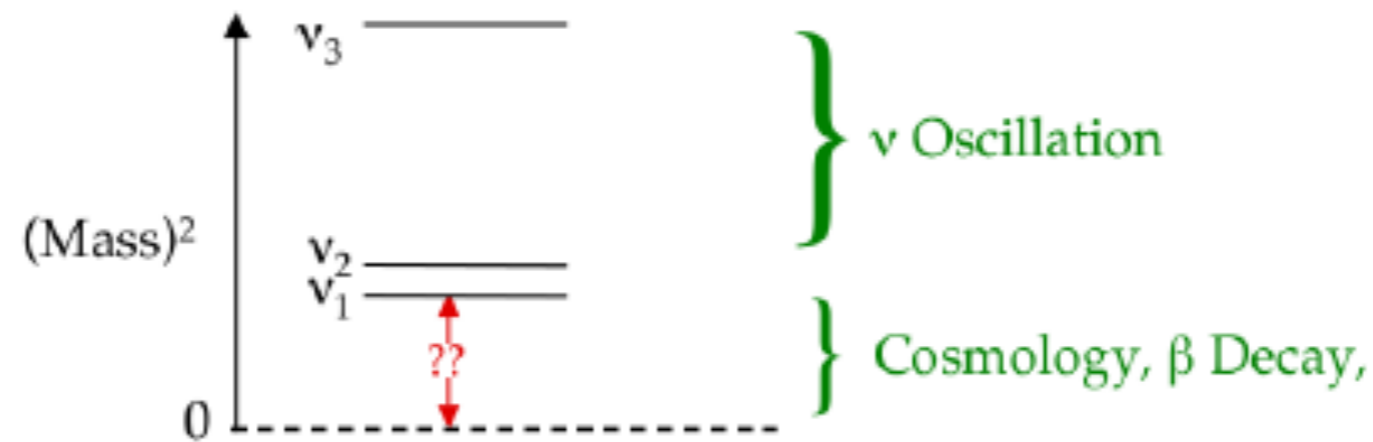
$$m_1^2 \simeq m_2^2 \simeq \Delta m_{atm}^2 \gg m_3^2$$



$\Delta m^2$

# Absolute Neutrino Mass Scale

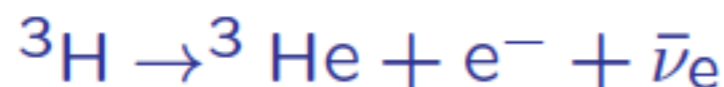
## The Absolute Scale of Neutrino Mass



How far above zero  
is the whole pattern?

Oscillation Data  $\Rightarrow \sqrt{\Delta m^2_{\text{atm}}} < \text{Mass}[\text{Heaviest } \nu_i]$

The Troitzk and Mainz  ${}^3\text{H}$   $\beta$ -decay experiments



$m_{\bar{\nu}_e} < 2.05 \text{ eV}$  (95% C.L.)

KATRIN

0.2 eV



Improved  $\beta$  energy resolution requires a **BIG**  $\beta$  spectrometer.

KATRIN

2006年11月25日

德國

Eggenstein-  
Leopoldshafen



Improved  $\beta$  energy resolution requires a **BIG**  $\beta$  spectrometer.

# KATRIN

**KATRIN JOURNEY**  
**9000KM**  
**(400KM)**



2006年11月25日  
德國



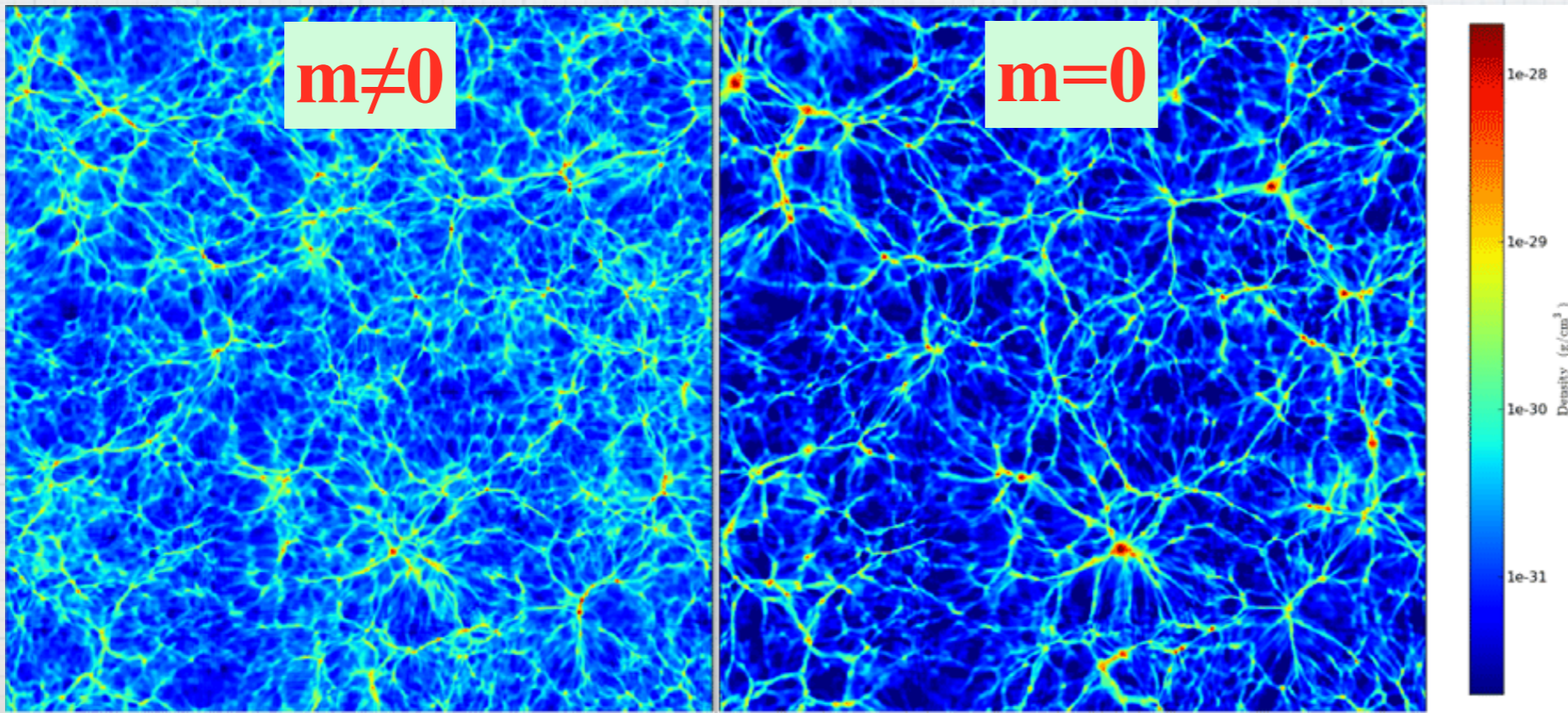
# The cosmological bound on $m_\nu$

- Number density

$$n_\nu = \int \frac{d^3 p}{(2\pi)^3} f_\nu(p, T_\nu) = \frac{3}{11} n_\gamma = \frac{6\zeta(3)}{11\pi^2} T_{CMB}^3$$

- Energy density

$$\rho_{\nu_i} = \int \sqrt{p^2 + m_{\nu_i}^2} \frac{d^3 p}{(2\pi)^3} f_\nu(p, T_\nu) \rightarrow \begin{cases} \frac{7\pi^2}{120} \left(\frac{4}{11}\right)^{4/3} T_{CMB}^4 & \text{Massless} \\ m_{\nu_i} n_\nu & \text{Massive } m_\nu \gg T \end{cases}$$



# The cosmological bound on $m_\nu$

- Number density

$$n_\nu = \int \frac{d^3 p}{(2\pi)^3} f_\nu(p, T_\nu) = \frac{3}{11} n_\gamma = \frac{6\zeta(3)}{11\pi^2} T_{CMB}^3$$

Contribution to the energy density of the Universe

$$\Omega_\nu h^2 = 1.7 \times 10^{-5}$$

Massless

$$\Omega_\nu h^2 = \frac{\sum_i m_i}{94.1 \text{ eV}}$$

Massive  $m_\nu \gg T$

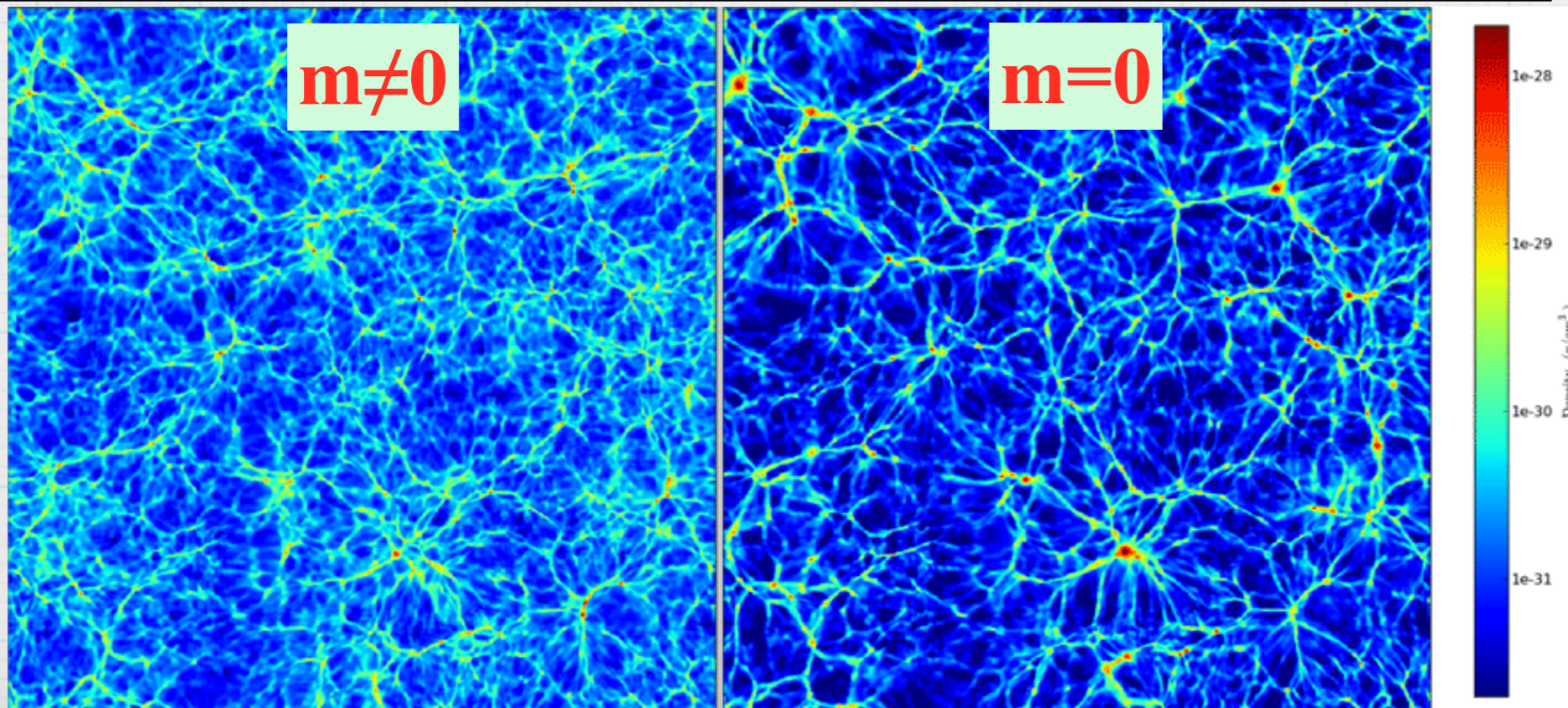
The best bound to absolute values of neutrino masses from Cosmology



$$\sum_i m_{\nu_i} < 0.23 \text{ eV}$$

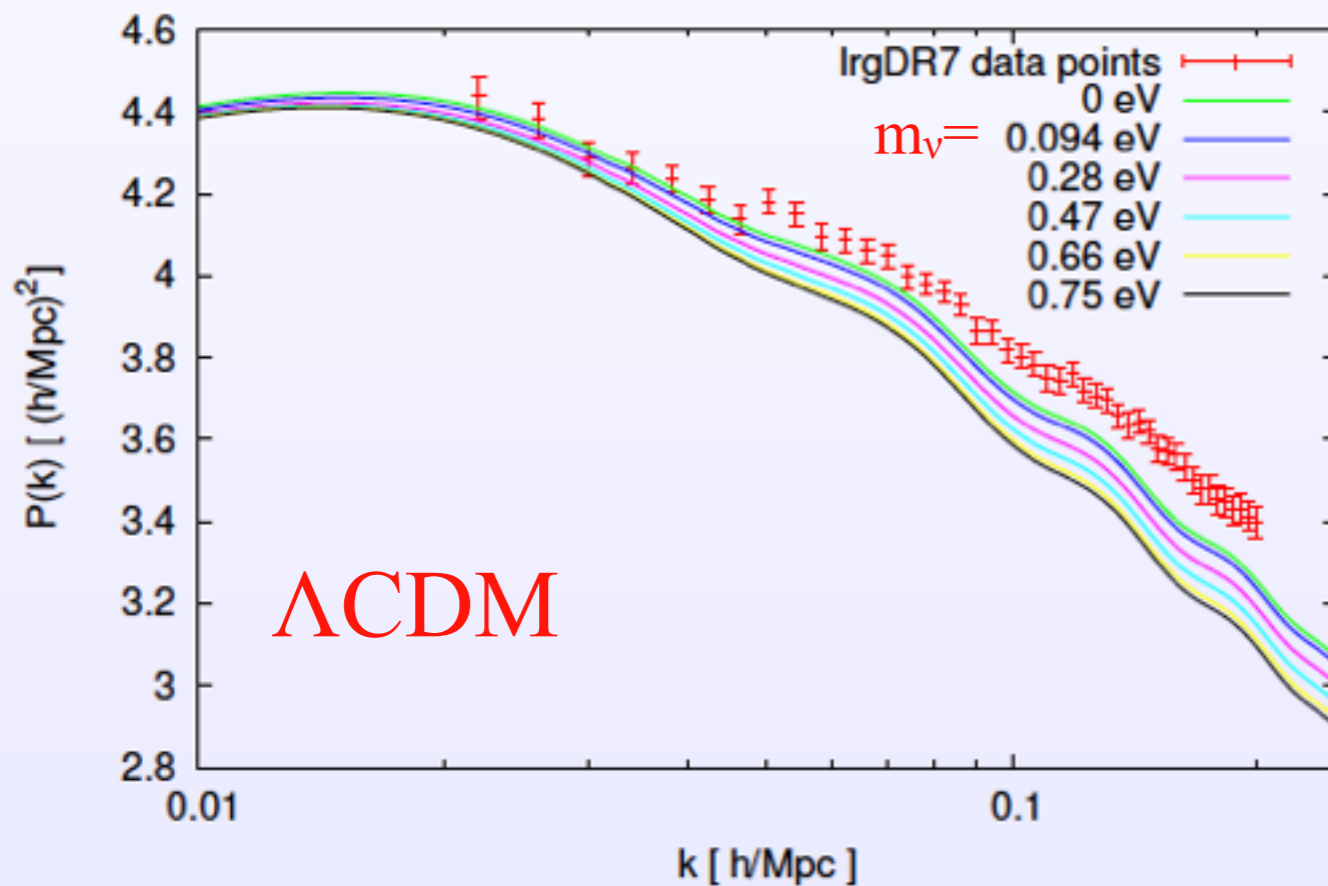
95%CL (Planck+other data)

$\Lambda$ CDM 

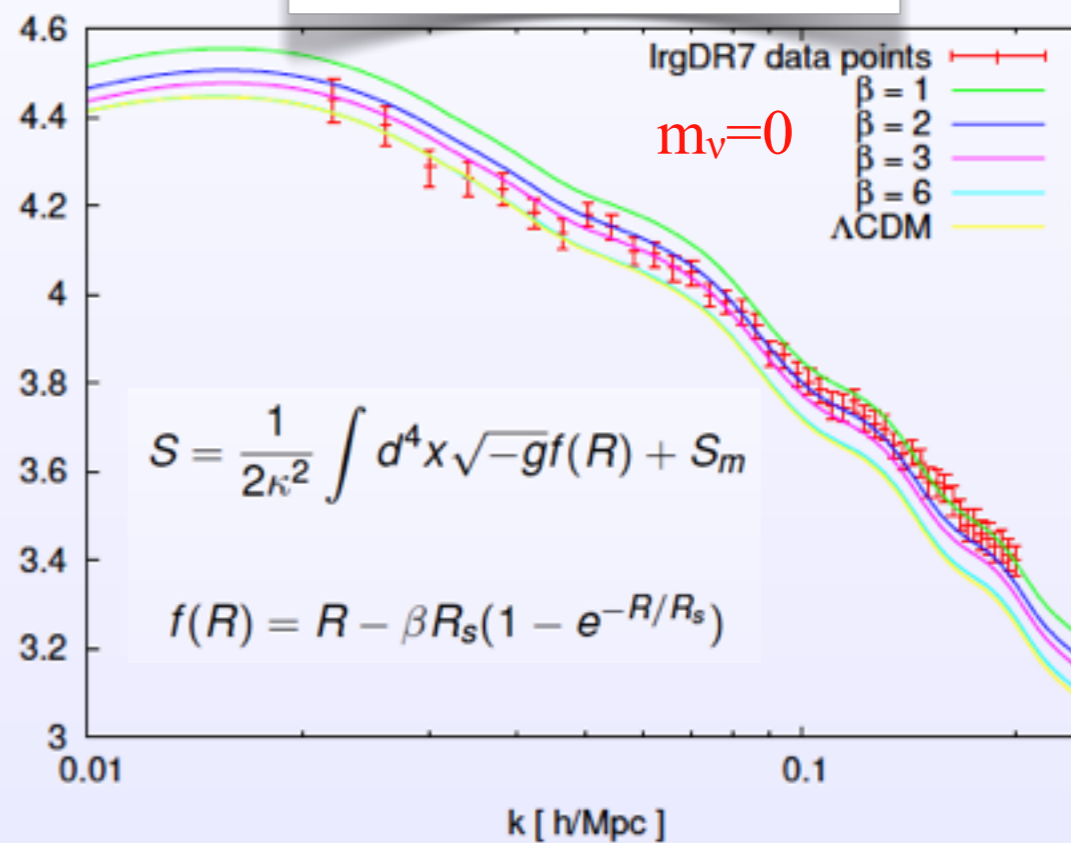


# Matter power spectrum in $\Lambda$ CDM and $f(R)$

$\Lambda$ CDM with different  $m_\nu$



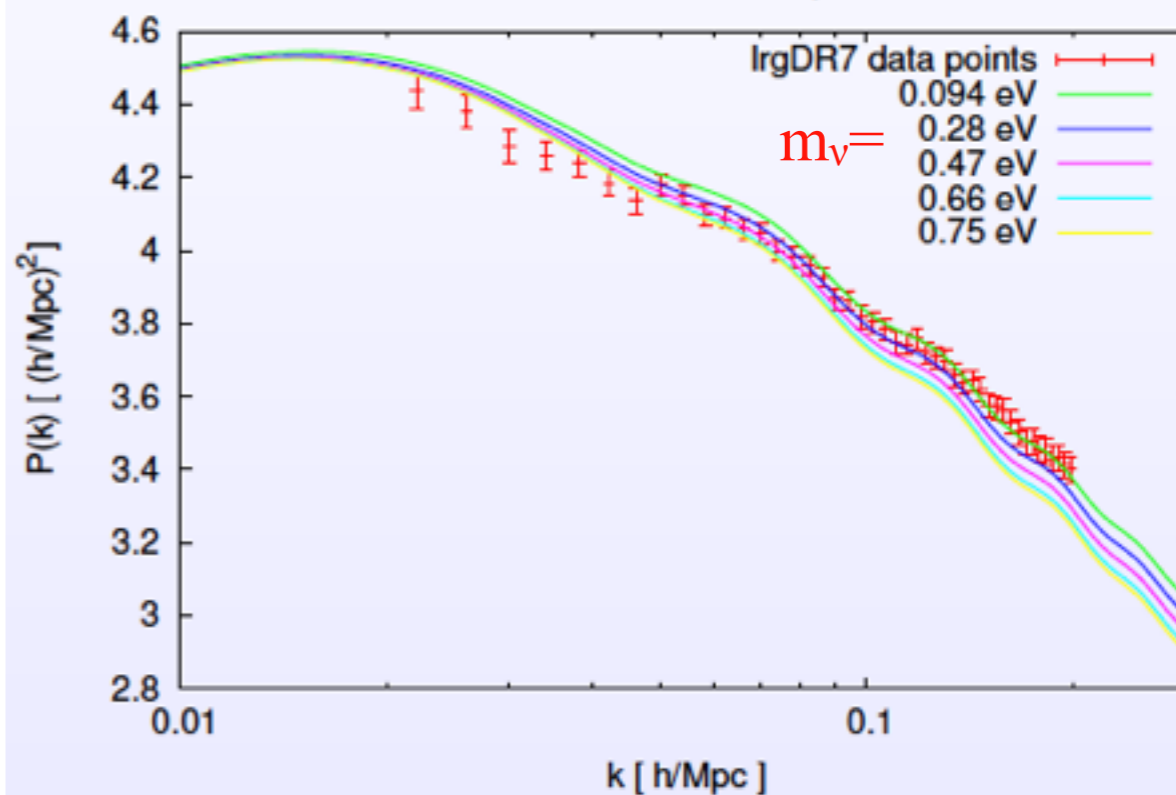
Exponential  $f(R)$  model



$f(R)$ model	$\Sigma m_\nu$
$\Lambda$ CDM	$< 0.200$ eV
Starobinsky	$0.248^{+0.203}_{-0.232}$ eV
Exponential	$< 0.214$ eV

CQG+CC.Lee, J.L.Shen,  
PLB740,285(2015)

$\beta = 1$ , different  $m_\nu$



In terms of the PMNS mixing matrix

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$V_{PMNS} = \begin{bmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{bmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\alpha_{21}/2} & 0 \\ 0 & 0 & e^{i\alpha_{31}/2} \end{pmatrix}$$

PDG2016

$$\sin^2 \theta_{12} = 0.308 \pm 0.017, \sin^2 \theta_{23} = 0.437^{+0.033}_{-0.023}, \sin^2 \theta_{13} = 0.0234^{+0.020}_{-0.019}, \delta = (1.39^{+0.38}_{-0.27})\pi$$

NH

$$\sin^2 \theta_{12} = 0.308 \pm 0.017, \sin^2 \theta_{23} = 0.455^{+0.039}_{-0.031}, \sin^2 \theta_{13} = 0.0240^{+0.019}_{-0.020}, \delta = (1.31^{+0.29}_{-0.33})\pi$$

IH

### Bimaximal Matrix

$$\begin{pmatrix} \frac{\sqrt{2}}{2} & \frac{\sqrt{2}}{2} & 0 \\ -\frac{1}{2} & \frac{1}{2} & \frac{\sqrt{2}}{2} \\ \frac{1}{2} & -\frac{1}{2} & \frac{\sqrt{2}}{2} \end{pmatrix}$$

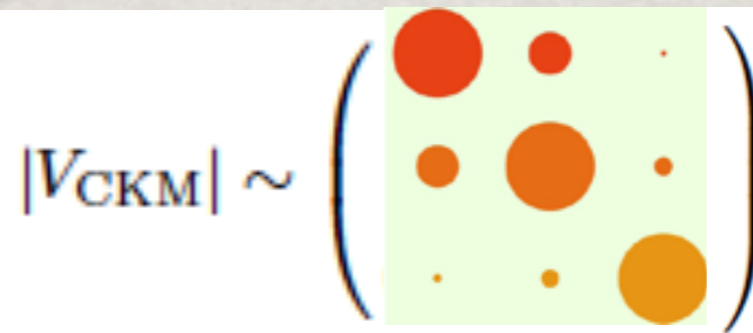
$$\theta_{12} = 45^\circ, \theta_{23} = 45^\circ, \theta_{13} = 0$$



### Tribimaximal Matrix

$$\begin{pmatrix} \frac{\sqrt{6}}{3} & \frac{\sqrt{3}}{3} & 0 \\ -\frac{\sqrt{6}}{6} & \frac{\sqrt{3}}{3} & \frac{\sqrt{2}}{2} \\ \frac{\sqrt{6}}{6} & -\frac{\sqrt{3}}{3} & \frac{\sqrt{2}}{2} \end{pmatrix}$$

$$\theta_{12} = 35.3^\circ, \theta_{23} = 45^\circ, \theta_{13} = 0$$



center values

$$\theta_{12} \approx 34^\circ$$

$$\theta_{23} \approx 42^\circ$$

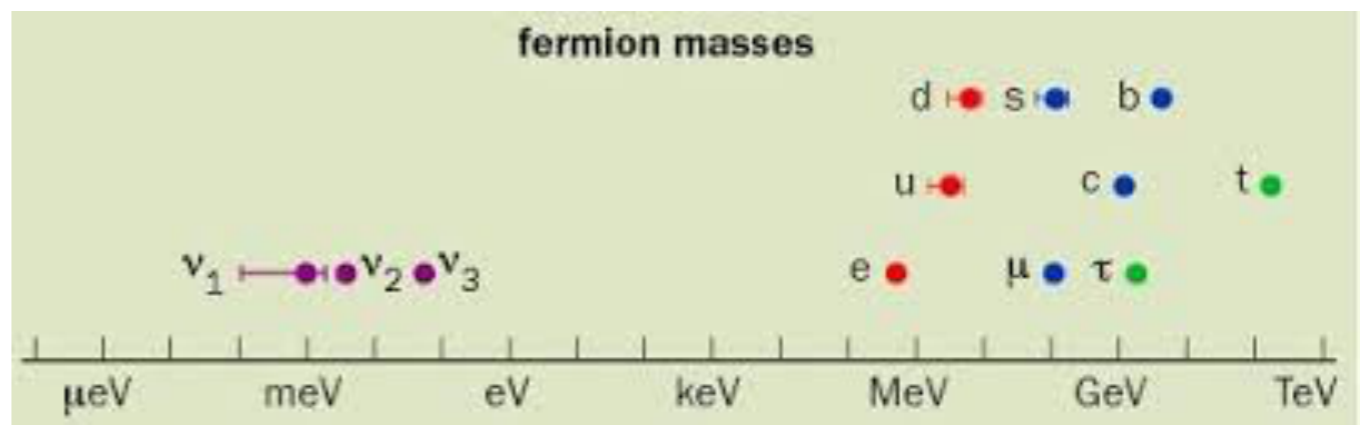
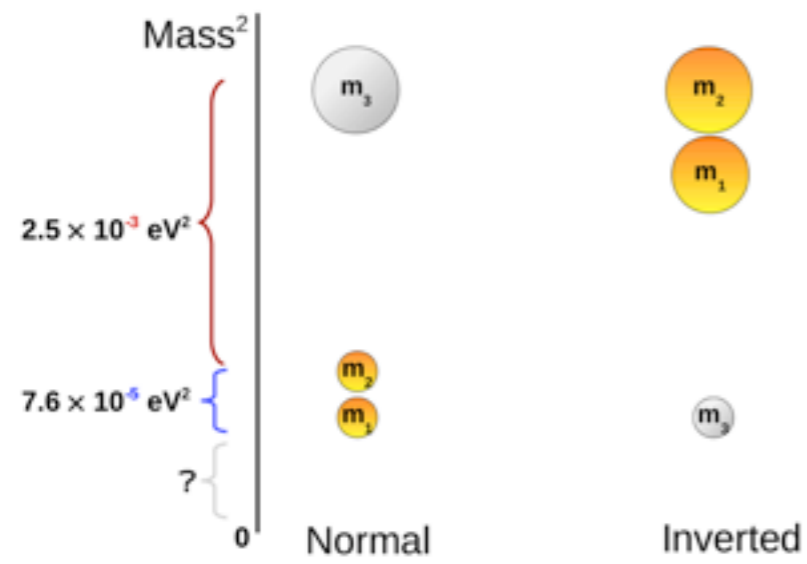
$$\theta_{13} \approx 9^\circ$$

$$\delta \approx 1.4\pi$$

Daya-Bay

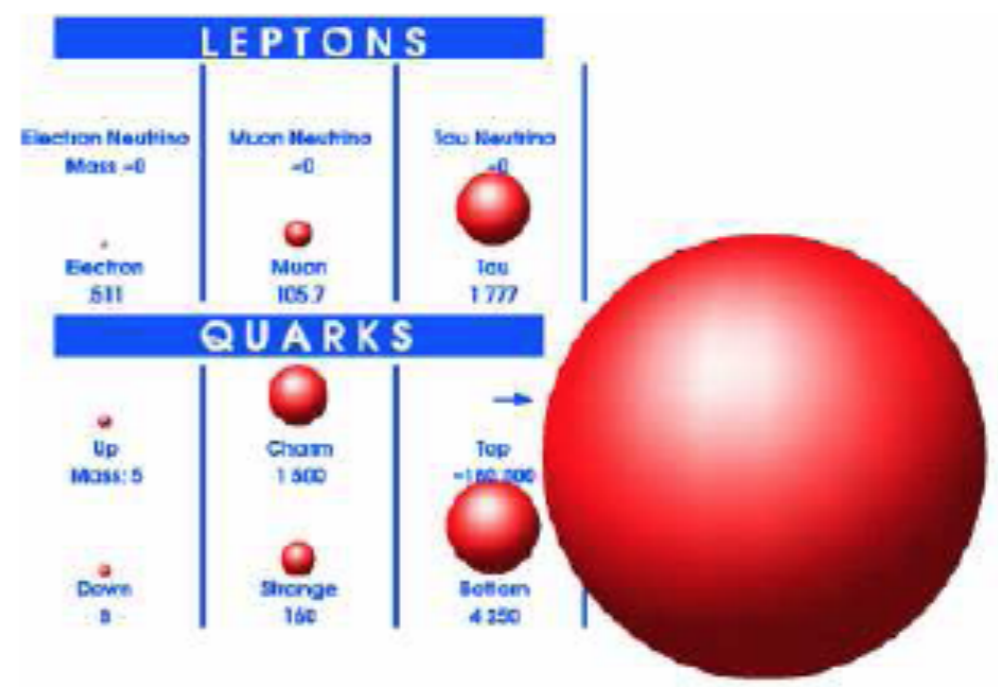
● A brief overview of neutrino mass generation

**Fermion Mass Problem**



$m_{\nu_j} \ll\ll m_{e,\mu,\tau}, m_q, q = u, c, t, d, s, b$

For  $m_{\nu_j} \lesssim 1 \text{ eV}$ :  $m_{\nu_j}/m_{l,q} \lesssim 10^{-6}$



**Questions:** Where do the

- quark mass hierachy
- small neutrino masses
- small quark mixings and
- large lepton mixings originate from?

如果中微子有質量，它們的質量為什麼遠小於相對的帶電輕子及夸克的質量？

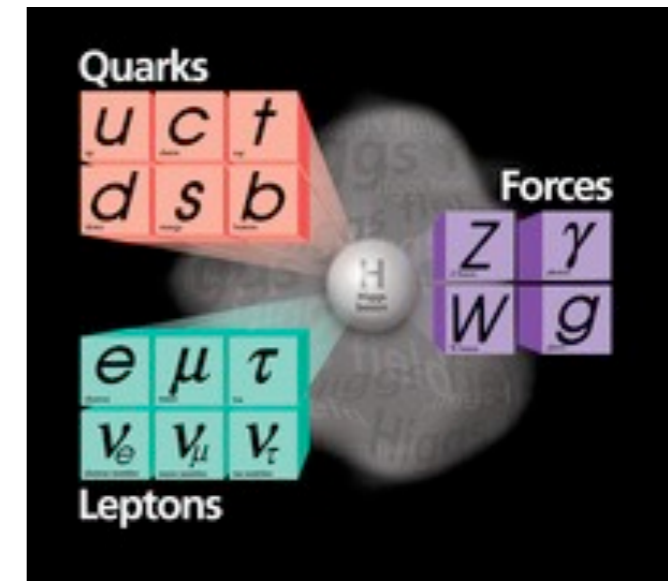
• The standard model:  $SU(3)_C \times SU(2)_L \times U(1)_Y$

$$Q_L : \begin{pmatrix} u \\ d \end{pmatrix}_L \quad \begin{pmatrix} c \\ s \end{pmatrix}_L \quad \begin{pmatrix} t \\ b \end{pmatrix}_L \quad L_L : \begin{pmatrix} \nu_e \\ e \end{pmatrix}_L \quad \begin{pmatrix} \nu_\mu \\ \mu \end{pmatrix}_L \quad \begin{pmatrix} \nu_\tau \\ \tau \end{pmatrix}_L$$

$$U_R : u_R \quad c_R \quad t_R$$

$$D_R : d_R \quad s_R \quad b_R \quad E_R : e_R \quad \mu_R \quad \tau_R$$

$$\text{Higgs : } H^0 \quad \text{Gauge Bosons : } W^\pm, Z, \gamma, g$$



$$\text{Yukawa interactions: } Y = \sum_{i,j} h_{ij}^d \bar{Q}_L \phi D_R + h_{ij}^u \bar{Q}_L \tilde{\phi} U_R + h_{ij}^e \bar{L}_L \phi E_R + h.c.$$

$$\Phi = \Phi_0 = (-\mu^2/2\lambda)^{1/2}$$

SSB

$$V_L^{d+} M_d V_R^d = M_d^{diag.}, \quad D_{L(R)j} = (V_{L(R)}^d)_{ji} D'_{L(R)i}$$

$$V_L^{u+} M_u V_R^u = M_u^{diag.}, \quad U_{L(R)j} = (V_{L(R)}^u)_{ji} U'_{L(R)i}$$

■ What about neutrinos?

■ Do neutrinos get their masses like charged fermions?



# Neutrino masses: Dirac or Majorana

Dirac neutrino mass:

$$\mathcal{L}_D = -m_D \bar{\nu}_L \nu_R + \text{h.c.}$$



**Introduce  $\nu_R$   
(not in the SM)**

😊 the lepton number L is conserved

Majorana neutrino mass:

$$\mathcal{L}_M = -m_M \bar{\nu}^c \nu + \text{h.c.} \quad \nu \leftrightarrow \bar{\nu}$$



**FORBIDDEN  
IN THE SM.**

• the lepton number L is violated

*Neutrino oscillations measure  $\Delta m^2$ , but they do not provide information about the absolute neutrino spectrum and cannot distinguish pure Dirac and Majorana neutrinos.*

***In the SM:***

- No Dirac mass term (no right-handed neutrino).
- No Majorana mass term either ( $\nu_L$  is an SU(2) doublet).

	$SU(3)_C \otimes SU(2)_L \otimes U(1)_Y$
$L_a = (\nu_a, l_a)_L^T$	(1, 2, -1)
$e_a^c$	(1, 1, 2)
$Q_a = (u_a, d_a)_L^T$	(3, 2, 1/3)
$u_a^c$	( $\bar{3}$ , 1, -4/3)
$d_a^c$	( $\bar{3}$ , 1, 2/3)
$\Phi$	(1, 2, 1)

Table 1: Matter and scalar multiplets of the Standard Model (SM)

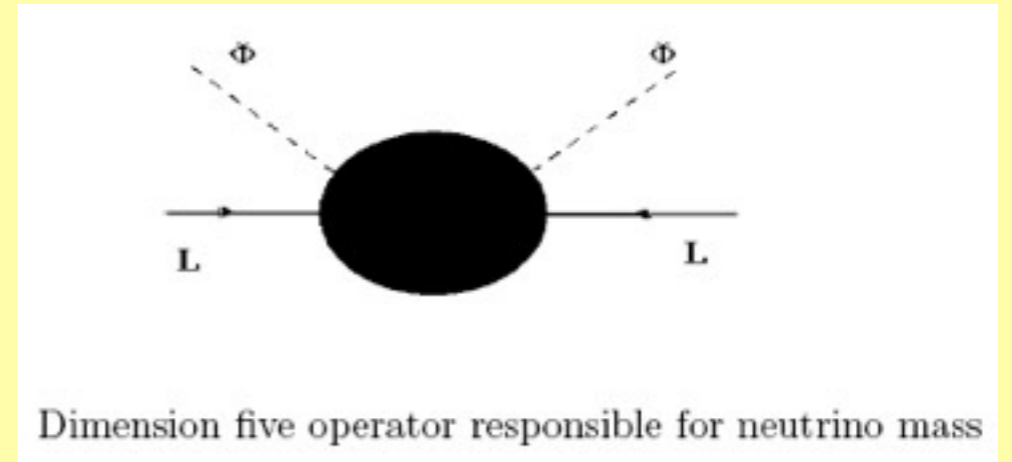
■ **Effective Dim-5 operator:**

$$O = (\lambda_0/M_X)L\Phi L\Phi$$

↓ *SSB*

$$m_\nu = \lambda_0 \frac{\langle \Phi \rangle^2}{M_X}, \quad (\text{Majorana})$$

S. Weinberg, Phys. Rev. D22, 1694 (1980).



For  $\lambda_0 \sim 1$ ,  $\langle \Phi \rangle \sim 100 \text{ GeV}$ ,  $M_X \sim M_P \rightarrow m_\nu \sim 10^{-6} \text{ eV}$  (too small)

$$\Delta m_{21}^2 \sim 7 \times 10^{-5} \text{ eV}^2 \quad |\Delta m_{31}^2| \sim 2 \times 10^{-3} \text{ eV}^2$$

**Neutrino masses beyond the SM:**

■ **If there are right handed neutrinos  $\nu_R$  :**  $\mathbf{v}_R = (1, 1, 0)$

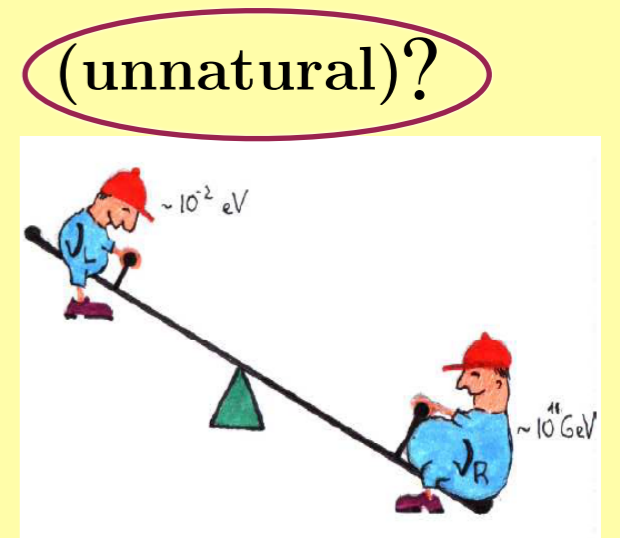
$$\mathcal{L}_Y = Y_\nu \bar{L} \Phi \nu_R + h.c. \Rightarrow m_\nu^D = Y_\nu \langle \Phi \rangle$$

The observed neutrino masses would require  $Y_\nu \leq 10^{-13} - 10^{-12}$  (unnatural)?

■ **Majorana mass for  $\nu_R$ :**  $M_{R\nu_R^T} C^{-1} \nu_R + h.c.$

**Type-I see-saw mechanism:**  $\mathcal{M}_\nu = -m_D^T M_R^{-1} m_D.$

(naturally small? + Majorana)



# The Seesaw Mechanism

The Seesaw mechanism refers to the neutrino mass matrix of the form:

$$L_m = -\frac{1}{2} (\nu_L^c, \nu_R) \begin{pmatrix} 0 & m_D \\ m_D^T & M_R \end{pmatrix} \begin{pmatrix} \nu_L \\ \nu_R^c \end{pmatrix}$$

For one generation, if  $M_R \gg m_D$ , the diagonal masses are

$$m_\nu \approx -m_D M_R^{-1} m_D^T, \quad m_N \approx M_R$$

$$\begin{aligned} m_{\nu_1} &= m_e^2 / M_R \\ m_{\nu_2} &= m_\mu^2 / M_R \\ m_{\nu_3} &= m_\tau^2 / M_R \end{aligned}$$

How large  $M_R$  needs to be?

For  $m_{\nu_1} = 0.1\text{eV}$ ,  $M_R = 2.5\text{TeV}$

For  $m_{\nu_2} = 0.1\text{eV}$ ,  $M_R = 10^8\text{GeV}$

For  $m_{\nu_3} = 0.1\text{eV}$ ,  $M_R = 3 \times 10^{10}\text{GeV}$ .

A very nice way to explain why light neutrino masses are so much lighter than their charged lepton partners

(Minkowski (1977); Gell-Mann, Ramond, and Slansky (1979); Yanagida (1979); Glashow (1980); Mohapatra and Senjanovic(1980))

# Inverse Seesaw

**Basis  $(\nu, \nu^c, S)$ :**

$$M_\nu = \begin{pmatrix} 0 & m_D & 0 \\ m_D^T & 0 & M \\ 0 & M^T & \mu \end{pmatrix},$$

Mohapatra & Valle,  
1986

**After EWSB, the effective light neutrino mass matrix is given by**

$$M_\nu = m_D M^{T-1} \mu M^{-1} m_D^T.$$

**“Inverse” seesaw, because:**

$$M_\nu \Rightarrow 0 \quad \text{IF} \quad \mu \Rightarrow 0$$

# Effective Operators

**d = 5:**

$$\mathcal{O}_W \propto \frac{c_{ij}}{\Lambda} (L_i H)(L_j H)$$

Weinberg, 1980

**d = 7:**

$$\mathcal{O}_2 \propto LLLe^c H$$

Babu & Leung, 2001

$$\mathcal{O}_3 \propto LLQd^c H$$

de Gouvea & Jenkins, 2007

$$\mathcal{O}_4 \propto LL\bar{Q}\bar{u}^c H$$

$$\mathcal{O}_8 \propto L\bar{e}^c \bar{u}^c d^c H$$

$$\mathcal{O} \propto (LH)(LH)(H_u H_d)$$

**d = 9:**

$$\mathcal{O}_5 \propto LLQd^c H H H^\dagger$$

$$\mathcal{O}_6 \propto LL\bar{Q}\bar{u}^c H H H^\dagger$$

$$\mathcal{O}_7 \propto LQ\bar{e}^c \bar{Q} H H H^\dagger$$

$$\mathcal{O}_9 \propto LLLe^c Le^c \quad .$$

$$\mathcal{O}_{10} \propto LLLe^c Qd^c$$

$$\mathcal{O}_{11} \propto LLQd^c Qd^c$$

.....

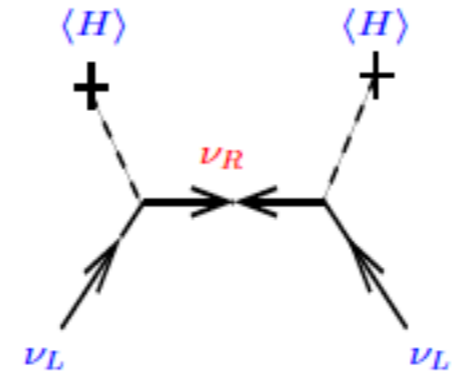
# Example realization:

d = 5:

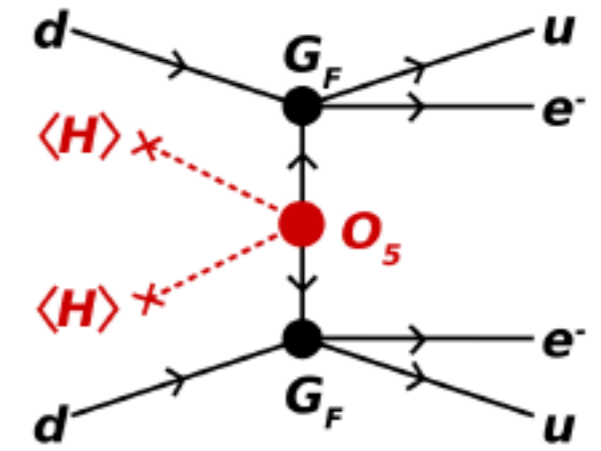
$$\mathcal{O}_W \propto \frac{c_{ij}}{\Lambda} (L_i H)(L_j H)$$

$$\Lambda \simeq M_{\nu R_k}$$

$$c_{ij} \propto Y_{ik}^\nu Y_{jk}^\nu$$

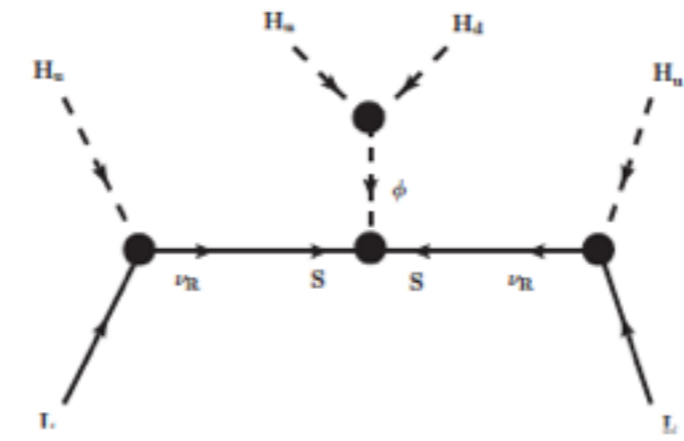


$0\nu\beta\beta$  decay:

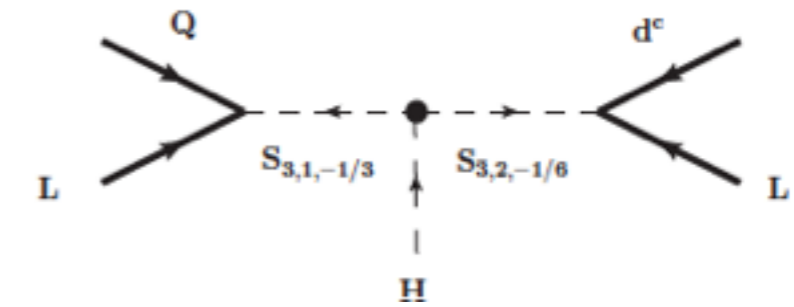
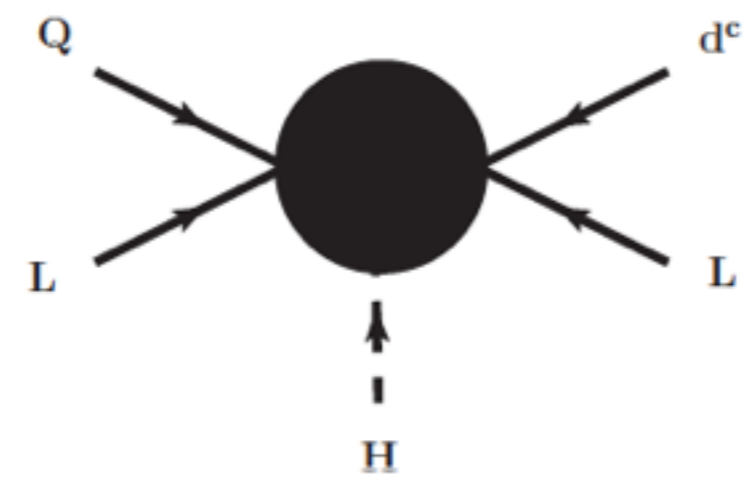


d = 7:

$$\mathcal{O} \propto (LH)(LH)(H_u H_d)$$



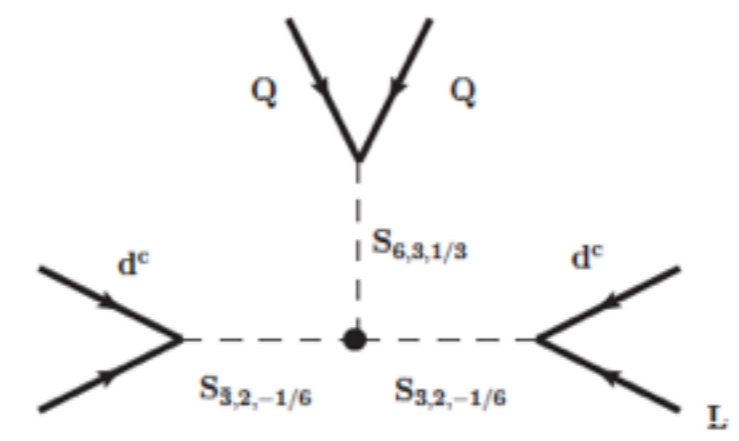
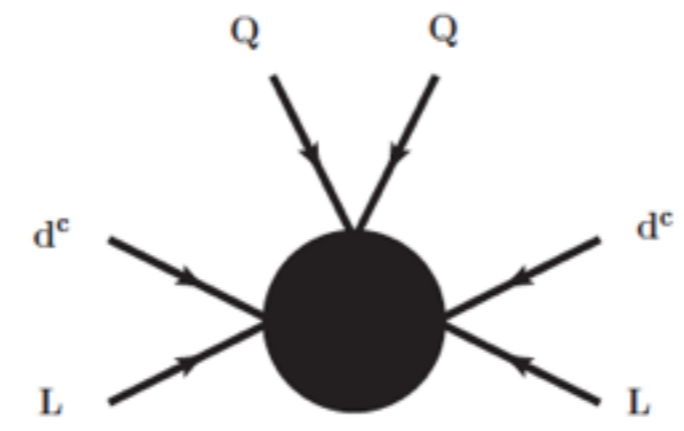
$$\mathcal{O}_3 \propto LLQd^c H$$



$S_{3,1,-1/3}$  - singlet leptoquark  
 $S_{3,2,1/6}$  - doublet leptoquark

d = 9:

$$\mathcal{O}_{11} \propto LLQd^c Qd^c$$



# ● 產生Majorana中微子質量簡介

Without  $\nu_R$

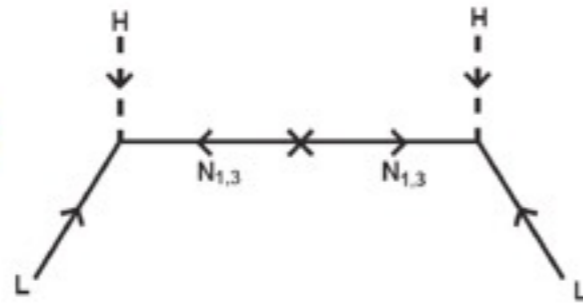
## 📌 Tree level

Minkowski 1977; ...

Xiao-Gang He et al 1989

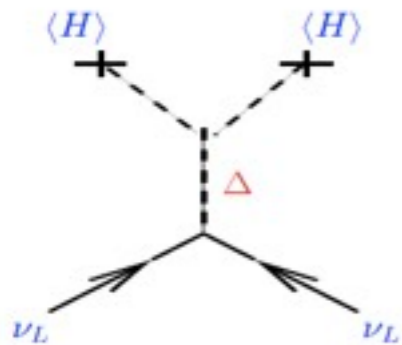
**Type-III See-Saw  
Xiao-Gang He**

Type (I,III) seesaw



$$N_1 : (1, 1, 0)$$

$$N_3 : (1, 3, 0)$$



Type II seesaw

Schechter & Valle, 1980, 1982  
Cheng & Li, 1980  
Mohapatra, Senjanovic, 1981  
...

$$\Delta \equiv \begin{pmatrix} H^- & -\sqrt{2} H^0 \\ \sqrt{2} H^{--} & -H^- \end{pmatrix} = (1, 3, 2) \quad \text{scalar triplet}$$

$$\mathcal{L}_{\text{Type II}} = \text{Tr}(D_\mu \Delta)^\dagger (D^\mu \Delta) - (Y_\nu l_L^T C i\sigma_2 \Delta l_L + \text{h.c.}) - V(H, \Delta),$$

$$V(H, \Delta) = M_\Delta^2 \text{Tr} \Delta^\dagger \Delta + (\mu H^T i\sigma_2 \Delta^\dagger H + \text{h.c.}) + \lambda_1 (H^\dagger H) \text{Tr} \Delta^\dagger \Delta + \lambda_2 (\text{Tr} \Delta^\dagger \Delta)^2 + \lambda_3 \text{Tr} (\Delta^\dagger \Delta)^2 + \lambda_4 H^\dagger \Delta \Delta^\dagger H.$$

$$M_\nu = \sqrt{2} Y_\nu v_\Delta$$

$$v_\Delta = \frac{\mu v^2}{\sqrt{2} M_\Delta^2}$$

$$M_\Delta \sim 250 \text{ GeV}, \mu \sim 0.1 \text{ eV}, Y_\nu \sim 1$$

→  $M_\nu \sim 0.1 \text{ eV}$

## Type III seesaw

Foot, Lew, X.G.He and Joshi, 1989

$$\Sigma = \begin{pmatrix} \Sigma_L^0/\sqrt{2} & \Sigma_L^+ \\ \Sigma_L^- & -\Sigma_L^0/\sqrt{2} \end{pmatrix} \quad \text{the triplet } \Sigma=(1, 3, 0)$$

$$\mathcal{L} = Tr[\bar{\Sigma}i\not{D}\Sigma] - \frac{1}{2}Tr[\bar{\Sigma}M_\Sigma\Sigma^c + \bar{\Sigma}^c M_\Sigma^*\Sigma] - \tilde{H}^\dagger \bar{\Sigma}^c \sqrt{2}Y_\Sigma L_L - \bar{L}_L \sqrt{2}Y_\Sigma^\dagger \Sigma^c \tilde{H}.$$

The mass terms are given by:

$$\mathcal{L}_\uparrow = -(\bar{l}_R \ \bar{\Psi}_R) \begin{pmatrix} m_l & 0 \\ Y_\Sigma v & M_\Sigma \end{pmatrix} \begin{pmatrix} l_L \\ \Psi_L \end{pmatrix} - (\bar{\nu}_L^c \ \bar{\Sigma}_L^{0c}) \begin{pmatrix} 0 & Y_\Sigma^T v/2\sqrt{2} \\ Y_\Sigma v/2\sqrt{2} & M_\Sigma/2 \end{pmatrix} \begin{pmatrix} \nu_L \\ \Sigma_L^0 \end{pmatrix}$$

$$v \equiv \sqrt{2}\langle\phi^0\rangle = 246 \text{ GeV}. \quad \Psi \equiv \Sigma_L^{+c} + \Sigma_L^-$$

$$M_\Sigma \sim 100 \text{ GeV} \Rightarrow Y_\Sigma \sim 10^{-7}$$



# Loop level

## 1-loop:

1980

**a. Zee model (with charged scalar singlet and additional scalar doublets).**

$$l^T \hat{f} i \sigma_2 l \eta^+ + \sum_{i=1,2} \bar{l} \hat{f}_i e H_i$$

2006

**b. Ma model (with fermion singlet  $N_i$  and additional scalar doublet  $\eta$ ).**

$$h_{\alpha i} (\nu_{\alpha} \eta^0 - l_{\alpha} \eta^+) N_i$$

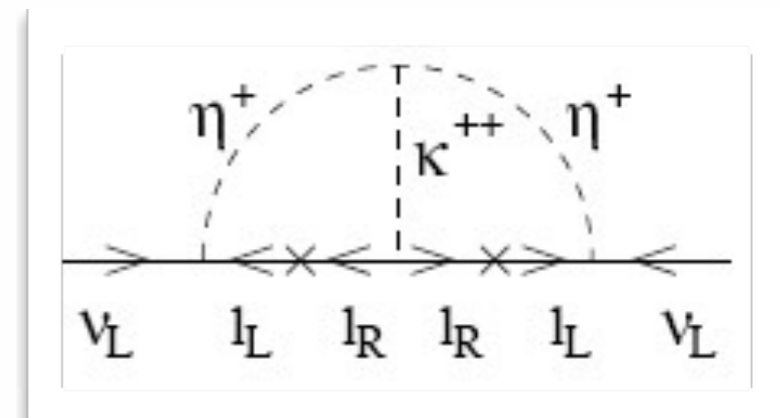
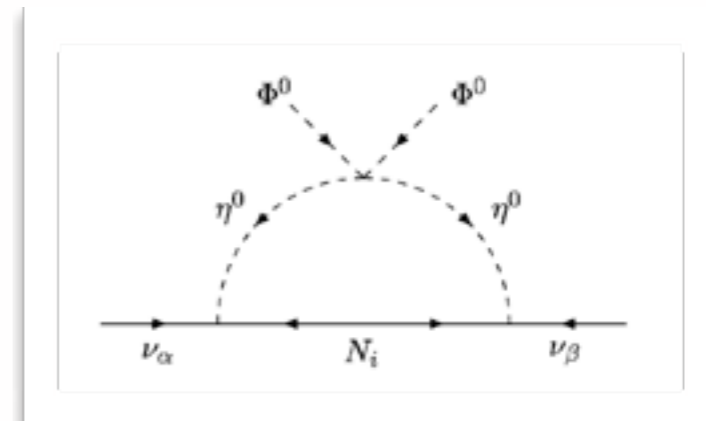
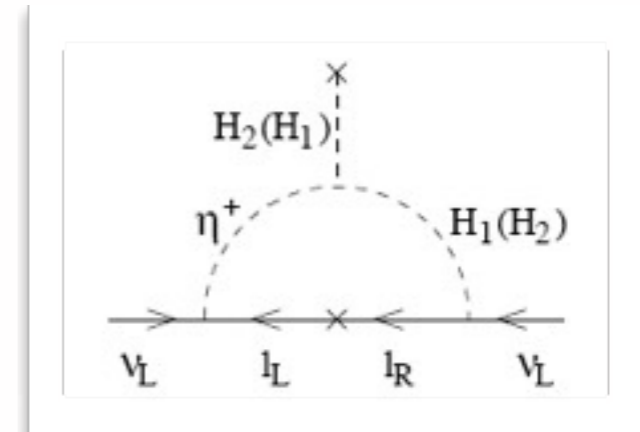
## 2-loop:

1986

1988

**Zee-Babu model (with doubly and singly charged scalars).**

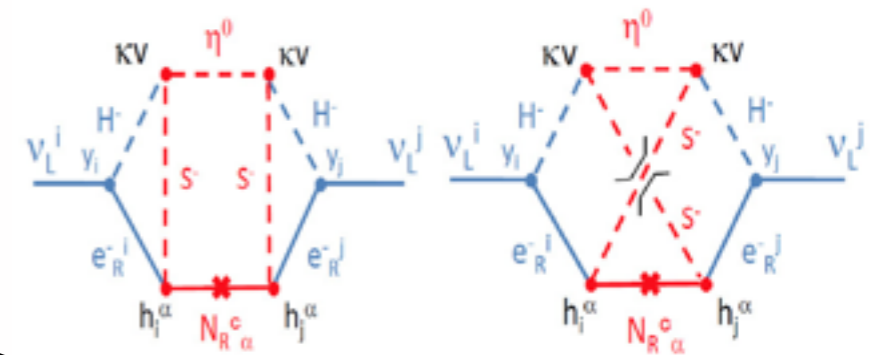
$$l^T \hat{f} l \eta^+ + l_R^T \hat{h} l_R k^{++}$$



# 3-loop:

	$Q^i$	$u_R^i$	$d_R^i$	$L^i$	$e_R^i$	$\Phi_1$	$\Phi_2$	$S^\pm$	$\eta$	$N_R^\alpha$
$Z_2$ (exact)	+	+	+	+	+	+	+	-	-	-
$\tilde{Z}_2$ (softly broken)	+	-	-	+	+	+	-	+	-	+

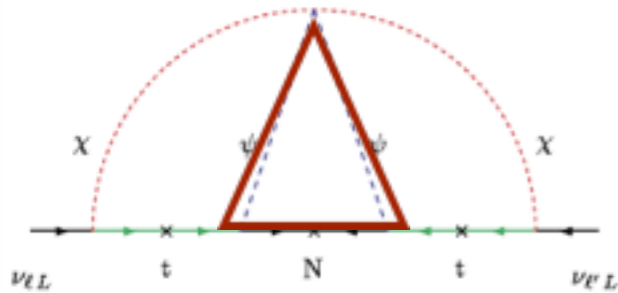
M.Aoki,S.Kanemura,O.Seto, PRL102,051805(2009)



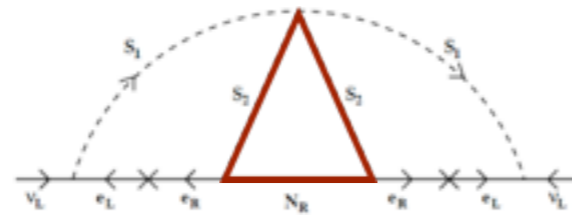
## Dark Matter

### Top quark as a dark portal

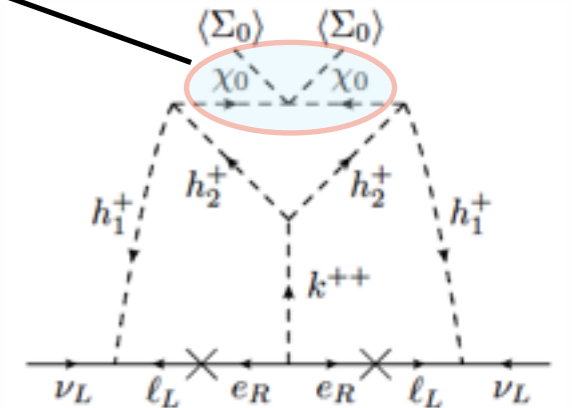
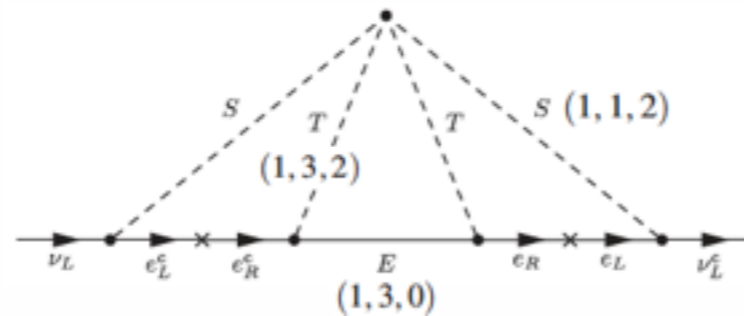
John N. Ng, Alejandro de la Puente 2013



L.Krauss,S.Nasri,and M.Trodden, 2002

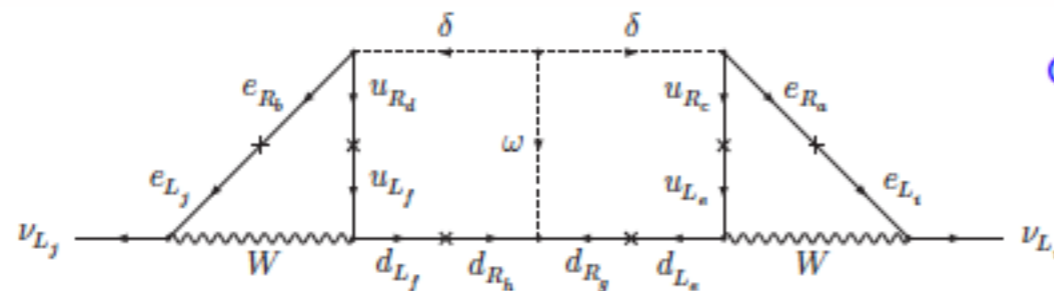


C.S.Chen,K.L.McDonald,S.Nasri, 2014

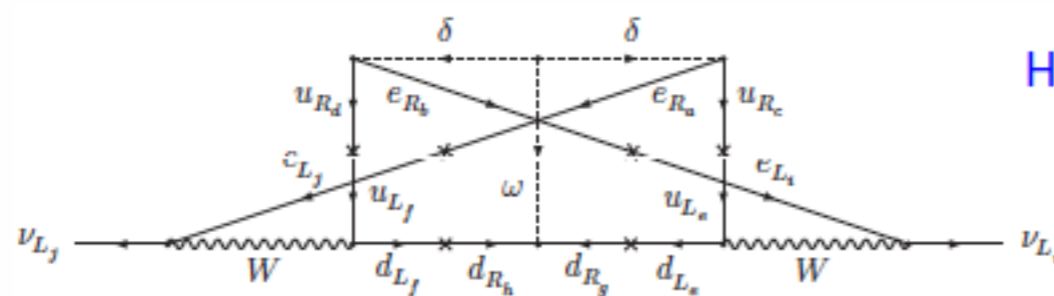


Hatanaka,Nishiwaki,Okada,Orikasa, arXiv:1412.8664

# 4-loop:



Gu, 2011

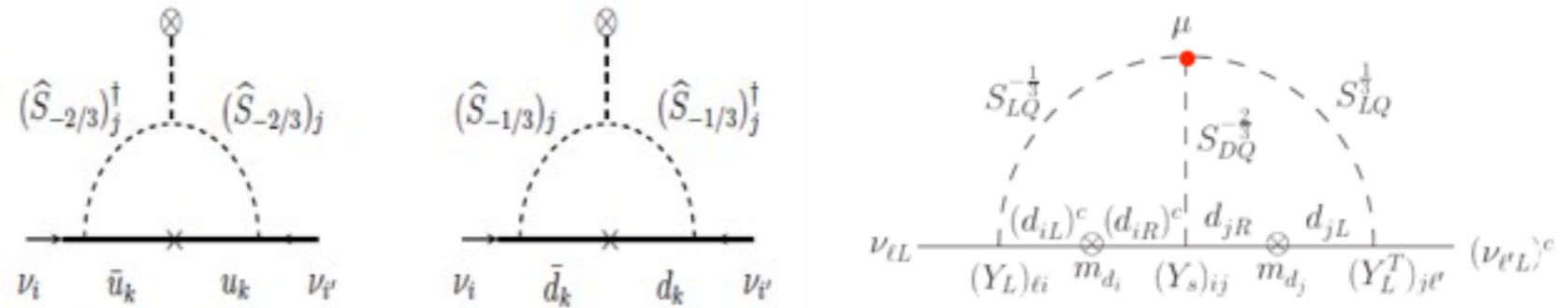


Helo et al., 2015

$m_\nu \simeq 10^{-8} \text{ eV}$   
 ... because  $d = 9$  4-loop  
 Needs (Quasi)-Dirac  $\nu$ 's  
 to explain oscillation data

# Other models with loops:

Hirsch et al. 1996, Aristizabal et al. 2008  
Leptoquarks



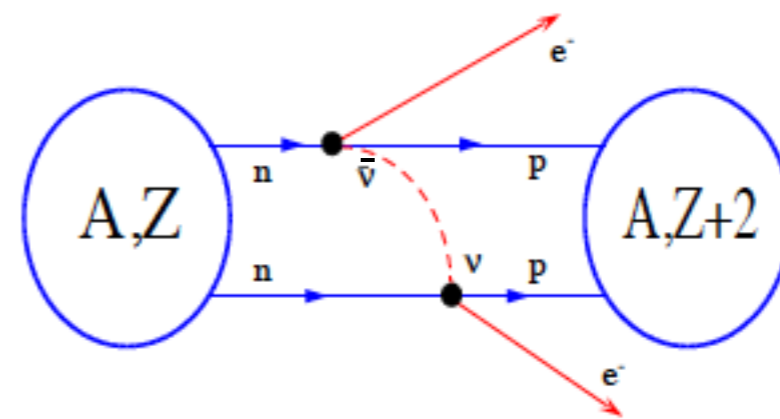
***Suppressed  $0\nu\beta\beta$  in all these loop models!***

Double beta decay:  $(A, Z) \rightarrow (A, Z + 2) + e^- + e^- + \bar{\nu}_e + \bar{\nu}_e$

Neutrinoless double beta decay:

$0\nu\beta\beta$

$(A, Z) \rightarrow (A, Z + 2) + e^- + e^-$



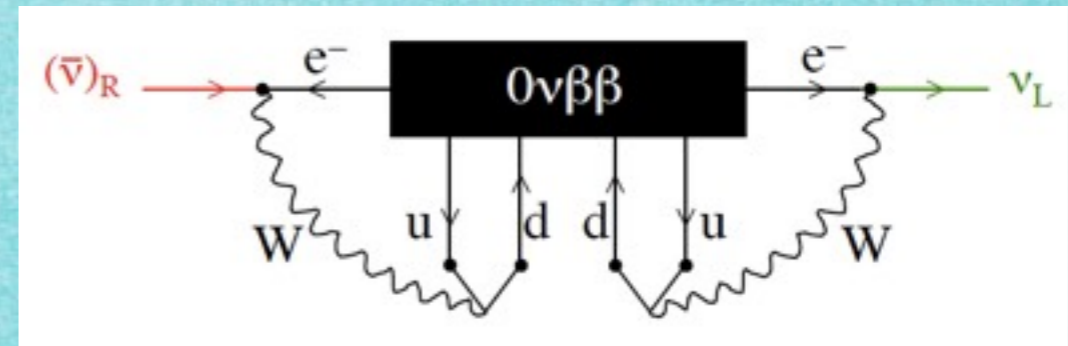
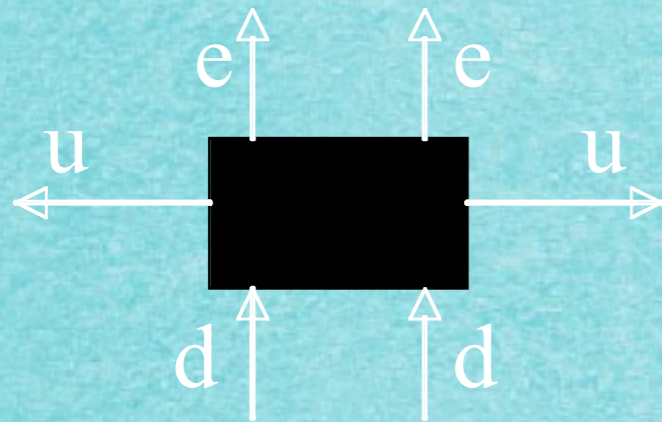
Majorana nature of  $\nu$

$$\nu \leftrightarrow \bar{\nu}$$

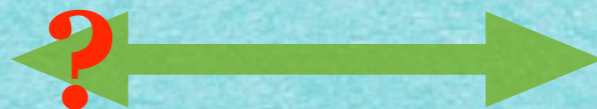
# “Black Box” theorem

*J. Schechter and J.W.F. Valle, Phys.Rev. D 25, 2951 (1982)*

“Any mechanism inducing the  $0\nu\beta\beta$  decay produces an effective Majorana neutrino mass term, which must therefore contribute to this decay.”



$0\nu\beta\beta$  decay



Majorana neutrino mass



The multi-loop with  $0\nu\beta\beta$  to  $m_\nu$  is too small,  
 $\sim O(10^{-25})$  eV.

*M.Duerr, M.Lindner, A.Merle, JHEP1106, 091 (2011) ....*

The theorem does not state if the mechanism for  $0\nu\beta\beta$  from  $M_\nu$  is the dominant one.

In some models, the dominant contributions to  $0\nu\beta\beta$  are generated without directly involving  $\nu_M$  or  $M_\nu$ .

● A special class of models to generate  $M_\nu$

C.S.Chen+CQG+J.N.Ng,  
PRD75,053004(2007)

	$SU(3)_C \otimes SU(2)_L \otimes U(1)_Y$
$L_a = (\nu_a, l_a)_L^T$	(1, 2, -1)
$e_{aL}^c$	(1, 1, 2)
$Q_a = (u_a, d_a)_L^T$	(3, 2, 1/3)
$u_{aL}^c$	( $\bar{3}$ , 1, -4/3)
$d_{aL}^c$	( $\bar{3}$ , 1, 2/3)
$\Phi$	(1, 2, 1)

Table 1: Matter and scalar multiplets of the Standard Model (SM)

**No  $\nu_R$  added**

**New scalars:** a triplet  $T$  (1,3,2) + a singlet  $\Psi$  (1,1,4)

$$\begin{aligned}
 V(\phi, T, \psi) = & -\mu^2 \phi^\dagger \phi + \lambda_\phi (\phi^\dagger \phi)^2 - \mu_T^2 \text{Tr}(T^\dagger T) + \lambda_T [\text{Tr}(T^\dagger T)]^2 + \lambda'_T \text{Tr}(T^\dagger T T^\dagger T) + m^2 \Psi^\dagger \Psi + \lambda_\Psi (\Psi^\dagger \Psi)^2 \\
 & + \kappa_1 \text{Tr}(\phi^\dagger \phi T^\dagger T) + \kappa_2 \phi^\dagger T T^\dagger \phi + \kappa_\Psi \phi^\dagger \phi \Psi^\dagger \Psi + \rho \text{Tr}(T^\dagger T \Psi^\dagger \Psi) \\
 & + [\lambda(\phi^T T \phi \Psi^\dagger) - M(\phi^T T^\dagger \phi) + \text{H.c.}]
 \end{aligned}$$

New Yukawa term:

$$Y_{ab} \bar{l}_{aR}^c l_{bR} \Psi$$

lepton # for  $\Psi$  is 2

No Yukawa coupling for the triplet:

~~$LLT$~~

Forbidden by a symmetry

$\Phi_1$  and  $\Phi_2$

$Z_2$  or  $T$ -parity:  $\Phi_1 \rightarrow \Phi_1$ ;  $\Phi_2 \rightarrow -\Phi_2$ ;  $T \rightarrow -T$ ;  $L \rightarrow L$

C.S.Chen+CQG,  
PRD82,105004(2010)

**\*Symmetry:** two Higgs doublets ( $\Phi_1$  and  $\Phi_2$ )  
with  $Z_2$ -symmetry or T-parity

CS.Chen,CQG,PRD82,105004(2010)

**T-parity:**  $\Phi_1 \rightarrow \Phi_1$ ;  $\Phi_2 \rightarrow -\Phi_2$ ;  $T \rightarrow -T$ ;  $L \rightarrow L$

~~LLT~~

$$\begin{aligned}
 V = & -\mu_1^2 \phi_1^\dagger \phi_1 + \lambda_1 (\phi_1^\dagger \phi_1)^2 - \mu_2^2 \phi_2^\dagger \phi_2 + \lambda_2 (\phi_2^\dagger \phi_2)^2 \\
 & - \mu_T^2 \text{Tr}(T^\dagger T) + \lambda_T [\text{Tr}(T^\dagger T)]^2 + \lambda'_T \text{Tr}(T^\dagger T T^\dagger T) \\
 & + m^2 \Psi^\dagger \Psi + \lambda_\Psi (\Psi^\dagger \Psi)^2 + \kappa_{\phi_1} \text{Tr}(\phi_1^\dagger \phi_1 T^\dagger T) \\
 & + \kappa'_{\phi_1} \phi_1^\dagger T T^\dagger \phi_1 + \kappa_{\Psi_1} \phi_1^\dagger \phi_1 \Psi^\dagger \Psi \\
 & + \kappa_{\phi_2} \text{Tr}(\phi_2^\dagger \phi_2 T^\dagger T) + \kappa'_{\phi_2} \phi_2^\dagger T T^\dagger \phi_2 \\
 & + \kappa_{\Psi_2} \phi_2^\dagger \phi_2 \Psi^\dagger \Psi + \lambda_3 \phi_1^\dagger \phi_1 \phi_2^\dagger \phi_2 \\
 & + \lambda_4 \phi_1^\dagger \phi_2 \phi_2^\dagger \phi_1 + \rho \text{Tr}(T^\dagger T \Psi^\dagger \Psi) + (M \phi_1^T T^\dagger \phi_2 \\
 & + \lambda_5 \phi_1^\dagger \phi_2 \phi_1^\dagger \phi_2 + \lambda \bar{\phi}_1^\dagger T \bar{\phi}_2^* \Psi + \text{H.c.}),
 \end{aligned}$$

No effect for other couplings

Chen,CQG,Huang,Tsai, PRD87,077702 (2013)

**Without Symmetry:**

$\xi(1,N,2) + \Psi(1,1,4)$



~~LLξ if N>3~~

There are only three possible new renormalizable Yukawa interactions involving SM fermions with non-Higgs scalars:

$$\frac{f}{f_{ab}} \bar{L}_a^c L_b \mathcal{S}, \quad y_{ab} \bar{\ell}_{R_a}^c \ell_{R_b} \Psi, \quad \frac{g_{ab}}{g_{ab}} \bar{L}_a^c L_b T$$

(Colorless scalars)

Zee model

Zee-Babu model

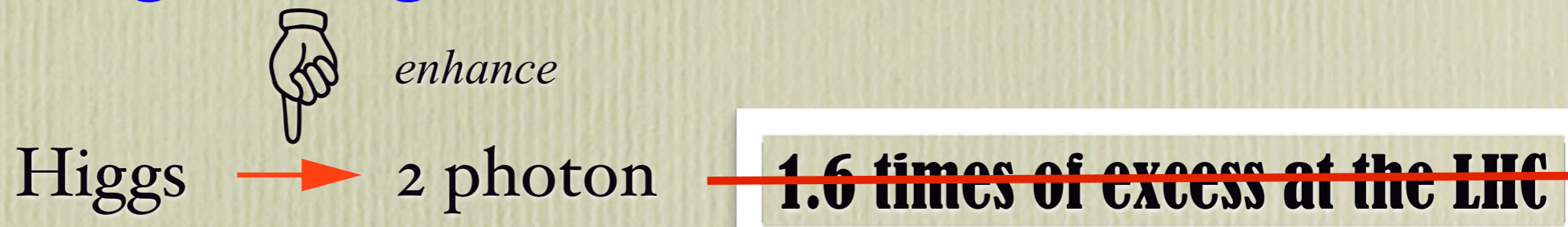
Type-II seesaw

**We will consider higher dimensional multiplets so that NO LL-like term is allowed in the Yukawa interactions.**

We replace  $s=(1,1,2)$  and  $T=(1,3,2)$  by  $\xi=(1,N,2)$

$N>3$  (=4, 5, 6, 7,...) is the quantum # under  $SU(2)_L$   
and  $Y=2$  is the hypercharge with  $Q_{em}=I_3+Y/2$

→ **Multi High Charged Scalars** e.g. for  $N=5$   $\xi = (\xi^{+++}, \xi^{++}, \xi^+, \xi^0, \xi^-)^T$



The scalar potential reads

$$V(\Phi, \xi, \Psi) = -\mu_\Phi^2 |\Phi|^2 + \lambda_\Phi |\Phi|^4 + \mu_\xi^2 |\xi|^2 + \lambda_\xi^\alpha |\xi|_\alpha^4 + \mu_\Psi^2 |\Psi|^2 + \lambda_\Psi |\Psi|^4$$

$$+ \lambda_{\Phi\xi}^\beta (|\Phi|^2 |\xi|^2)_\beta + \lambda_{\Phi\Psi} |\Phi|^2 |\Psi|^2 + \lambda_{\xi\Psi} |\xi|^2 |\Psi|^2$$

$$+ [\mu \xi \xi \Psi + \text{h.c.}]$$

No  $N=4, 6, 8, 10, \dots$ , even dimensions  
due to their antisymmetric products



$N=5, 7, \dots$ , odd dimensions

Some detail calculations:

doublet  $\Phi(1,2,-1)$  + triplet  $T(1,3,2)$  + singlet  $\Psi(1,1,4)$

$$\phi = \begin{pmatrix} \phi^0 \\ \phi^- \end{pmatrix}_{-1}$$

$$T = \begin{pmatrix} T^0 & \frac{T^-}{\sqrt{2}} \\ \frac{T^-}{\sqrt{2}} & T^{--} \end{pmatrix}_{-2}$$

$$\psi_4^{++}$$

The most general potential is

$$\begin{aligned} V(\phi, T, \psi) &= -\mu^2 \phi^\dagger \phi + \lambda_\phi (\phi^\dagger \phi)^2 - \mu_T^2 \text{Tr}(T^\dagger T) + \lambda_T [\text{Tr}(T^\dagger T)]^2 + \lambda'_T \text{Tr}(T^\dagger T T^\dagger T) \\ &+ m^2 \psi^\dagger \psi + \lambda_\psi (\psi^\dagger \psi)^2 + \kappa_{T1} \text{Tr}(\phi^\dagger \phi T^\dagger T) + \kappa_{T2} \phi^\dagger T T^\dagger \phi + \kappa_\psi \phi^\dagger \phi \psi^\dagger \psi \\ &+ \rho \text{Tr}(T^\dagger T \psi^\dagger \psi) + \lambda (\phi^\dagger T \phi \psi + h.c.) \\ &= -\mu^2 (\phi^+ \phi^- + |\phi^0|^2) + \lambda_\phi (|\phi^0|^4 + 2|\phi^0|^2 \phi^+ \phi^- + (\phi^+ \phi^-)^2) \\ &- \mu_T^2 (|T^0|^2 + T^{++} T^{--} + T^+ T^-) \\ &+ (\lambda_T + \lambda'_T) [ |T^0|^4 + 2|T^0|^2 T^+ T^- + (T^{++} T^{--})^2 + 2T^{++} T^{--} T^+ T^- ] \\ &+ (\lambda_T + \frac{\lambda'_T}{2}) (T^+ T^-)^2 + \lambda'_T (T^0 T^{++} T^- T^- + T^{0*} T^{--} T^+ T^+) \\ &+ m^2 (\psi^{++} \psi^{--}) + \lambda_\psi (\psi^{++} \psi^{--})^2 \\ &+ (\kappa_{T1} + \kappa_{T2}) (|\phi^0|^2 |T^0|^2 + \phi^+ \phi^- T^{++} T^{--}) \\ &+ (\kappa_{T1} + \frac{\kappa_{T2}}{2}) (|\phi^0|^2 T^+ T^- + \phi^+ \phi^- T^+ T^-) + \kappa_{T1} (|T^0|^2 \phi^+ \phi^- + |\phi^0|^2 T^{++} T^{--}) \\ &- \frac{\kappa_{T2}}{\sqrt{2}} (\phi^0 T^0 \phi^+ T^- + \phi^0 T^+ \phi^+ T^{--} + h.c.) + \kappa_\psi (|\phi^0|^2 + \phi^+ \phi^-) \psi^{++} \psi^{--} \\ &+ \rho (|T^0|^2 + T^{++} T^{--} + T^+ T^-) \psi^{++} \psi^{--} \\ &+ \lambda (T^{0*} \phi^- \phi^- \psi^{++} + \sqrt{2} \phi^{0*} T^- \phi^- \psi^{++} + \phi^{0*} \phi^{0*} T^{--} \psi^{++} + h.c.) \end{aligned} \quad (4)$$



The vacuum energy

$$-\frac{\mu^2 v^2}{2} + \frac{\lambda_\phi v^4}{4} - \frac{\mu_T^2 v_T^2}{2} + (\lambda_T + \lambda'_T) \frac{v_T^4}{4}$$

Tadpole terms

$$+[-\mu^2 v + \lambda_T v^3 + (\kappa_{T1} + \kappa_{T2}) \frac{v v_T^2}{2}] \phi_1 + [-\mu_T^2 v_T + (\lambda_T + \lambda'_T) v_T^3 + (\kappa_{T1} + \kappa_{T2}) \frac{v^2 v_T}{2}] T_1$$

The mass terms

$$\begin{aligned} & -\mu^2 \left( \frac{\phi_1^2 + \phi_2^2}{2} + \phi^+ \phi^- \right) + \frac{\lambda_T}{2} (3v^2 \phi_1^2 + v^2 \phi_2^2 + 2v^2 \phi^+ \phi^-) - \mu_T^2 \left( \frac{T_1^2 + T_2^2}{2} + T^+ T^- + T^{++} T^{--} \right) \\ & + \frac{\lambda_T + \lambda'_T}{4} (6v_T^2 T_1^2 + 2v_T^2 T_2^2 + 4v_T^2 T^+ T^-) + m^2 (\psi^{++} \psi^{--}) \\ & + \frac{\kappa_{T1} + \kappa_{T2}}{4} [v^2 (T_1^2 + T_2^2) + v_T^2 (\phi_1^2 + \phi_2^2) + 4v v_T \phi_1 T_1] + (\kappa_{T1} + \frac{\kappa_{T2}}{2}) \frac{v^2}{2} T^+ T^- \\ & + \kappa_{T1} \left( \frac{v_T^2}{2} \phi^+ \phi^- + \frac{v^2}{2} T^{++} T^{--} \right) - \frac{\kappa_{T2}}{\sqrt{2}} \frac{v v_T}{2} (\phi^+ T^- + \phi^- T^+) + \frac{\lambda v^2}{2} (T^{--} \psi^{++}) \\ & + \left( \frac{\kappa_\psi v^2}{2} + \frac{\rho v_T^2}{2} \right) \psi^{++} \psi^{--} \end{aligned} \quad (7)$$

Trilinear couplings

$$\begin{aligned} & \frac{\lambda_\phi}{4} (4v \phi_1^3 + 4v \phi_1 \phi_2^2 + 8v \phi_1 \phi^+ \phi^-) + \frac{\lambda_{T1} + \lambda'_T}{4} (4v_T T_1^3 + 4v_T T_1 T_2^2 + 8v_T T_1 T^+ T^-) \\ & + \frac{\lambda'_T}{\sqrt{2}} (v_T T^{++} T^- T^- + v_T T^{--} T^+ T^+) + \frac{\kappa_{T1} + \kappa_{T2}}{2} [v \phi_1 (T_1^2 + T_2^2) + v_T T_1 (\phi_1^2 + \phi_2^2)] \\ & + (\kappa_{T1} + \frac{\kappa_{T2}}{2}) v \phi_1 T^+ T^- + \kappa_{T1} (v_T T_1 \phi^+ \phi^- + v \phi_1 T^{++} T^{--}) \\ & - \frac{\kappa_{T2}}{\sqrt{2}} \left[ \frac{v}{2} (T_1 + iT_2) \phi^+ T^- + \frac{v_T}{2} (\phi_1 + i\phi_2) \phi^+ T^- + \frac{v}{\sqrt{2}} T^+ \phi^+ T^{--} + h.c. \right] \\ & + (\kappa_\psi v \phi_1 + \rho v_T T_1) \psi^{++} \psi^{--} \\ & + \lambda \left[ \frac{v_T}{\sqrt{2}} \phi^- \phi^- \psi^{++} + v T^- \phi^- \psi^{++} + v (\phi_1 - i\phi_2) T^{--} \psi^{++} + h.c. \right] \end{aligned}$$

# The quartic couplings

$$\begin{aligned}
 & \frac{\lambda_\phi}{4} [\phi_1^4 + \phi_2^4 + \phi_1^2 \phi_2^2 + 4\phi_1^2 \phi^+ \phi^- + 4\phi_2^2 \phi^+ \phi^- + 4(\phi^+ \phi^-)^2] \\
 & + (\lambda_T + \lambda'_T) \left[ \frac{1}{4} (T_1^4 + T_2^4 + T_1^2 T_2^2) + (T_1^2 + T_2^2) T^+ T^- + (T^{++} T^{--})^2 + 2T^{++} T^{--} T^+ T^- \right] \\
 & + (\lambda_T + \frac{\lambda'_T}{2}) (T^+ T^-)^2 + \frac{\lambda'_T}{\sqrt{2}} [(T_1 + iT_2) T^{++} T^- T^- + (T_1 - iT_2) T T^{--} T^+ T^+] \\
 & + (\kappa_{T1} + \kappa_{T2}) \left[ \frac{1}{4} (\phi_1^2 + \phi_2^2) (T_1^2 + T_2^2) + \phi^+ \phi^- T^{++} T^{--} \right] \\
 & + (\kappa_{T1} + \frac{\kappa_{T2}}{2}) \left( \frac{\phi_1^2 + \phi_2^2}{2} T^+ T^- + \phi^+ \phi^- T^+ T^- \right) + \frac{\kappa_{T1}}{2} [(T_1^2 + T_2^2) \phi^+ \phi^- + (\phi_1^2 + \phi_2^2) T^{++} T^{--}] \\
 & - \frac{\kappa_{T2}}{\sqrt{2}} \left[ \left( \frac{\phi_1 T_1 - \phi_2 T_2}{2} + i \frac{\phi_2 T_1 + \phi_1 T_2}{2} \right) \phi^+ T^- + \frac{\phi_1 + i\phi_2}{\sqrt{2}} T^+ \phi^+ T^{--} + h.c. \right] \\
 & + [\kappa_\psi \phi^+ \phi^- + \frac{\kappa_\psi}{2} (\phi_1^2 + \phi_2^2) + \frac{\rho}{2} (T_1^2 + T_2^2) + \rho T^{++} T^{--} + \rho T^+ T^-] \psi^{++} \psi^{--} \\
 & + \lambda \left[ \frac{T_1 - iT_2}{\sqrt{2}} \phi^- \phi^- \psi^{++} + (\phi_1 - i\phi_2) T^- \phi^- \psi^{++} + \frac{\phi_1^2 - \phi_2^2 - 2i\phi_1 \phi_2}{2} T^{--} \psi^{++} + h.c. \right] \quad (9)
 \end{aligned}$$

Let

$$\phi^0 = \frac{v + \phi_1 + i\phi_2}{\sqrt{2}} \quad \text{and} \quad T^0 = \frac{v_T + T_1 + iT_2}{\sqrt{2}}$$

Mass mixing matrices

$$\begin{pmatrix} \phi_1 & T_1 \end{pmatrix} \begin{pmatrix} -\frac{\mu^2}{2} + \frac{3}{2} \lambda_\phi v^2 + \frac{(\kappa_{T1} + \kappa_{T2}) v_T^2}{4} & \frac{(\kappa_{T1} + \kappa_{T2}) v v_T}{2} \\ \frac{(\kappa_{T1} + \kappa_{T2}) v v_T}{2} & -\mu_T^2 + \frac{3}{2} (\lambda_T + \lambda'_T) v_T^2 + \frac{(\kappa_{T1} + \kappa_{T2}) v^2}{4} \end{pmatrix} \begin{pmatrix} \phi_1 \\ T_1 \end{pmatrix}$$

$$\begin{pmatrix} \phi_2 & T_2 \end{pmatrix} \begin{pmatrix} -\frac{\mu^2}{2} + \frac{\lambda_\phi}{2} v^2 + \frac{(\kappa_{T1} + \kappa_{T2}) v_T^2}{4} & 0 \\ 0 & -\frac{\mu_T^2}{2} + \frac{(\lambda_T + \lambda'_T)}{2} v_T^2 + \frac{(\kappa_{T1} + \kappa_{T2}) v^2}{4} \end{pmatrix} \begin{pmatrix} \phi_2 \\ T_2 \end{pmatrix}$$

$$\begin{pmatrix} \phi^+ & T^+ \end{pmatrix} \begin{pmatrix} -\mu^2 + \lambda_\phi v^2 + \frac{\kappa_{T1}}{2} v_T^2 & -\frac{\kappa_{T2}}{2\sqrt{2}} v v_T \\ -\frac{\kappa_{T2}}{2\sqrt{2}} v v_T & -\mu_T^2 + (\lambda_T + \lambda'_T) v_T^2 + \left( \frac{\kappa_{T1}}{2} + \frac{\kappa_{T2}}{4} \right) v^2 \end{pmatrix} \begin{pmatrix} \phi^- \\ T^- \end{pmatrix}$$

$$\begin{pmatrix} T^{++} & \psi^{++} \end{pmatrix} \begin{pmatrix} -\mu_T^2 + \lambda_T v_T^2 v_T^2 + \frac{\kappa_{T1}}{2} v^2 & \frac{\lambda}{2} v^2 \\ \frac{\lambda}{2} v^2 & m^2 + \frac{\kappa_\psi}{2} v^2 + \frac{\rho}{2} v_T^2 \end{pmatrix} \begin{pmatrix} T^{--} \\ \psi^{--} \end{pmatrix}$$

# The covariant derivative

$$D_\mu T = \partial_\mu T - i\frac{g}{2}([\vec{\sigma} \cdot \vec{W}_\mu T] + [\vec{\sigma} \cdot \vec{W}_\mu]^t) - i\frac{g'}{2}Y B_\mu T$$

$$\begin{aligned} Tr[(D_\mu T)^\dagger(D^\mu T)] &= (\partial_\mu T^{0*})(\partial^\mu T^0) + (\partial_\mu T^-)(\partial^\mu T^+) + (\partial_\mu T^{--})(\partial^\mu T^{++}) \\ &+ (\partial^\mu T^0)[ig(T^{0*}W_\mu^3 + T^+W_\mu^-) - ig'B_\mu T^{0*}] + h.c. \\ &+ (\partial^\mu T^+)[-ig(T^{--}W_\mu^+ + T^0W_\mu^-) + ig'B_\mu T^-] + h.c. \\ &+ (\partial^\mu T^{++})[-ig(W_\mu^-T^- - T^{--}W_\mu^3) + ig'B_\mu T^{--}] + h.c. \\ &+ [g(T^{0*}W_\mu^3 + T^+W_\mu^-) - g'B_\mu T^{0*}]^2 \\ &+ [g(W_\mu^+T^{--} + T^0W_\mu^-) - ig'B_\mu T^-]^2 \\ &+ [g(W_\mu^-T^- - T^{--}W_\mu^3) - g'B_\mu T^{--}]^2 \end{aligned}$$

$$\begin{aligned} (D_\mu \phi)^\dagger(D^\mu \phi) &= (\partial_\mu \phi^0)(\partial^\mu \phi^{0*}) + (\partial_\mu \phi^+)(\partial^\mu \phi^-) \\ &+ \frac{1}{2}(\partial_\mu \phi^+)[ig(W_\mu^3\phi^- - \sqrt{2}W_\mu^-\phi^0) + ig'B_\mu\phi^-] + h.c. \\ &+ \frac{1}{2}(\partial_\mu \phi^0)[ig(\sqrt{2}W^{-\mu}\phi^+ + W^{3\mu}\phi^{0*}) - ig'B^\mu\phi^{0*}] + h.c. \\ &+ \frac{1}{4}[g(\sqrt{2}W_\mu^-\phi^0 - W_\mu^3\phi^-) - g'B^\mu\phi^-]^2 \\ &+ \frac{1}{4}[g(\sqrt{2}W_\mu^-\phi^+ - W_\mu^3\phi^{0*}) - g'B_\mu\phi^{0*}]^2 \end{aligned}$$

$$\begin{aligned} (D_\mu \psi)^\dagger(D^\mu \psi) &= (\partial_\mu \psi^{--})(\partial^\mu \psi^{++}) + 2ig'B_\mu\psi^{--}(\partial^\mu \psi^{++}) + h.c. \\ &+ 4g'^2 B_\mu B^\mu \psi^{--}\psi^{++} \end{aligned}$$

The masses of gauge bosons are

$$M_W^2 W_\mu^+ W^{-\mu} = \frac{g^2}{4} (v^2 + 2v_T^2) W_\mu^+ W^{-\mu}$$

$$M_W^2 = \frac{g^2}{4} (v^2 + 2v_T^2)$$

The mixing matrix

$$\frac{1}{2} \begin{pmatrix} W_\mu^3 & B_\mu \end{pmatrix} \begin{pmatrix} \frac{g^2}{8} (v^2 + 4v_T^2) & -\frac{gg'}{8} (v^2 - 4v_T^2) \\ -\frac{gg'}{8} (v^2 - 4v_T^2) & \frac{g'^2}{8} (v^2 + 4v_T^2) \end{pmatrix} \begin{pmatrix} W^{3\mu} \\ B^\mu \end{pmatrix}$$

$$\begin{pmatrix} Z_\mu \\ A_\mu \end{pmatrix} = \begin{pmatrix} \cos \theta_W & -\sin \theta_W \\ \sin \theta_W & \cos \theta_W \end{pmatrix} \begin{pmatrix} W_\mu^3 \\ B_\mu \end{pmatrix}$$

$$\begin{pmatrix} Z_\mu \\ A_\mu \end{pmatrix} = \begin{pmatrix} \cos \theta_W & -\sin \theta_W \\ \sin \theta_W & \cos \theta_W \end{pmatrix} \begin{pmatrix} W_\mu^3 \\ B_\mu \end{pmatrix}$$

$$\tan 2\theta_W = \frac{2 \frac{g'}{g} \frac{v^2 - 4v_T^2}{v^2 + 4v_T^2}}{1 - \frac{g'^2}{g^2}}$$

$$M_Z^2 = \frac{g^2}{4 \cos^2 \theta_W} (v^2 + 4v_T^2)$$

$$\rho = \frac{M_W^2}{M_Z^2 \cos^2 \theta_W} = \frac{v^2 + 2v_T^2}{v^2 + 4v_T^2}$$

$$\begin{pmatrix} G^- \\ P^- \end{pmatrix} = \frac{1}{\sqrt{v^2 + 2v_T^2}} \begin{pmatrix} v & \sqrt{2}v_T \\ -\sqrt{2}v_T & v \end{pmatrix} \begin{pmatrix} \phi^- \\ T^- \end{pmatrix}$$

# Gauge-scalar trilinear interactions :

$$W_\mu^+ - T^+ - T^0 : \quad \frac{ig}{\sqrt{2}} W^{-\mu} [(\partial_\mu T_1) T^+ - (\partial_\mu T^+) T_1] + \frac{g}{\sqrt{2}} W^{-\mu} [(\partial_\mu T^+) T_2 - (\partial_\mu T^+) T_2] + h.c.$$

$$W_\mu^+ - T^+ - T^{--} : \quad ig W^{+\mu} [(\partial_\mu T^{--}) T^+ - (\partial_\mu T^+) T^{--}] + h.c.$$

$$Z_\mu (A_\mu) - T^+ - T^- : \quad ig' \sin \theta_W (\partial_\mu T^-) T^+ Z^\mu - ig \sin \theta_W (\partial_\mu T^-) T^+ A^\mu + h.c.$$

$$A_\mu - T^{++} - T^{--} : \quad -i \frac{g}{\cos \theta_W} (\partial_\mu T^{--}) T^{++} A^\mu \quad Z_\mu - T^0 - T^0 : \quad \frac{g}{\cos \theta_W} Z^\mu [(\partial_\mu T_1) T_2 - (\partial_\mu T_2) T_1]$$

$$A_\mu - \phi^+ - \phi^- : \quad -\frac{ig}{2 \cos \theta} (\partial_\mu \phi^-) \phi^+ A^\mu + h.c.$$

$$W_\mu^+ - \phi^- - \phi^0 : \quad \frac{ig}{2} W_\mu^+ [(\partial^\mu \phi^-) \phi_1 - (\partial^\mu \phi_1) \phi^-] + \frac{g}{2} W_\mu^+ [(\partial^\mu \phi^-) \phi_2 - (\partial^\mu \phi_2) \phi^-] + h.c.$$

# Gauge-scalar trilinear interactions :

$$W_\mu^+ - T^+ - T^0 : \quad \frac{ig}{\sqrt{2}} W^{-\mu} [(\partial_\mu T_1) T^+ - (\partial_\mu T^+) T_1] + \frac{g}{\sqrt{2}} W^{-\mu} [(\partial_\mu T^+) T_2 - (\partial_\mu) T^+] + h.c.$$

$$W_\mu^+ - T^+ - T^{--} : \quad ig W^{+\mu} [(\partial_\mu T^{--}) T^+ - (\partial_\mu T^+) T^{--}] + h.c.$$

$$Z_\mu (A_\mu) - T^+ - T^- : \quad ig' \sin \theta_W (\partial_\mu T^-) T^+ Z^\mu - ig \sin \theta_W (\partial_\mu T^-) T^+ A^\mu + h.c.$$

$$A_\mu - T^{++} - T^{--} : \quad -i \frac{g}{\cos \theta_W} (\partial_\mu T^{--}) T^{++} A^\mu \quad Z_\mu - T^0 - T^0 : \quad \frac{g}{\cos \theta_W} Z^\mu [(\partial_\mu T_1) T_2 - (\partial_\mu T_2) T_1]$$

$$A_\mu - \phi^+ - \phi^- : \quad -\frac{ig}{2 \cos \theta} (\partial_\mu \phi^-) \phi^+ A^\mu + h.c.$$

$$W_\mu^+ - \phi^- - \phi^0 : \quad \frac{ig}{2} W_\mu^+ [(\partial^\mu \phi^-) \phi_1 - (\partial^\mu \phi_1) \phi^-] + \frac{g}{2} W_\mu^+ [(\partial^\mu \phi^-) \phi_2 - (\partial^\mu \phi_2) \phi^-] + h.c.$$

$$Z_\mu - \phi^0 - \phi^0 : \quad \frac{g}{2 \cos \theta_W} Z^\mu [(\partial_\mu \phi_1) \phi_2 - (\partial_\mu \phi_2) \phi_1] \quad Z_\mu - Z_\mu - T^0 : \quad \frac{g^2}{\cos^2 \theta_W} v_T Z_\mu Z^\mu T_1$$

$$T^+ - W_\mu^- - Z_\mu : \quad \frac{g^2 + g^2 \sin^2 \theta_W}{\cos \theta_W} \left( \frac{v_T}{\sqrt{2}} \right) Z^\mu W_\mu^- T^+ + h.c. \quad T^+ - W_\mu^- - A_\mu : \quad -g^2 \sin \theta_W \frac{v_T}{\sqrt{2}} T^+ W_\mu^- A^\mu + h.c.$$

$$T^0 - W^{+\mu} - W_\mu^- : \quad g^2 v_T T_1 W^{+\mu} W_\mu^- \quad T^{--} - W^{+\mu} - W_\mu^+ : \quad \frac{g^2}{\sqrt{2}} v_T T^{--} W^{+\mu} W_\mu^+ + h.c.$$

$$\phi^0 - W^{+\mu} - W_\mu^- : \quad \frac{g^2 v}{2} W^{+\mu} W_\mu^- \phi_1 \quad W^{-\mu} - A_\mu - \phi^+ : \quad -\frac{g^2 v}{4 \cos \theta_W} W^{-\mu} A_\mu \phi^+ + h.c.$$

$$\phi^0 - Z^\mu - Z_\mu : \quad \frac{g^2 v}{4 \cos^2 \theta_W} Z^\mu Z_\mu \phi_1 \quad W^{-\mu} - Z_\mu - \phi^+ : \quad \frac{g^2 v}{4 \cos \theta_W} W^{-\mu} Z_\mu \phi^+ + h.c.$$

$$\begin{aligned}
& T^+ - T^- - Z_\mu - Z_\mu \\
& \qquad \qquad \qquad -g'^2 \sin^2 \theta_W T^+ T^- Z^\mu Z_\mu \\
& T^+ - T^- - A^\mu - A_\mu \\
& \qquad \qquad \qquad g'^2 \cos^2 \theta_W T^+ T^- A^\mu A_\mu \\
& T^+ - T^- - Z^\mu - A_\mu \\
& \qquad \qquad \qquad -2g'^2 \sin \theta_W \cos \theta_W T^+ T^- Z^\mu A_\mu \\
& T^{--} - T^+ - W^{+\mu} - Z_\mu \\
& \qquad \qquad \qquad gg' \sin \theta_W W^{+\mu} Z_\mu T^+ T^{--} + h.c.
\end{aligned}$$

$$\begin{aligned}
& T^{--} - T^+ - W^{+\mu} - A_\mu \\
& \qquad \qquad \qquad -(gg' \cos \theta_W + \frac{g^2}{\cos \theta_W}) W^{+\mu} A_\mu T^+ T^{--} + h.c. \\
& T^0 - T^0 - W^{+\mu} - W_\mu^- \\
& \qquad \qquad \qquad \frac{g^2}{2} (T_1^2 + T_2^2) W^{+\mu} W_\mu^- \\
& T^0 - T^+ - W^{-\mu} - Z_\mu \\
& \qquad \qquad \qquad \frac{g^2}{\sqrt{2}} \left( \frac{1 + \sin^2 \theta_W}{\cos \theta_W} \right) T^+ Z_\mu W^{-\mu} (T_1 + iT_2) + h.c. \\
& T^0 - T^+ - W^{-\mu} - A_\mu \\
& \qquad \qquad \qquad -\frac{g^2 \sin \theta_W}{\sqrt{2}} T^+ A_\mu W^{-\mu} (T_1 + iT_2) + h.c.
\end{aligned}$$

$$\begin{aligned}
& T^0 - T^{--} - W^{+\mu} - W_\mu^+ \\
& \qquad \qquad \qquad \frac{g^2}{\sqrt{2}} (T_1 - iT_1) T^{--} W^{+\mu} W_\mu^+ + h.c.
\end{aligned}$$

$$\begin{aligned}
& T^0 - T^0 - Z_\mu - Z^\mu \\
& \qquad \qquad \qquad \frac{g^2}{2 \cos^2 \theta_W} (T_1^2 + T_2^2) Z^\mu Z_\mu \\
& T^+ - T^- - W^{+\mu} - W_\mu^- \\
& \qquad \qquad \qquad 2g^2 W^{+\mu} W_\mu^- T^+ T^- \\
& T^{++} - T^{--} - W^{+\mu} - W_\mu^- \\
& \qquad \qquad \qquad g^2 W^{+\mu} W_\mu^- T^{++} T^{--} \\
& T^{++} - T^{--} - A^\mu - A_\mu \\
& \qquad \qquad \qquad \frac{g^2}{\cos^2 \theta_W} A^\mu A_\mu T^{++} T^{--} \\
& \phi^+ - \phi^- - A^\mu - A_\mu \\
& \qquad \qquad \qquad \frac{g^2}{4 \cos^2 \theta_W} A^\mu A_\mu \phi^+ \phi^-
\end{aligned}$$

$$\begin{aligned}
& \phi^0 - \phi^0 - W^{+\mu} - W_\mu^- \\
& \qquad \qquad \qquad \frac{g^2}{4} W^{+\mu} W_\mu^- (\phi_1^2 + \phi_2^2) \\
& \phi^0 - \phi^+ - W^{-\mu} - A_\mu \\
& \qquad \qquad \qquad -\frac{g^2}{4 \cos \theta_W} W^{-\mu} A_\mu \phi^+ (\phi_1 + i\phi_2) + h.c. \\
& \phi^0 - \phi^0 - Z^\mu - Z_\mu \\
& \qquad \qquad \qquad \frac{g^2}{8 \cos^2 \theta_W} Z^\mu Z_\mu (\phi_1^2 + \phi_2^2) \\
& \phi^+ - \phi^- - W^{-\mu} - W_\mu^+ \\
& \qquad \qquad \qquad \frac{g^2}{2} W^{+\mu} W_\mu^- \phi^+ \phi^- \\
& \phi^0 - \phi^+ - W^{-\mu} - Z_\mu \\
& \qquad \qquad \qquad \frac{g^2}{4 \cos \theta_W} W^{-\mu} Z_\mu \phi^+ (\phi_1 + i\phi_2) + h.c. \\
& Z^\mu (A^\mu) - Z_\mu (A_\mu) - \psi^{++} - \psi^{--} \\
& \qquad \qquad \qquad 4g'^2 \sin^2 \theta_W Z^\mu Z_\mu \psi^{++} \psi^{--} + 4g'^2 \cos^2 \theta_W A^\mu A_\mu \psi^{++} \psi^{--} - 8g'^2 \sin \theta_W \cos \theta_W A^\mu Z_\mu
\end{aligned}$$

## Constraints on the models:

$$\text{VEVs: } \langle \phi^0 \rangle \equiv \frac{v}{\sqrt{2}} \text{ and } \langle T^0 \rangle \equiv \frac{v_T}{\sqrt{2}}.$$

$$M_W^2 = \frac{g^2}{4}(v^2 + 2v_T^2), \quad M_Z^2 = \frac{g^2}{4 \cos^2 \theta_W}(v^2 + 4v_T^2),$$

$$\rho = 1.0002_{-0.0004}^{+0.0007}$$



$$v_T < 4.41 \text{ GeV}$$

Two doubly charged scalars:

$$T = \begin{pmatrix} T^0 & \frac{T^-}{\sqrt{2}} \\ \frac{T^-}{\sqrt{2}} & T^{--} \end{pmatrix} \text{ and } \Psi_{++}$$

Mass eigenstates:

or for N=5  $\xi = (\xi^{+++}, \xi^{++}, \xi^+, \xi^0, \xi^-)^T$

$$\begin{pmatrix} P_1^{\pm\pm} \\ P_2^{\pm\pm} \end{pmatrix} = \begin{pmatrix} \cos \delta & \sin \delta \\ -\sin \delta & \cos \delta \end{pmatrix} \begin{pmatrix} T^{\pm\pm} \\ \Psi^{\pm\pm} \end{pmatrix}$$

$$\sin 2\delta = \left[ 1 + \left( \frac{2m^2 + (2\lambda'_T + \rho)v_T^2}{2\lambda v^2} + \frac{\kappa_2 + \kappa_\Psi}{2\lambda} - \frac{\omega}{\lambda} \right)^2 \right]^{-\frac{1}{2}}$$

$$M_{P_{1,2}}^2 = \frac{1}{2} \left[ a + c \mp \sqrt{4b^2 + (c - a)^2} \right]$$

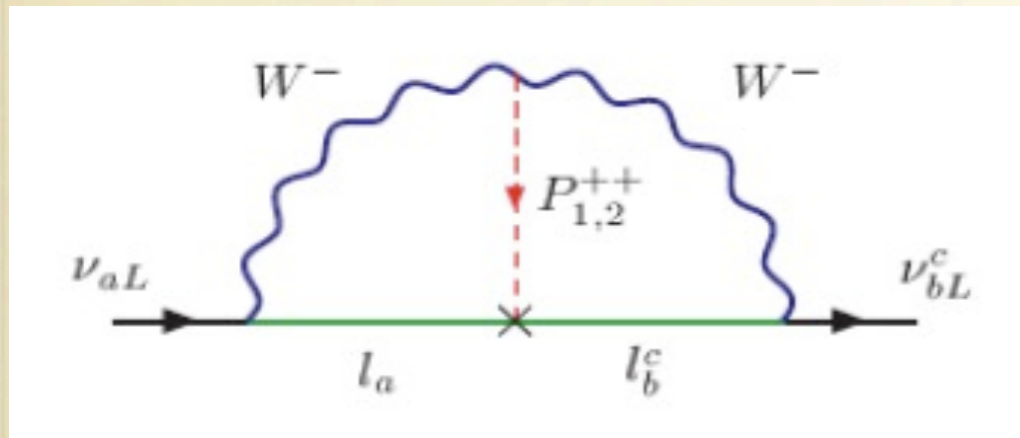
$$\omega \equiv \frac{M}{\sqrt{2}v_T}$$

$$a = \frac{1}{2}(2\omega - \kappa_2)v^2 - \lambda'_T v_T^2, \quad b = \frac{1}{2}\lambda v^2, \quad c = m^2 + \frac{1}{2}(\kappa_\Psi v^2 + \rho v_T^2).$$



● **Neutrino mass generation:**

The neutrino masses are generated radiatively at two-loop level



$$a, b = e, \mu, \tau.$$

$$(m_\nu)_{ab} = \frac{1}{\sqrt{2}} g^4 m_a m_b v_T Y_{ab} \sin(2\delta) [I(M_W^2, M_{P_1}^2, m_a, m_b) - I(M_W^2, M_{P_2}^2, m_a, m_b)]$$

$$I(M_W^2, M_{P_i}^2, m_a, m_b) =$$

$$\int \frac{d^4 q}{(2\pi)^4} \int \frac{d^4 k}{(2\pi)^4} \frac{1}{k^2 - m_a^2} \frac{1}{k^2 - M_W^2} \frac{1}{q^2 - M_W^2} \frac{1}{q^2 - m_b^2} \frac{1}{(k-q)^2 - M_{P_i}^2}$$

$$M_{P_{1,2}} > M_W$$

$$I(M_W^2, M_{P_i}^2, 0, 0) \sim \frac{1}{(4\pi)^4} \frac{1}{M_{P_i}^2} \log^2 \left( \frac{M_W^2}{M_{P_i}^2} \right)$$

$$m_\nu = \tilde{f}(M_{P_1}, M_{P_2}) \times \begin{pmatrix} m_e^2 Y_{ee} & m_e m_\mu Y_{e\mu} & m_e m_\tau Y_{e\tau} \\ m_e m_\mu Y_{e\mu} & m_\mu^2 Y_{\mu\mu} & m_\tau m_\mu Y_{\mu\tau} \\ m_e m_\tau Y_{e\tau} & m_\tau m_\mu Y_{\mu\tau} & m_\tau^2 Y_{\tau\tau} \end{pmatrix}$$

$$= f(M_{P_1}, M_{P_2}) \times \begin{pmatrix} 2.6 \times 10^{-7} Y_{ee} & 5.4 \times 10^{-5} Y_{e\mu} & 9.1 \times 10^{-4} Y_{e\tau} \\ 5.4 \times 10^{-5} Y_{e\mu} & 1.1 \times 10^{-2} Y_{\mu\mu} & 0.19 Y_{\mu\tau} \\ 9.1 \times 10^{-4} Y_{e\tau} & 0.19 Y_{\mu\tau} & 3.17 Y_{\tau\tau} \end{pmatrix}$$

**normal hierarchy:**

$$\begin{pmatrix} \epsilon' & \epsilon & \epsilon \\ \epsilon & 1 + \eta & 1 + \eta \\ \epsilon & 1 + \eta & 1 + \eta \end{pmatrix}$$

$$\tilde{f}(M_{P_1}, M_{P_2}) = \frac{\sqrt{2} g^4 v_T \sin(2\delta)}{128\pi^4} \left[ \frac{1}{M_{P_1}^2} \log^2 \left( \frac{M_W}{M_{P_1}} \right) - \frac{1}{M_{P_2}^2} \log^2 \left( \frac{M_W}{M_{P_2}} \right) \right]$$

$$f = \tilde{f} \times (1\text{GeV}^2)$$

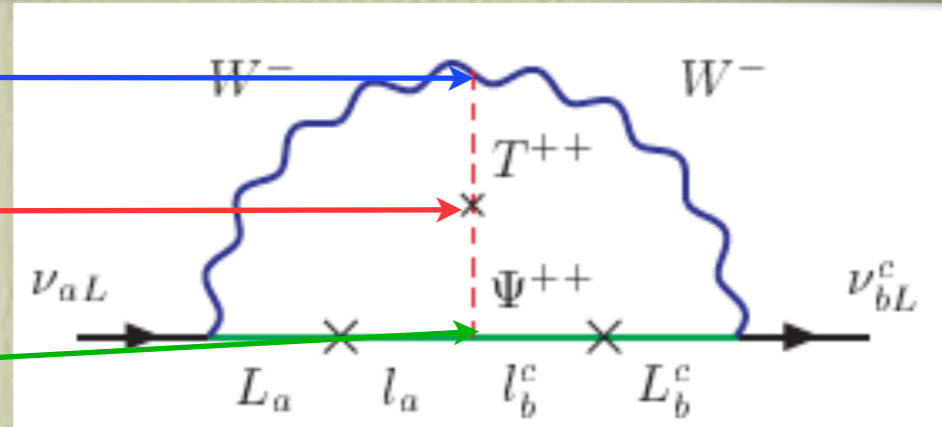
$$\begin{matrix} Y_{ee} < 0.17, & Y_{e\mu} < 0.2, & Y_{e\tau} < 0.2 \\ Y_{\mu\mu} < 3.5, & Y_{\mu\tau} < 0.2, & Y_{\tau\tau} < 0.02 \end{matrix}$$

# The neutrino masses are generated radiatively at two-loop level

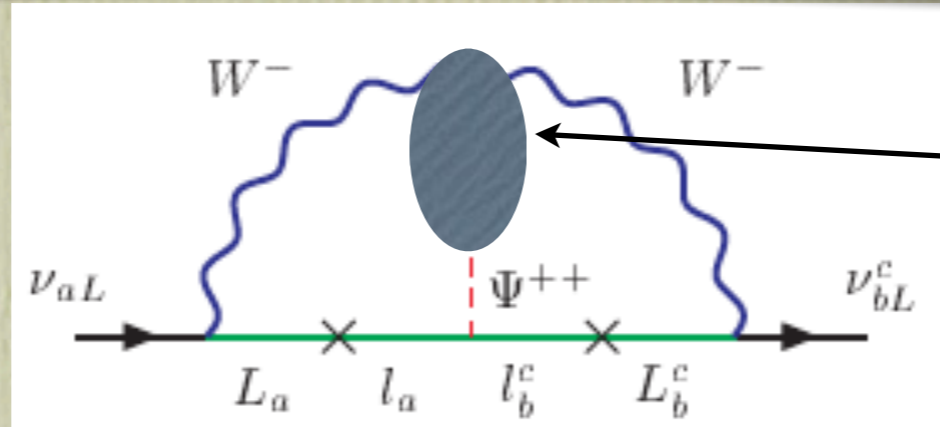
$$Tr[(D_\mu T)^\dagger (D^\mu T)]$$

$$\lambda \Phi^T T \Phi \Psi$$

$$Y_{ab} \bar{l}_{aR}^c l_{bR} \Psi$$

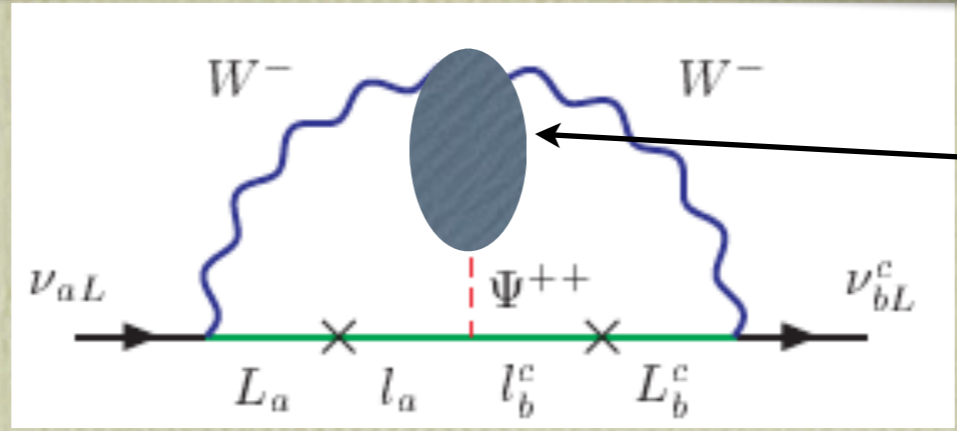


# The neutrino masses are generated radiatively at two-loop level



$$W^\pm W^\pm \Psi^+ \bar{\Psi}^+$$

# The neutrino masses are generated radiatively at two-loop level

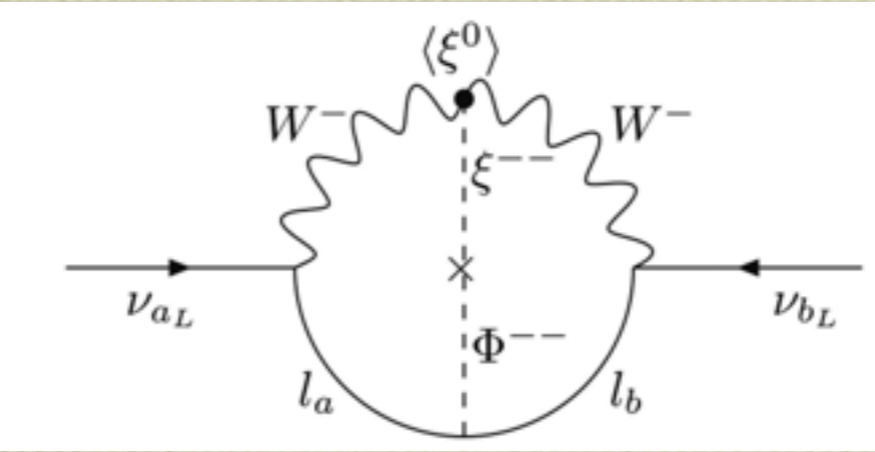


$$W^\pm W^\pm \Psi^\mp \bar{\Psi}$$

**New model:** a multiplet  $\xi (1,5,2)$  + a singlet  $\Phi (1,1,4)$

*Chen, CQG, Huang, Tsai, PRD87, 077702 (2014)*

**Without Symmetry:**

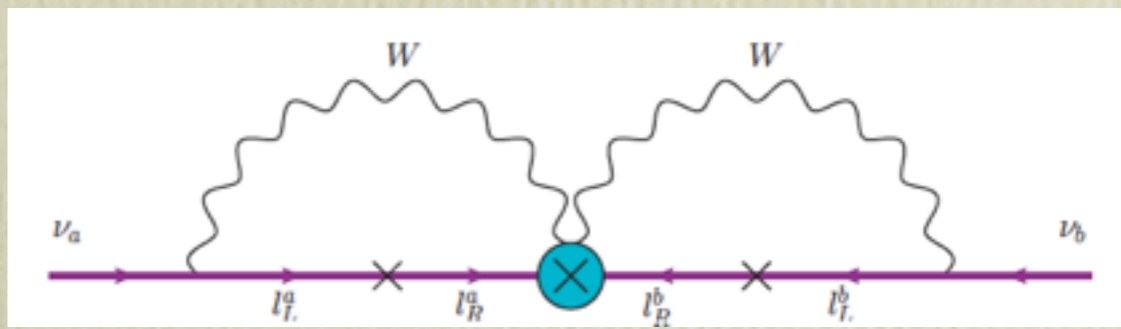


*M. Gustafsson, J.M.No, M.A.Rivera, PRD90, 013012 (2014)*

**dimension-9  
L violating O**

$$\mathcal{O}^9 \equiv C_{ab}^{(9)} \bar{\ell}_{R_a}^c \ell_{R_b} \left[ (D_\mu H)^T i\sigma_2 H \right]^2$$

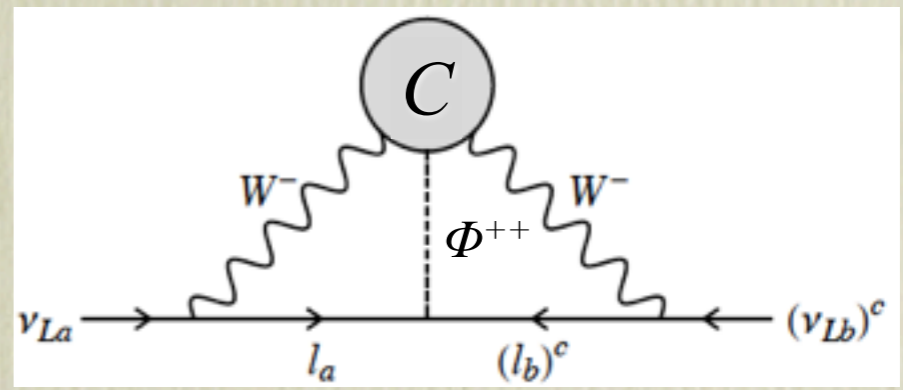
$$\frac{C_{ab}^{(9)}}{\tilde{\Lambda}} \bar{\ell}_{R_a}^c \ell_{R_b} W_\mu^+ W^{+\mu}$$



*S.F.King, A.Merle, L.Panizzi, JHEP1411, 124 (2014)*

**dimension-7 O**

$$\begin{aligned} \mathcal{O}_7^{(a)} &= \Phi (H \otimes H)_{\underline{3}} [(D_\mu H) \otimes (D^\mu H)]_{\underline{3}} \\ \mathcal{O}_7^{(b)} &= \Phi [(D_\mu H) \otimes H]_{\underline{1}} [(D^\mu H) \otimes H]_{\underline{1}} \\ \mathcal{O}_7^{(c)} &= \Phi [(D_\mu H) \otimes H]_{\underline{3}} [(D^\mu H) \otimes H]_{\underline{3}} \end{aligned}$$



$$(M_\nu)_{ab} \propto m_a m_b Y_{ab}$$

For  $Y_{ab} \sim O(1)$

$$(M_\nu)_{ee} \ll (M_\nu)_{e\mu} \ll (M_\nu)_{e\tau} \ll (M_\nu)_{\mu\mu} \ll (M_\nu)_{\mu\tau} \ll (M_\nu)_{\tau\tau}$$

→ **Normal hierarchy**

Z.-z.Xing, *PLB*530,159(2002); *PLB*539,85(2002);  
 Frampton, Glashow, Marfatia, *PLB*536,79(2002);  
 W.L.Guo, Z.-z.Xing, *PRD*67,053002(2003);  
 .....

With  $(M_\nu)_{ee} \approx (M_\nu)_{e\mu} \approx 0$  and the center values of PDG2014 :

$$\sin^2 \theta_{12} = 0.308 \pm 0.017, \quad \sin^2 \theta_{23} = 0.437_{-0.023}^{+0.033}, \quad \sin^2 \theta_{13} = 0.0234_{-0.0019}^{+0.0020},$$

$$\Delta m_{21}^2 = \left(7.54_{-0.22}^{+0.26}\right) \times 10^{-5} \text{ eV}, \quad \Delta m_{32}^2 = (2.43 \pm 0.06) \times 10^{-3} \text{ eV},$$

CQG+L.H.Tsai, *Annals Phys.* 365,210 (2016)  
 CQG, *MPLA*30,1530018(2015)

$$M_\nu \simeq \begin{pmatrix} 0 & 0 & 1.0e^{-i\eta} \\ 0 & 2.4e^{i(\frac{\pi}{2}+\eta)} & 2.3e^{i\frac{\pi}{2}} \\ 1.0e^{-i\eta} & 2.3e^{i\frac{\pi}{2}} & 2.8e^{i(\frac{\pi}{2}+\frac{2\eta}{3})} \end{pmatrix} \times 10^{-2} \text{ eV}$$

Dirac and Majorana phases:

$$\delta = \frac{3}{2}\pi - \frac{3}{2}\eta, \quad \alpha_{21} = \pi + \frac{3}{2}\eta, \quad \alpha_{31} = \frac{3}{2}\pi - \frac{1}{2}\eta, \quad (\eta \simeq 0.07\pi)$$

Agree well with

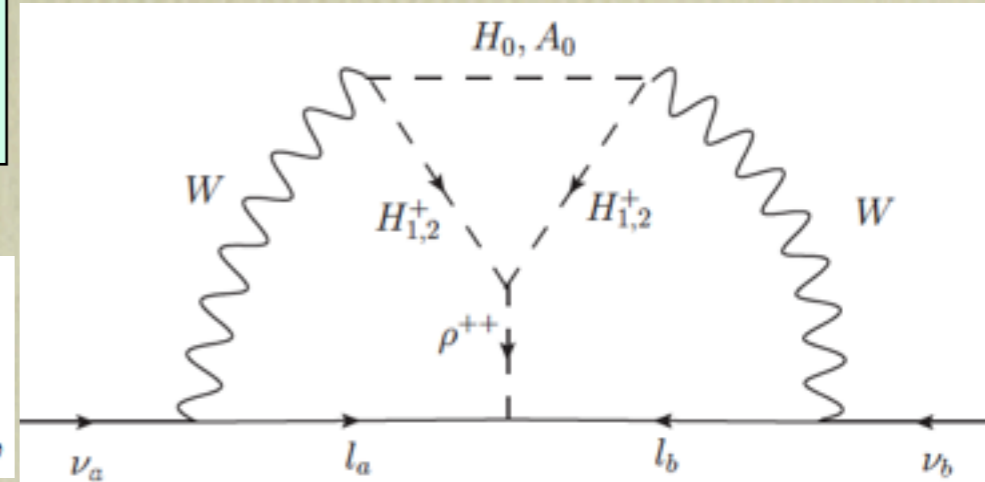
Global Best-Fit

$$\delta/\pi = 1.39_{-0.27}^{+0.38}$$

# At three-loop level

# THE COCKTAIL MODEL FOR NEUTRINO MASSES

*M. Gustafsson, J.M.No, M.A.Rivera, PRL110, 211802 (2013); Erratum, PRL112, 259902 (2014).*



	$SU(2)_L$	$U(1)_Y$	$Z_2$
$\Phi_2$	2	1	-
$S^+$	1	2	-
$\rho^{++}$	1	4	+

$$-\mathcal{L}_{\text{dark}} = \frac{\lambda_5}{2} (\Phi_1^\dagger \Phi_2)^2 + \kappa_1 \Phi_2^T i\sigma_2 \Phi_1 S^- + \kappa_2 \rho^{++} S^- S^- + \xi \Phi_2^T i\sigma_2 \Phi_1 S^+ \rho^{--} + C_{ab} \bar{\ell}_a^c \ell_b R \rho^{++} + \text{h.c.},$$

$$(m_\nu)_{ab} = (x_a C_{ab} x_b) \frac{s_{2\beta}}{(16\pi^2)^3} (\mathcal{A}_1 \mathcal{I}_1 + \mathcal{A}_2 \mathcal{I}_2),$$

*CQG, D.Huang, L.H.Tsai, PRD90, 113005 (2014)*

$$\mathcal{A}_1 = \frac{[\kappa_2 s_{2\beta} + (\xi v) c_{2\beta}]}{m_\rho^2} \frac{(\Delta m_+^2)^2 \Delta m_0^2}{m_\rho^2 v^2},$$

$$\Delta m_+^2 = m_{H_2}^2 - m_{H_1}^2, \quad \Delta m_0^2 = m_{A^0}^2 - m_{H^0}^2$$

$$\mathcal{A}_2 = \frac{\xi v}{m_\rho^2} \frac{\Delta m_+^2 \Delta m_0^2}{v^2},$$

$$x_a = m_a/v, \quad I_1 \sim I_2 \sim O(1)$$

For  $C_{ab} \sim O(1)$

$$m_\nu = \begin{pmatrix} \approx 0 & \approx 0 & 10.1 \\ \approx 0 & -5.01 & 0.0980 \\ 10.1 & 0.0980 & -4.77 \end{pmatrix} \times 10^{-3} + i \begin{pmatrix} \approx 0 & \approx 0 & 0.23 \\ \approx 0 & -2.37 & -2.33 \\ 0.23 & -2.33 & -2.74 \end{pmatrix} \times 10^{-2} \text{ eV}$$

$$C_{ab} = \begin{pmatrix} \leq O(10^{-2}) & \leq O(10^{-2}) & e^{0.224i} \\ \leq O(10^{-2}) & 1.90 \times 10^{-1} e^{-1.78i} & 1.08 \times 10^{-2} e^{-1.56i} \\ e^{0.224i} & 1.08 \times 10^{-2} e^{-1.56i} & 7.73 \times 10^{-4} e^{-1.74i} \end{pmatrix} \times |C_{e\tau}|$$

●  $0\nu\beta\beta$  decays:

$$Y_{ee}$$

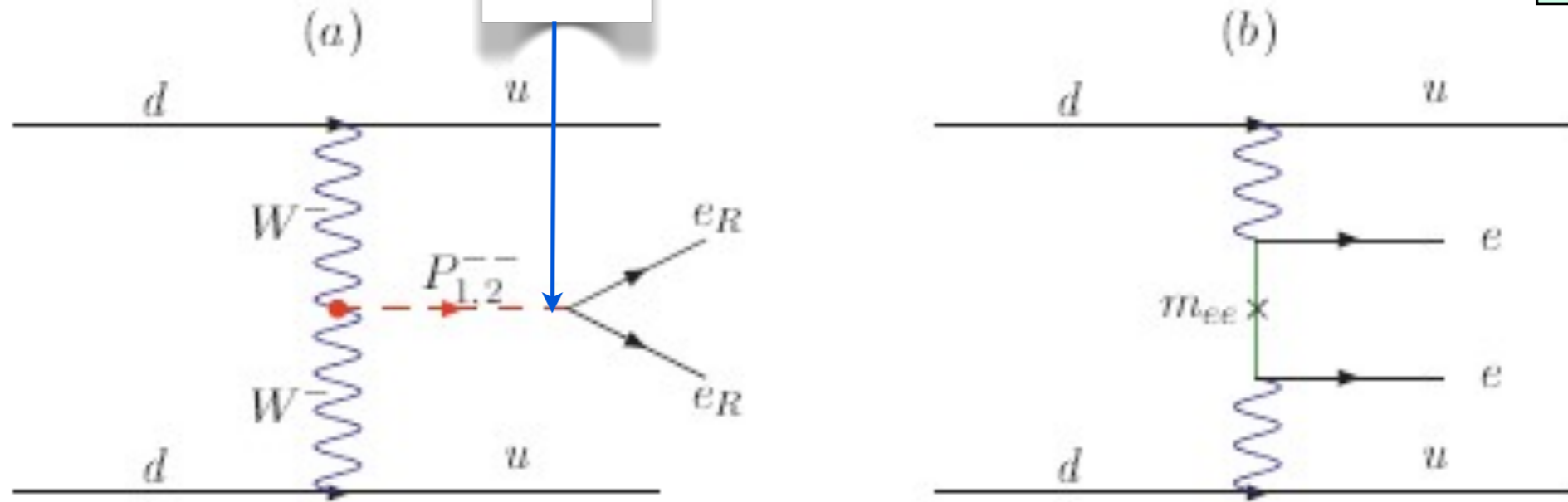


Figure 9:  $0\nu\beta\beta$  decays via exchange of: (a) doubly charged Higgs and (b) light Majorana neutrinos.

$$A_{P_{1,2}^{--}} \sim \frac{g^4 Y_{ee} v_T \sin 2\delta}{16\sqrt{2}M_W^4} \left( \frac{1}{M_{P_1}^2} - \frac{1}{M_{P_2}^2} \right)$$



$$A_\nu \sim \frac{g^4}{M_W^4} \frac{m_{ee}}{\langle p \rangle^2}$$

$$\langle p \rangle \sim 0.1 \text{ GeV}$$

$$A_\nu / A_{P_{1,2}^{--}} \lesssim 10^{-7}$$

Black box theorem is irrelevant as  $0\nu\beta\beta$  dominantly arises from the SD contribution

No other strong constraint on  $Y_{ee}$  except the rate of  $0\nu\beta\beta$  itself. So the rate of  $0\nu\beta\beta$  can be very large, which would correlate with the LHC searches.

●  $0\nu\beta\beta$  decays:

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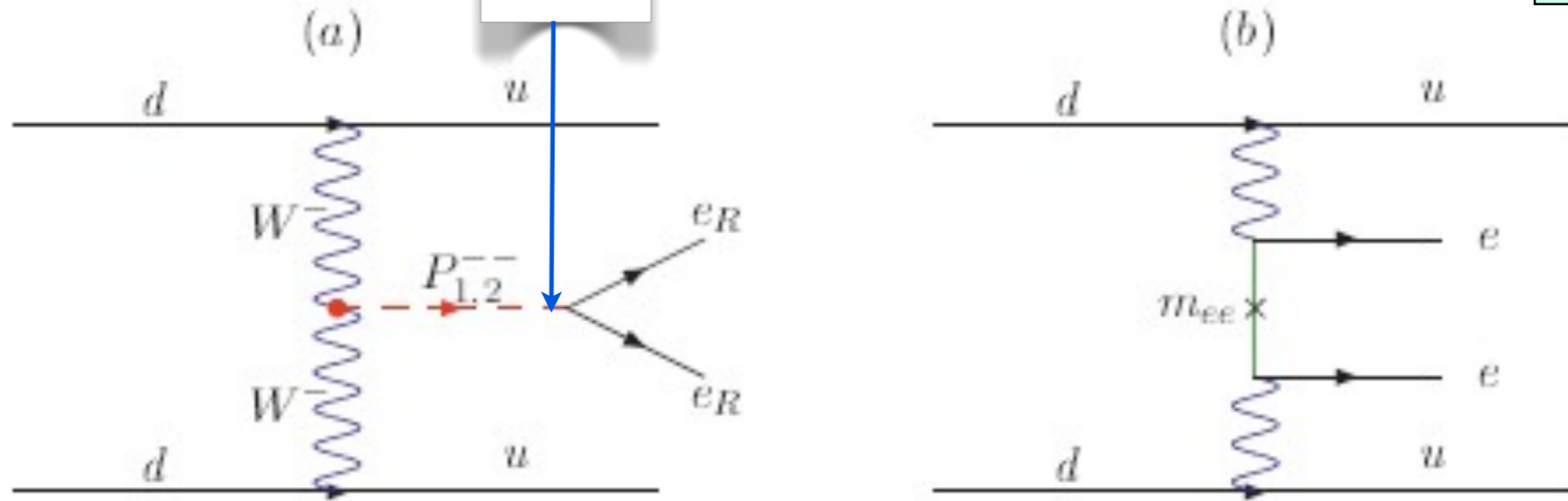


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$$\langle p \rangle \sim 0.1 \text{ GeV}$$

$$A_\nu / A_{P_{1,2}^{--}} \lesssim 10^{-7}$$

The smallness of this ratio is due to the fact that in our model,  $m_{ee}$  is suppressed not only by a two-loop factor, it is also suppressed by the electron mass factor  $(m_e/M_W)^2$  coming from the doubly charged scalar coupling.

**Black box theorem is irrelevant as  $0\nu\beta\beta$  dominantly arises from the SD contribution**

No other strong constraint on  $Y_{ee}$  except the rate of  $0\nu\beta\beta$  itself. So the rate of  $0\nu\beta\beta$  can be very large, which would correlate with the LHC searches.

Currently,  $Y_{ee} < 0(10^{-2})$   
for  $M_P \sim 0(1) \text{ TeV}$



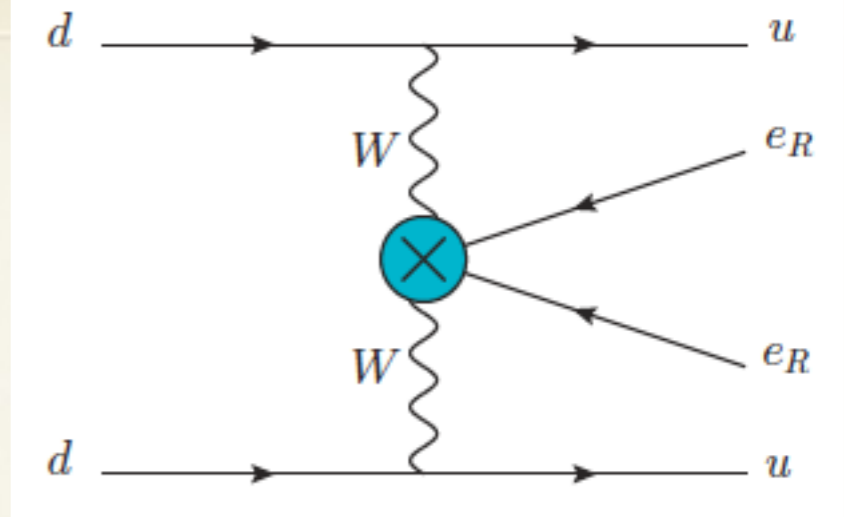
## Dimension-9 L violating O

$$\frac{C_{ab}^{(9)}}{\tilde{\Lambda}} \bar{\ell}^c_{R_a} \ell_{R_b} W_\mu^+ W^{+\mu}$$

M. Gustafsson, J.M.No, M.A.Rivera,  
PRD90, 013012 (2014)

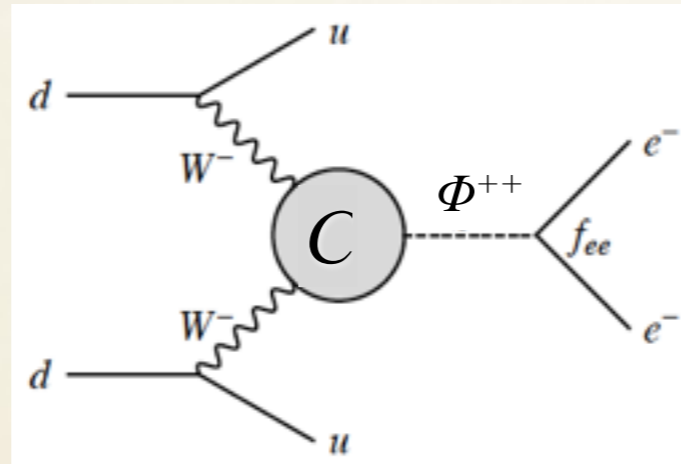
$$\mathcal{L}_{0\nu\beta\beta} = \frac{G_F^2}{2m_p} \epsilon_3 J^\mu J_\mu \bar{e}(1 - \gamma_5)e^c$$

$$J^\mu = \bar{u}\gamma^\mu(1 - \gamma_5)d, \quad \epsilon_3 = -2m_p \mathcal{A}_{0\nu\beta\beta}^{\text{SD}}$$



## Dimension-7 O

S.F.King, A.Merle, L.Panizzi,  
JHEP1411, 124 (2014)



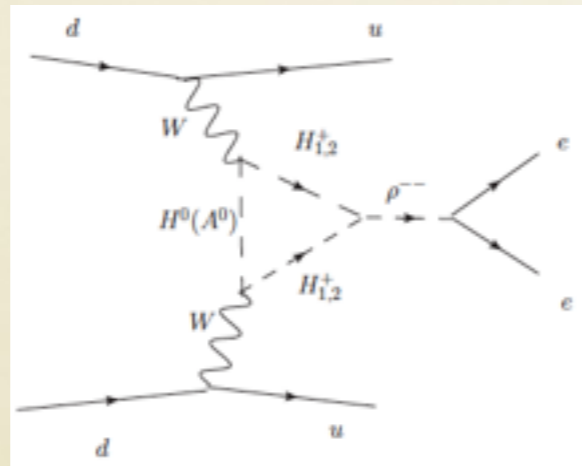
$$\mathcal{L}_{0\nu\beta\beta}^{\text{eff}} = \frac{C_{fee}}{4M_\phi^2 \Lambda^3} J_{L\mu} J_L^\mu \bar{e}(1 - \gamma_5)e^c$$

$$\frac{C_{fee}}{M_\phi^2 \Lambda^3} < 4.0 \times 10^{-3} \text{TeV}^{-5}$$

## Cocktail model

Gustafsson, No, Rivera,  
PRD90, 013012 (2014)

CQG, D.Huang, L.H.Tsai,  
PRD90, 113005 (2014)



	$> T_{\text{exp}} (10^{25} \text{yr})$	$ C_{ee} _{\text{max}}$
GERDA-1( $^{76}\text{Ge}$ ) [22]	2.1	0.0015
KamLAND-Zen( $^{136}\text{Xe}$ ) [23]	1.9	0.0011
NEMO-3( $^{150}\text{Nd}$ ) [24]	0.0018	0.0060
CUORICINO( $^{130}\text{Te}$ ) [25]	0.3	0.0016
NEMO-3( $^{82}\text{Se}$ ) [26, 27]	0.036	0.0059
NEMO-3( $^{100}\text{Mo}$ ) [27]	0.11	0.0021

$$\mathcal{A}_{0\nu\beta\beta}^{\text{loop}} = \frac{\Delta m_+^2 s_{2\theta^+}}{8\pi^2 m_\rho^2} C_{ee} \{ [\Delta m_+^2 s_{2\theta^+} - \xi v (c_{\theta^+}^2 m_{H_2^+}^2 + s_{\theta^+}^2 m_{H_1^+}^2)] [F_{H_1^+, H_2^+, H_0} - F_{H_1^+, H_2^+, A_0}] - \xi v [m_{H_0}^2 F_{H_1^+, H_2^+, H_0} - m_{A_0}^2 F_{H_1^+, H_2^+, A_0}] \}$$

# Multi Charged Scalars

## • Other physics:

### a. Lepton flavor physics:

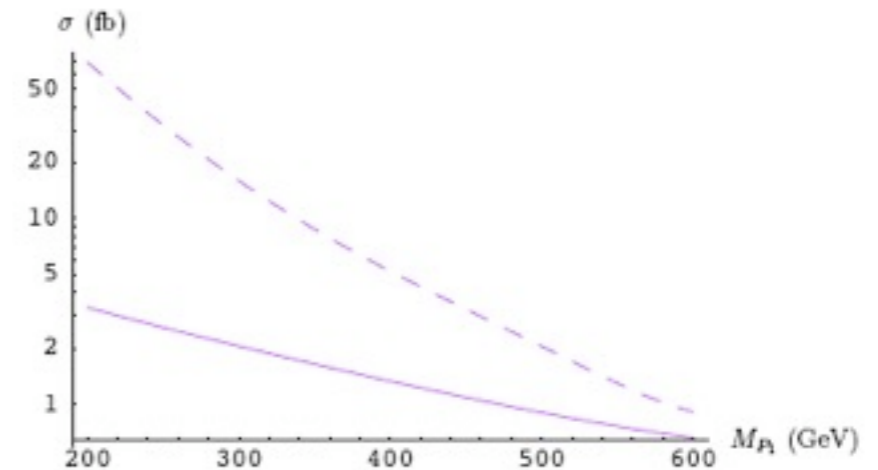
1. Muonium anti-muonium conversion  $\mu^+e^- - \mu^-e^+$   $H_{M\bar{M}} = \frac{Y_{ee}Y_{\mu\mu}}{2M_{--}^2} \bar{\mu}\gamma^\mu e_R \bar{e}\gamma_\mu \mu_R + h.c.$ ,
2. Effective  $e^+e^- \rightarrow l^+l^-$ ,  $l = e, \mu, \tau$ , contact interactions  $\frac{Y_{ee}^2}{M_{--}^2} \bar{e}_R\gamma^\mu e_R \bar{e}_R\gamma_\mu e_R$
3. Rare  $\mu \rightarrow 3e$  decays and its  $\tau$  counterparts
4. Radiative flavor violating charged leptonic decays  $Br(\mu \rightarrow e\gamma) = \frac{\alpha}{3\pi G_F^2} \sum_{l=e,\mu,\tau} \left( \frac{Y_{l\mu}Y_{le}}{M_{--}^2} \right)^2$

### b. Doubly charged scalars at the LHC:

#### 1 Production of the doubly charged Higgs

The WW fusion processes similar to  $0\nu\beta\beta$  decays + the Drell-Yan annihilation processes:

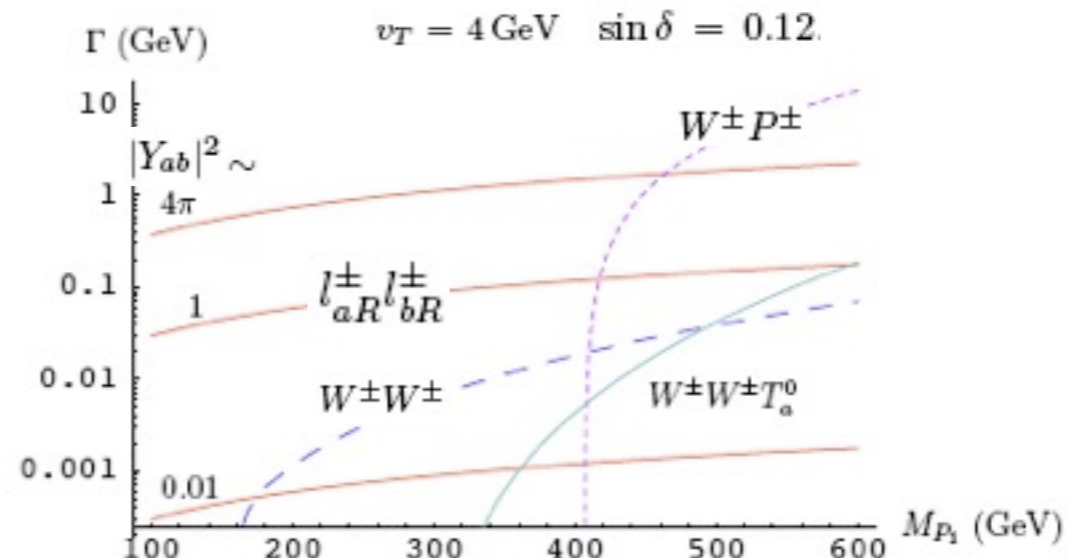
$$q\bar{q} \rightarrow \gamma^*, Z^* \rightarrow P_1^{++}P_1^{--} \quad (q = u, d)$$



#### 2 The decay of $P_1^{\pm\pm}$

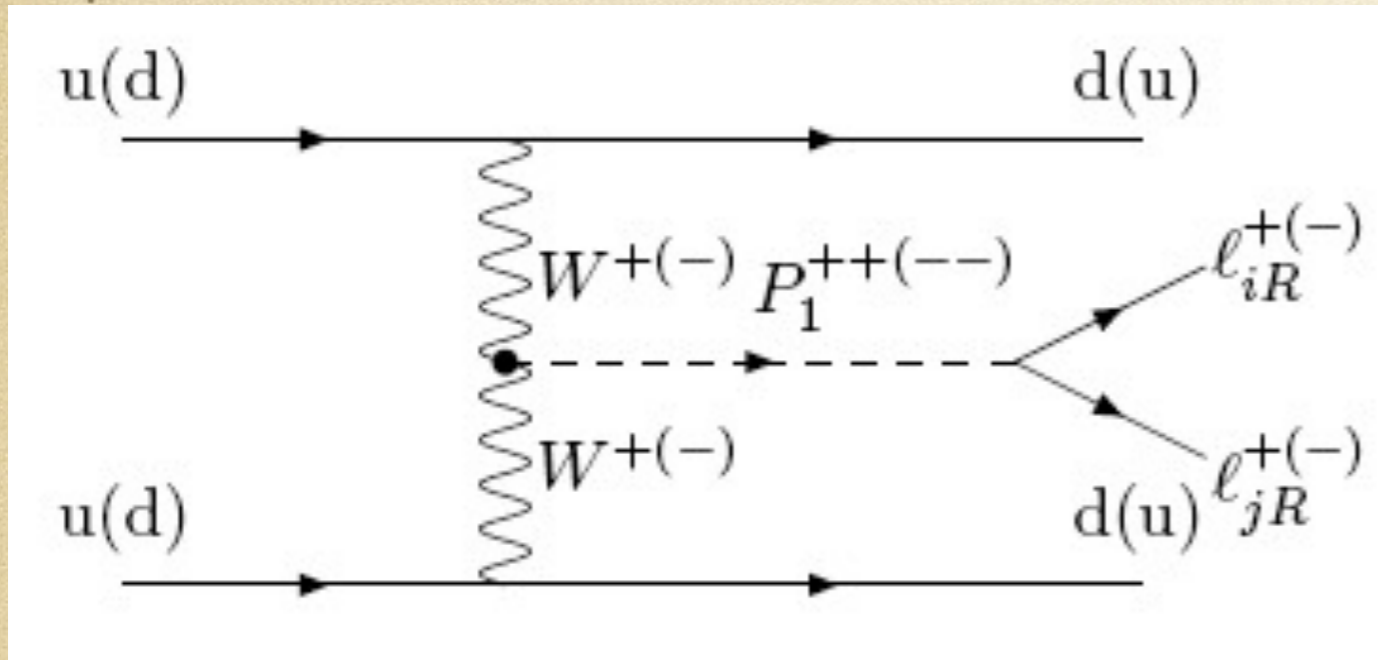
- (1)  $P_1^{\pm\pm} \rightarrow l_{aR}^\pm l_{bR}^\pm$  ( $a, b = e, \mu, \tau$ ),
- (2)  $P_1^{\pm\pm} \rightarrow W^\pm W^\pm$ ,
- (3)  $P_1^{\pm\pm} \rightarrow P^\pm W^\pm$ ,
- (4)  $P_1^{\pm\pm} \rightarrow P^\pm P^\pm$ ,
- (5)  $P_1^{\pm\pm} \rightarrow W^\pm W^\pm X^0$ ,  $X^0 = T_a^0, h^0, P^0$
- (6)  $P_1^{\pm\pm} \rightarrow P^\pm P^\pm X^0$ .

(4) and (6) are not allowed in our model



c. Same-sign single dilepton signatures:

$$pp \rightarrow \ell_i^\pm \ell_j^\pm X \leftarrow JJ$$



Chen, CQG, Zhuridov,  
Eur.Phys.J.C60,119(2009)

$$\frac{d\sigma_{\pm}^{pp}}{d \cos \theta} = A (\lambda_1^{ij})^2 H_{\pm}^{pp}$$

$$A = \frac{G_F^4 M_W^6}{2^7 \pi^5} = 50 \text{ ab}, \quad \lambda_1^{ij} = \sqrt{2 - \delta_{ij}} |Y_{ij}| c_{\delta} s_{\delta},$$

$$H_{\pm}^{pp} = \left( \frac{v_T}{M_W} \right)^2 \int_{z_0}^1 \frac{dz}{z} \int_z^1 \frac{dy}{y} \int_y^1 \frac{dx}{x} p_{\pm}(x, xs) p_{\pm} \left( \frac{y}{x}, \frac{y}{x} s \right) l \left( \frac{z}{y} \right) h \left( \frac{s}{M_{P_1}^2} z \right)$$

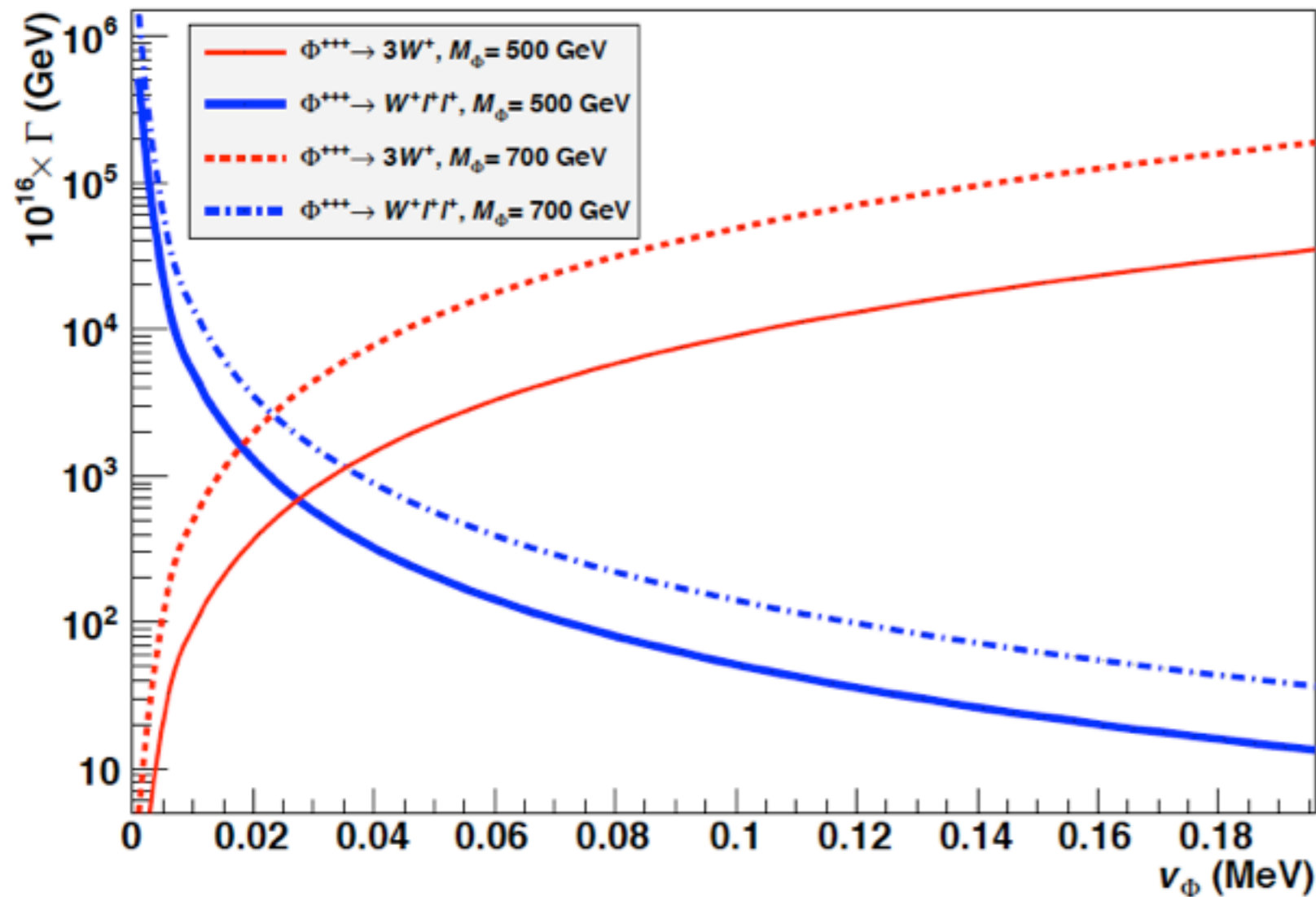
**Remarks:**

- (i) In our model, the final state charged leptons are right-handed. Hence, in principle, helicity measurements can be used to distinguish between our model and those whose doubly charged Higgs coupling only to left-handed leptons (LLT).
- (ii)  $P_1^{\pm\pm}$  will directly produce spectacular lepton # violating signals from like-sign dileptons such as  $e\mu$ ,  $e\tau$  and  $\mu\tau$ .

## d. Triply charged scalar decays:

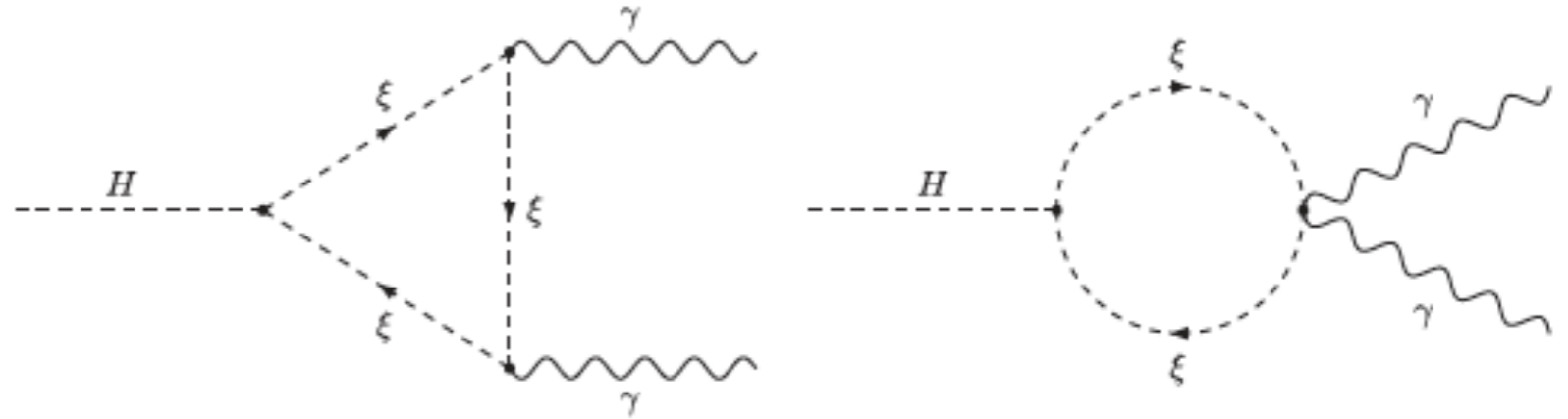
$$\Gamma(\Phi^{+++} \rightarrow 3W) = \frac{3g^6}{2048\pi^3} \frac{v_\Phi^2 M_\Phi^5}{m_W^6}$$

$$\Gamma(\Phi^{+++} \rightarrow W^+ \ell^+ \ell^+) = \frac{g^2}{6144\pi^3} \frac{M_\Phi \sum_i m_i^2}{v_\Phi^2}$$



## e. Multi charged scalar contributions to $H \rightarrow \gamma\gamma$ and $H \rightarrow Z\gamma$ :

$$\Gamma(H \rightarrow \gamma\gamma) = \frac{G_F \alpha^2 m_H^3}{128 \sqrt{2} \pi^3} \left| \sum_f N_f^c Q_f^2 A_{\frac{1}{2}}(\tau_f) + A_1(\tau_W) + \sum_{I_3} (I_3 + 1)^2 \frac{v}{2} \frac{\mu_s}{m_s^2} A_0(\tau_s) \right|^2,$$



$\xi = (1, N, 2)$  with  $N=3, 5, \dots$

e.g.  $\xi = (\xi^{+++}, \xi^{++}, \xi^+, \xi^0, \xi^-)^T$  for  $N=5$

$I_3 = (-N+3)/2$  to  $(N+1)/2$

*Chen, CQG, Huang, Tsai,  
PRD87,077702 (2013)*

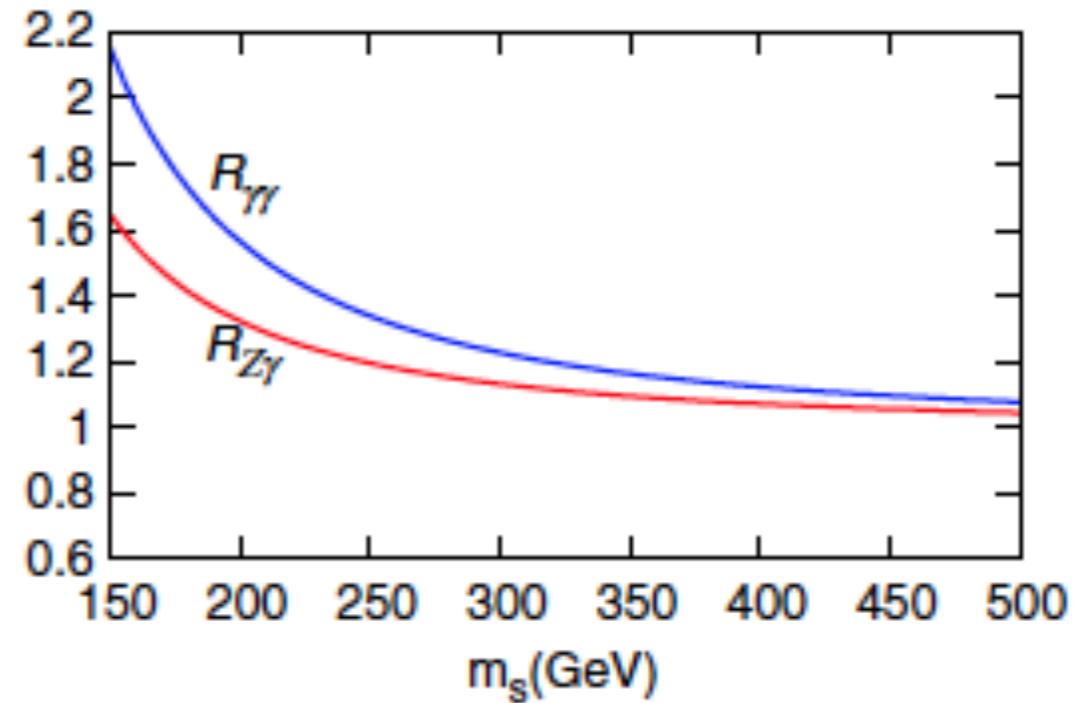


FIG. 4 (color online).  $R_{\gamma\gamma} \equiv \Gamma(H \rightarrow \gamma\gamma)/\Gamma(H \rightarrow \gamma\gamma)_{\text{SM}}$  and  $R_{Z\gamma} \equiv \Gamma(H \rightarrow Z\gamma)/\Gamma(H \rightarrow Z\gamma)_{\text{SM}}$  as functions of the degenerate mass factor  $m_s$  of the multicharged scalar states with  $\mathbf{n} = 5$  and the universal trilinear coupling to Higgs,  $\mu_s = -100$  GeV.

# ☺ Open questions in neutrino physics:

**1. What are the masses of the neutrino mass eigenstates ( $\nu_i$ )?**

**2. Are the neutrino mass eigenstates Dirac or Majorana particles?**

**3. If  $0\nu\beta\beta$  is observed, is  $\nu$  a Majorana particle?**

**4. Can we understand the mixing angles in the neutrino sector? Is there a symmetry behind them?**

**5. Is there CP violation in the neutrino sector?**

**6. Others: sterile neutrino, dark radiation?**

**New field: Astro-Neutrino Physics**

**A real window for new physics**

# The Growing Excitement of Neutrino Physics

Pauli Predicts the Neutrino

Fermi's theory of weak interactions

Reines & Cowan discover (anti)neutrinos

2 distinct flavors identified  
Davis discovers the solar deficit

Kamioka II confirms solar deficit

LEP shows 3 active flavors

SAGE and Gallex see the solar deficit

Kamioka II and IMB see atmospheric neutrino anomaly

Kamioka II and IMB see supernova neutrinos

Kamioka II and IMB see

supernova neutrinos

Nobel prize for discovery of distinct flavors!

LSND sees possible indication of oscillation signal

Nobel Prize for  $\bar{\nu}$  discovery!

Super K sees evidence of atmospheric neutrino oscillations

Super K confirms solar deficit and "images" sun

SNO shows solar oscillation to active flavor

Nobel Prize for neutrino astroparticle physics!

KamLAND confirms solar oscillations

K2K confirms atmospheric oscillations

Daya Bay; Reno see  $\theta_{13}$

Icecube see high energy cosmic neutrinos

2015 Nobel Prize for neutrino oscillation

JUNO reactor experiment

astro-neutrino physics

1930

1955

1980

2005

2015



謝謝！